

A new bound on $|U_{e3}|^2$ from CHOOZ data Nicolo
#. Donato
S. Petcov
S. B

Consider the case of 3 massive ν_i

$$m_1 < m_2 < m_3$$

Probability of $\nu \rightarrow \nu'$, in vacuum

$$P(\nu \rightarrow \nu') = \left| \delta_{\nu\nu'} + \sum_{i=2}^3 U_{\nu'i} (e^{-i\Delta m_{i1}^2 \frac{L}{2E}} - 1) U_{\nu'i}^* \right|^2$$

$\Delta m_{i1}^2 = m_i^2 - m_1^2$

Δm_{21}^2 is relevant for oscillations of solar neutrinos

Δm_{31}^2 is relevant for atm ν 's oscillations

$$\Delta m_{21}^2 \ll \Delta m_{31}^2$$

Two main consequences of this inequality

- ① For atmospheric and LBL values of $\frac{L}{E}$
- $$\Delta m_{21}^2 \frac{L}{2E} \ll 1$$

atm, LBL

$$P(\nu_{\alpha} \rightarrow \nu_{\alpha'}) \approx \delta_{\alpha\alpha'} + U_{\alpha 3} U_{\alpha' 3}^* (e^{-i\Delta m_{31}^2 \frac{L}{2E}} - 1) / 2$$

only $U_{\alpha 3}$ and Δm_{31}^2 determine oscillations

For transitions $\alpha' \neq \alpha$

$$P(\nu_{\alpha} \rightarrow \nu_{\alpha'}) = \frac{1}{2} A_{\alpha\alpha'} (1 - \cos \Delta m_{31}^2 \frac{L}{2E})$$

$$0 \leq A_{\alpha\alpha'} = 4 |U_{\alpha 3}|^2 |U_{\alpha' 3}|^2 \leq 1$$

For survival probability $\alpha' = \alpha$

$$P(\nu_{\alpha} \rightarrow \nu_{\alpha}) = 1 - \frac{1}{2} B_{\alpha\alpha} (1 - \cos \Delta m_{31}^2 \frac{L}{2E})$$

$$0 \leq B_{\alpha\alpha} = \sum_{\alpha' \neq \alpha} A_{\alpha\alpha'} = 4 |U_{\alpha 3}|^2 (1 - |U_{\alpha 3}|^2) \leq 1$$

Notice only $|U_{\alpha 3}|^2$ enters, no phase

$$P(\nu_{\alpha} \rightarrow \nu_{\alpha'}) = P(\bar{\nu}_{\alpha} \rightarrow \bar{\nu}_{\alpha'}) \text{ is}$$

satisfied automatically

CP violation can not be revealed in the case of Δm_{31}^2 dominance

② consequence of $\Delta m_{31}^2 \gg \Delta m_{21}^2$

$|U_{e3}|$ enters incoherently in the survival probability of solar ν_e 's

$$P(\nu_e \rightarrow \nu_e) = |U_{e3}|^4 + (1 - |U_{e3}|^2)^2 P^{(1,2)}(\nu_e \rightarrow \nu_e)$$

$$\cos^2 \theta_{12} = \frac{|U_{e1}|^2}{\sum_{i=1,2} |U_{ei}|^2}, \quad \sin^2 \theta_{12} = \frac{|U_{e2}|^2}{\sum_{i=1,2} |U_{ei}|^2}$$

if $|U_{e3}| \ll 1$

for solar neutrinos

$$P(\nu_e \rightarrow \nu_e) \approx P^{(1,2)}(\nu_e \rightarrow \nu_e)$$

information on $\Delta m_{21}^2, \theta_{12}$

$\nu_e \rightarrow \nu_\mu, \nu_\tau$ transitions

atm, LBL experiments

$\nu_\mu \rightarrow \nu_\tau$

information on $\Delta m_{31}^2, \theta_{23}$

correspond to observation
corrections to this picture?

Direct inf on $|U_{e3}|^2$ from LBL
 CHOOZ experiment
 in the case of Δm_{31}^2 dominance

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - \frac{1}{2} B_{ee} \left(1 - \cos \Delta m_{31}^2 \frac{L}{2E}\right)$$

$$B_{ee} = 4|U_{e3}|^2(1 - |U_{e3}|^2) = \sin^2 2\theta$$

From the CHOOZ exclusion curve

$$B_{ee} \leq B_{ee}^0(\Delta m_{31}^2) \quad \text{Fig}$$

Two bounds on $|U_{e3}|^2$ $\Delta m_{31}^2 \approx 3 \cdot 10^{-2}$
 $B_{ee} \leq 10^{-1}$

$$|U_{e3}|^2 \leq \frac{1}{2} (1 - \sqrt{1 - B_{ee}^0}) \approx \frac{1}{4} B_{ee}^0 \ll 1$$

$$|U_{e3}|^2 \geq \frac{1}{2} (1 + \sqrt{1 - B_{ee}^0}) \approx 1 - \frac{1}{4} B_{ee}^0 \approx 1$$

if $|U_{e3}|^2 \approx 1$, $P(\bar{\nu}_e \rightarrow \bar{\nu}_e) \approx 1$, thus

