



# OPERA: ready to run!

Maria Teresa Muciaccia  
Università degli Studi & INFN-Bari  
on behalf of the OPERA Collaboration



The detector is ready for the next phase

...waiting for the first  
neutrino interaction  
in emulsion.....



# Summary

- Aim and strategy of the experiment
- OPERA detector
- CNGS 2007 run
- $\nu_{\mu}$  interactions in a lead-emulsion target
- CNGS 2008 run
- Conclusions



# The Oscillation Project with Emulsion tRacking Apparatus **OPERA**

Long baseline experiment searching for the  $\nu_\tau$  *appearance* in a pure  $\nu_\mu$  beam produced at CERN

CNGS (CERN to Gran Sasso beam)

$\langle E \rangle = 17$  GeV,  $L = 732$  km

Hybrid set-up (nuclear emulsions + electronic detectors)



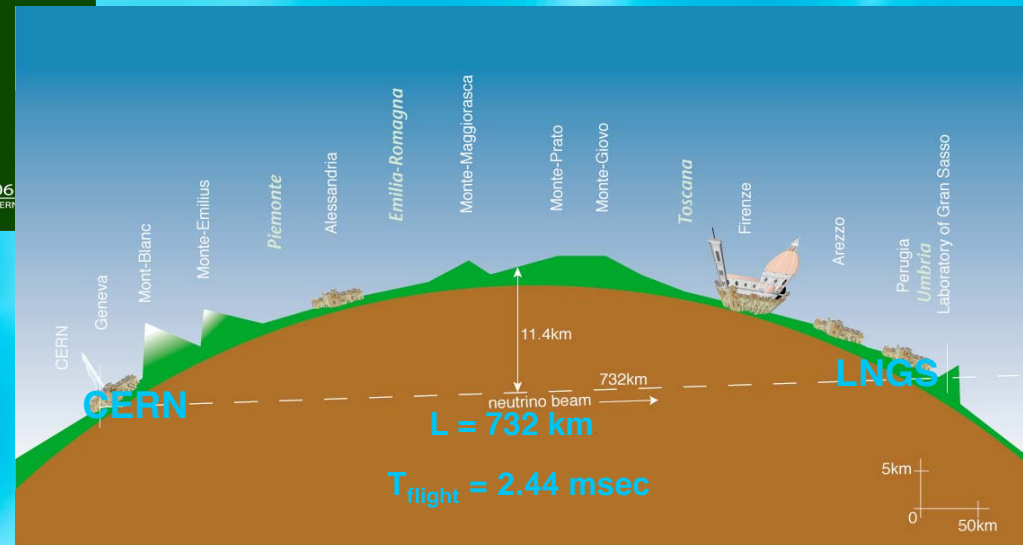
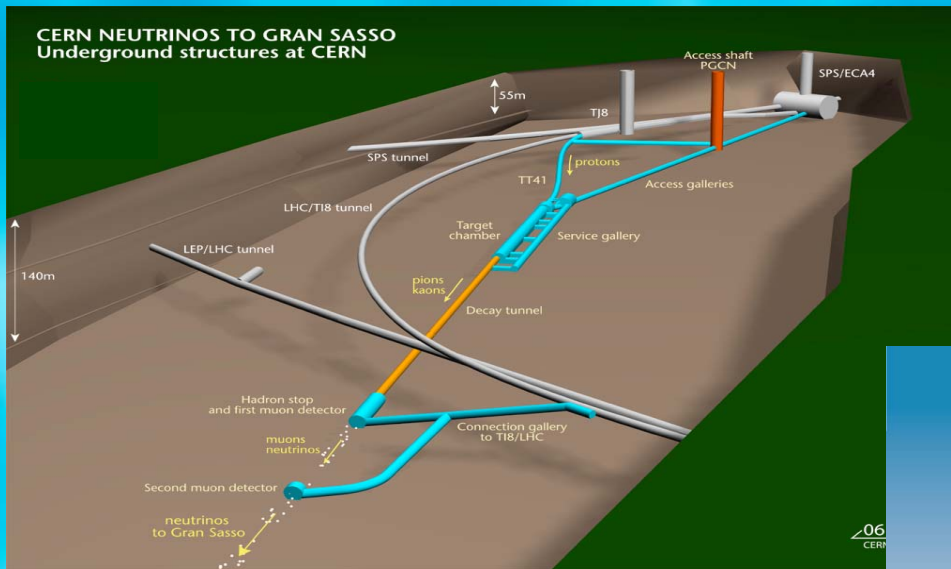
# The CERN Neutrino to Gran Sasso beam

## CNGS

400 GeV/c protons (CERN SPS) on graphite target  
 $\nu$  beam produced in the decay in flight of secondary  $\pi$ , K's  
 in 1km-long decay tunnel towards LNGS

optimized for appearance

$\langle E_{\nu_{\mu}} \rangle$	17 GeV
$(\nu_e + \bar{\nu}_e) / \nu_{\mu}$	0.87%
$\bar{\nu}_{\mu} / \nu_{\mu}$	2.1%
$\nu_{\tau}$ prompt	negligible



$4.5 \times 10^{19}$  p.o.t./year, 200 days/year  
 $\epsilon = 55\%$  shared mode



# Aim of the experiment

All the experiments indicate  $\nu_{\mu} \rightarrow \nu_{\tau}$   
as the dominant oscillation mode

**but still missing:**  
**direct observation of**  
**oscillated  $\nu_{\tau}$**

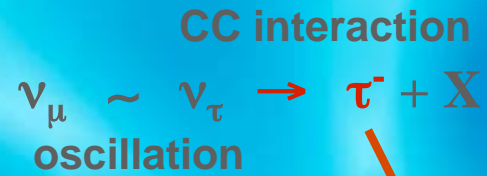
**Detection of  $\nu_{\tau}$  CC interactions and *direct* observation of  $\tau$  decays**



**Provide an unambiguous evidence for  $\nu_{\mu} \rightarrow \nu_{\tau}$  oscillations in the  
parameter region indicated by the atmospheric neutrino data**



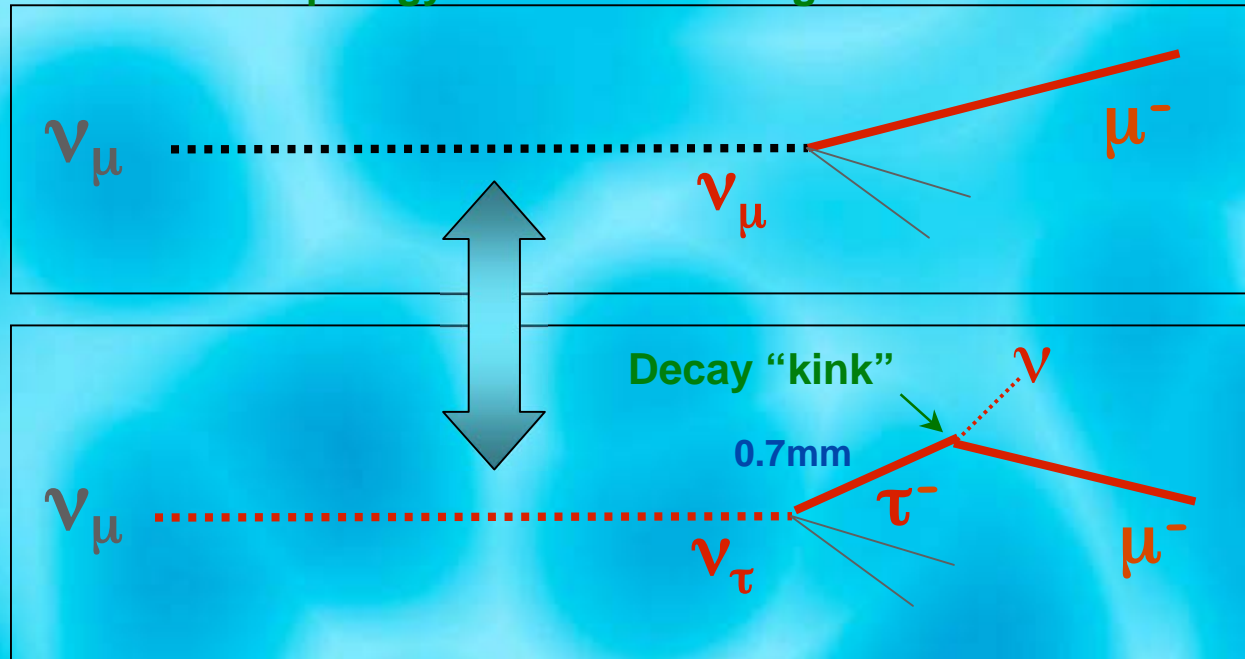
# Experimental signature of the $\nu_\tau$ appearance