$|U_{e3}|^2$ is small

$$\underline{|U_{e3}|^2 \leq 2.5 \cdot 10^{-2}}$$

from S-K spectrum measurement
 and θ/N LMA is the most
preferred allowed region

after SNO the preference of LMA
increased

for LMA Δm_{21}^2 can be (a few) 10^{-4} eV^2

in the CHOOZ experiments

$$L \approx 1000 \text{ m}, E \approx 3 \text{ MeV}$$

$$\Delta m_{31}^2 \frac{L}{2E} \approx 0.8 \cdot 10^{-1}$$

corrections to CHOOZ bound on

$$\underline{|\theta_{e3}|^2}?$$

$$P(\bar{\nu}_e \rightarrow \bar{\nu}_e) = 1 - 2|\theta_{e3}|^2 (1 - |\theta_{e3}|^2) (1 - \cos \Delta m_{31}^2 \frac{L}{2E})$$

$$- 2|\theta_{e2}|^2 |\theta_{e1}|^2 (1 - \cos \Delta_{21})$$

$$+ 2|\theta_{e3}|^2 |\theta_{e2}|^2 (\cos(\Delta_{31} - \Delta_{21}) - \cos \Delta_{31})$$

additional terms

at $\Delta m_{21}^2 \leq 4 \cdot 10^{-4} \text{ eV}^2$ no changes

at $\Delta m_{21}^2 = 6 \cdot 10^{-4} \text{ eV}^2$ $|\theta_{e3}| \leq 2 \cdot 10^{-2} \frac{\sin^2 \theta_{12}}{2}$

$\Delta m_{21}^2 = 8 \cdot 10^{-4} \text{ eV}^2$ $|\theta_{e3}| \leq 1 \cdot 10^{-2}$

More exact determination of
LMA allowed region

(S-K, SNO, Kamland, Borexino,...)

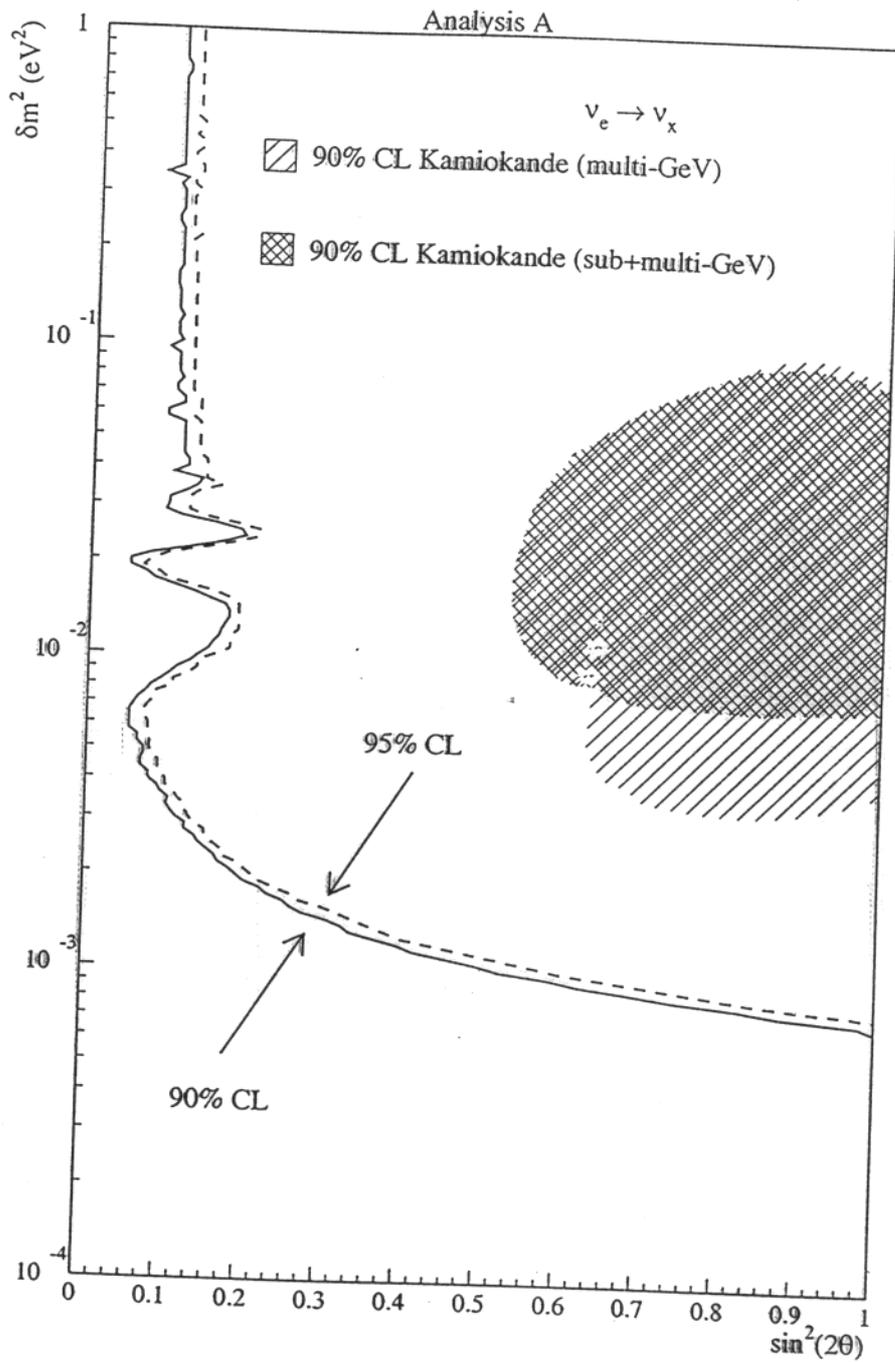


Figure 9: Exclusion plot for the oscillation parameters based on the absolute comparison of measured vs. expected positron yields.