Detection of the  $\tau$  decay:  
 the challenge is to distinguish  $\nu_\tau$   
 interactions from  $\nu_\mu$  interactions

- $\mu^- \nu_\tau \bar{\nu}_\mu$  B. R. ~ 17%
- $h^- \nu_\tau n(\pi^0)$  B. R. ~ 50%
- $e^- \nu_\tau \bar{\nu}_e$  B. R. ~ 18%
- $\pi^+ \pi^- \pi^0 \nu_\tau n(\pi^0)$  B. R. ~ 14%

## Topology selection: kink signature





# Detecting $\nu_\tau$

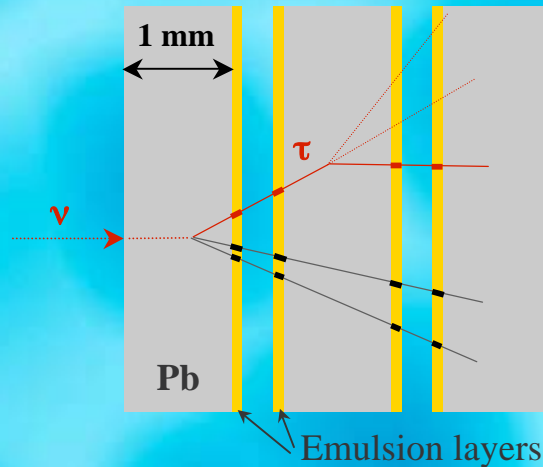
Two conflicting requirements:

large mass (low cross-section)

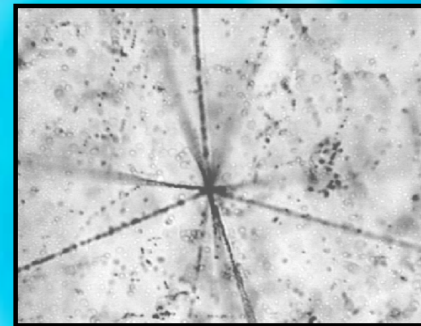
high granularity (signal selection / background rejection)



lead-nuclear emulsion target  
segmented into basic units  
called *bricks*



**Nuclear emulsions** (high granularity  $\sim 300$  hits/mm)



This technology allows a sub-micron spatial resolution, and a 3D reconstruction of particle tracks, necessary for  $\tau$  detection

What the brick cannot do:

- signal a neutrino interaction
- identify muons





# OPERA Detector Design

LNGS Hall C

Installation started in  
**May 2003**

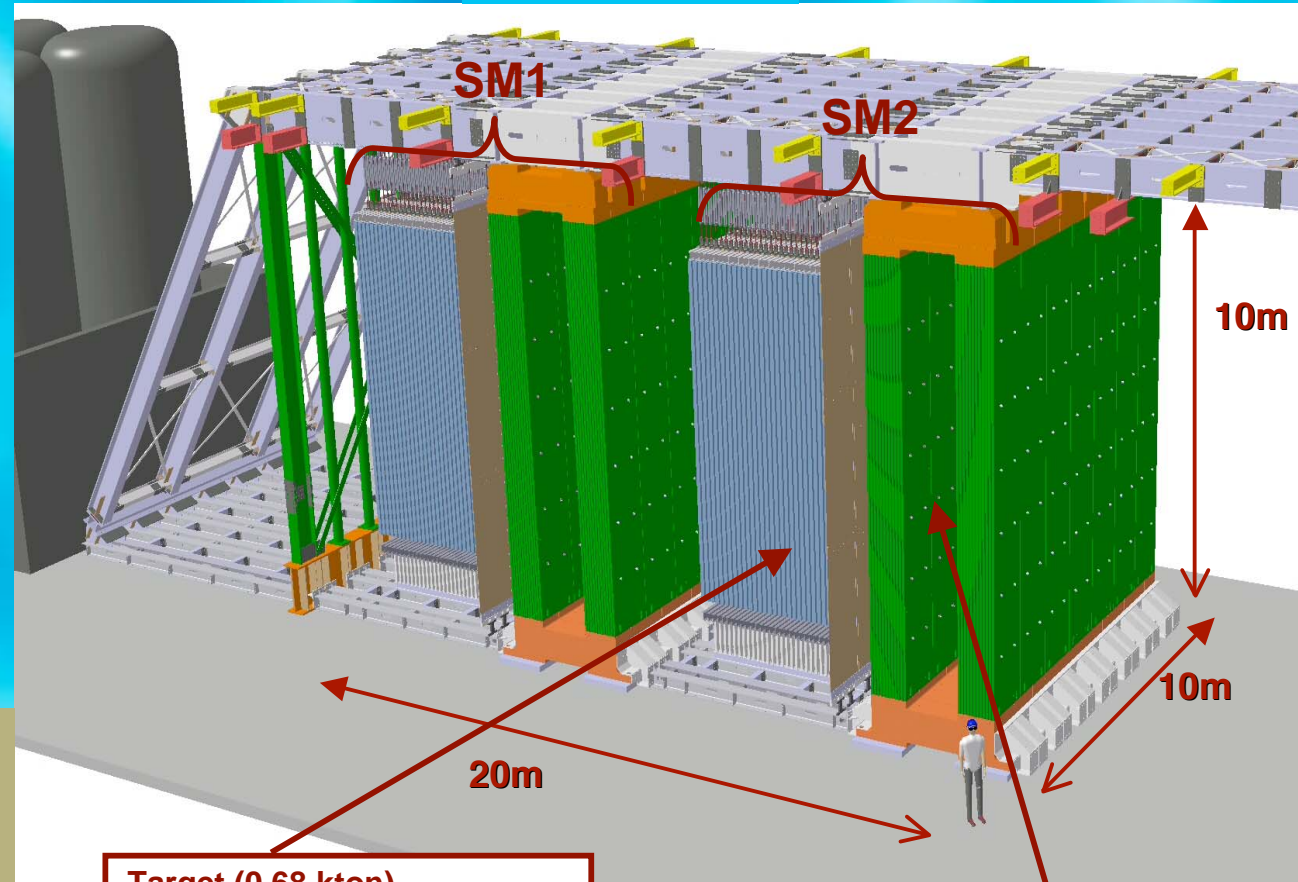
First observation of the  
CNGS neutrinos on  
**August 18<sup>th</sup>, 2006**

First observation of the  
CNGS neutrinos  
in emulsion on  
**October 2<sup>nd</sup>, 2007**

Target: 1350 tons

5 year running @  $4.5 \cdot 10^{19}$

- 22,000 neutrino interactions
- ~110  $\nu_\tau$  interactions
- ~11  $\nu_\tau$  identified
- <1 event of background



### Target (0.68 kton)

29 lead/emulsion brick walls,  
alternated to scintillator planes,  
(T T) to select the brick  
containing neutrino interaction

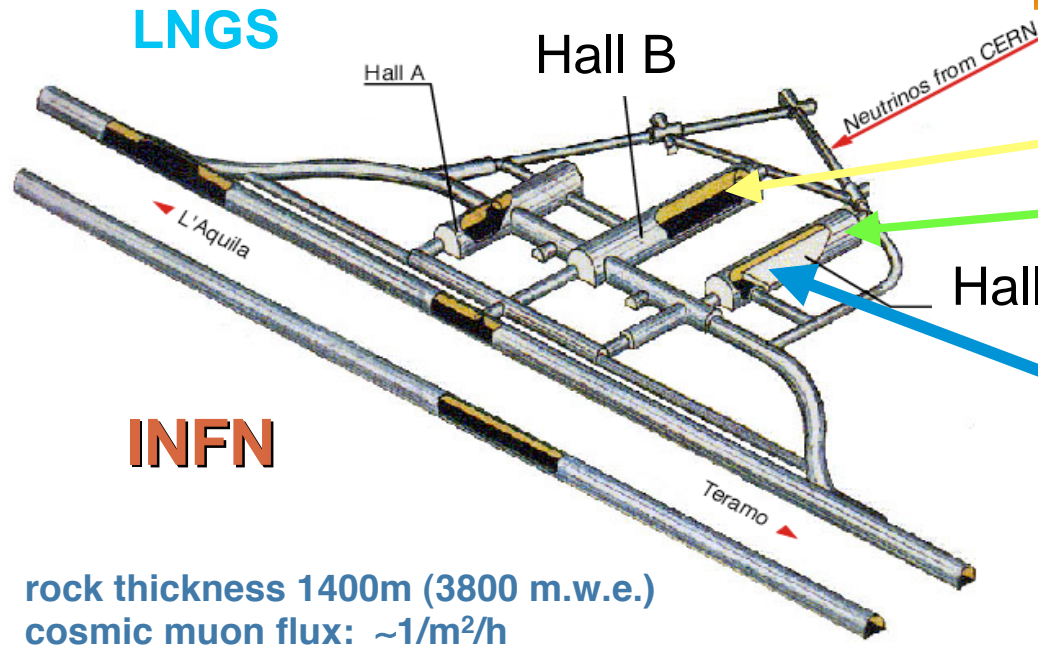
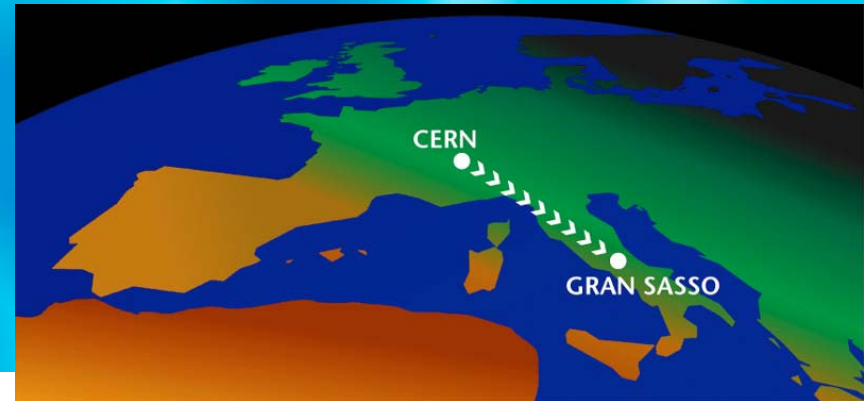
### Muon Spectrometer

Magnet equipped  
with 22 RPC planes:  
ID muons, charge and  
momentum measurement



# OPERA at LNGS

Largest underground laboratory  
for astro-particle physics



ICARUS

BOREXINO

Hall C

OPERA (CNGS1)

underground area: 18 000 m<sup>2</sup>  
external facilities  
easy access

rock thickness 1400m (3800 m.w.e.)  
cosmic muon flux:  $\sim 1/\text{m}^2/\text{h}$



# $\nu_{\mu} \rightarrow \nu_{\tau}$ Oscillation Search

Number of expected signal and background events

$\tau$ decay channel	Signal		Background
	$\Delta m^2 = 2.5 \times 10^{-3} \text{ eV}^2$	$\Delta m^2 = 3.0 \times 10^{-3} \text{ eV}^2$	
$\tau \rightarrow \mu$	2.9	4.2	0.17
$\tau \rightarrow e$	3.5	5.0	0.17
$\tau \rightarrow h$	3.1	4.4	0.24
$\tau \rightarrow 3h$	0.9	1.3	0.17
<b>ALL</b>	<b>10.4</b>	<b>14.9</b>	<b>0.75</b>

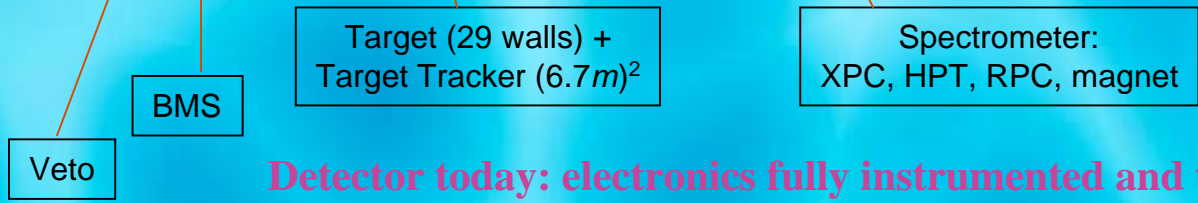
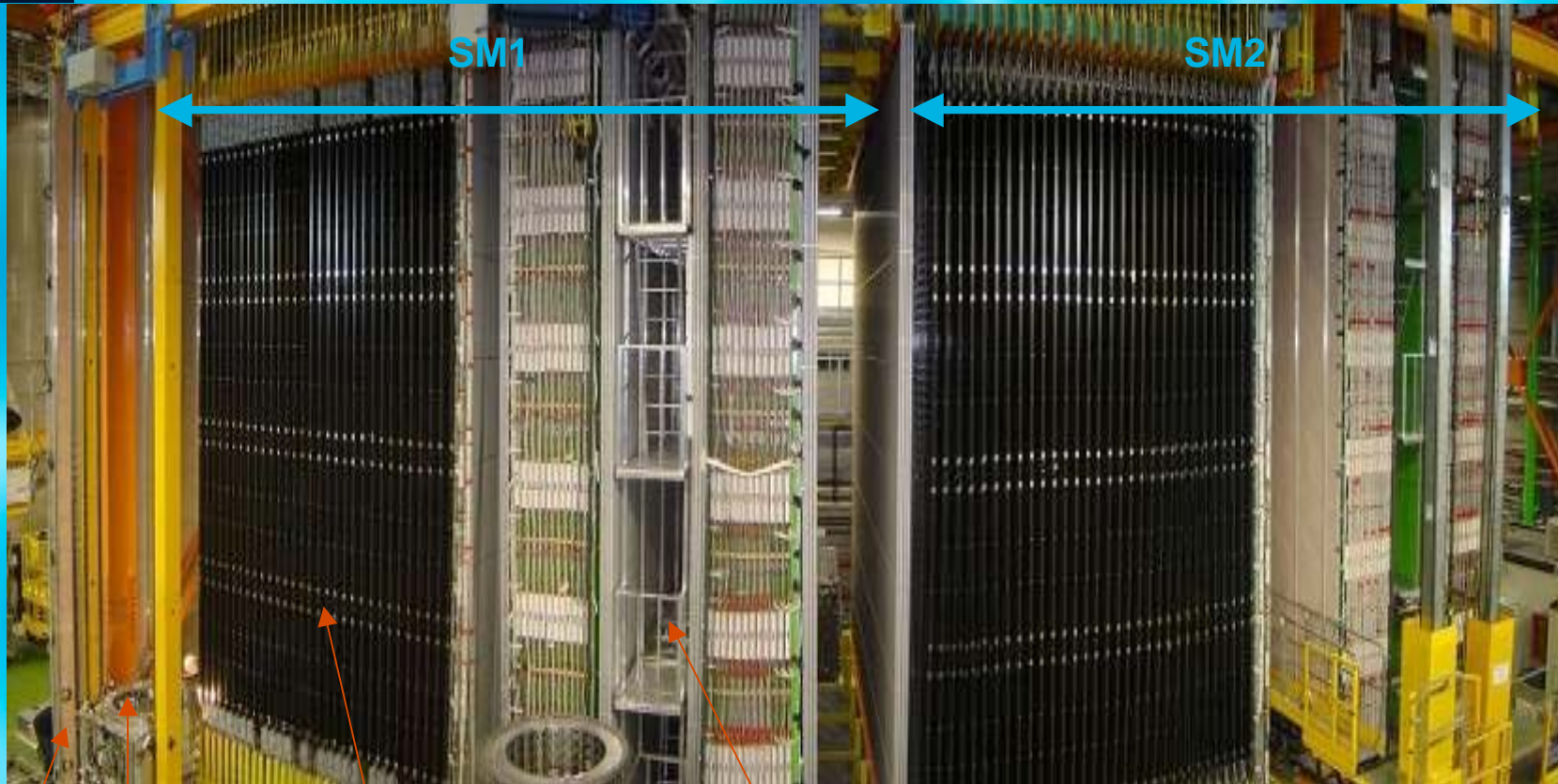
Main background sources:

- charm production and decays
- hadron re-interactions in lead
- large-angle muon scattering in lead

full mixing, 5 year run @  $4.5 \times 10^{19}$  pot/year  
 25% target mass reduction = 1.35 kton  
 ~150000 bricks



# OPERA Detector Today



**Detector today: electronics fully instrumented and tested**  
**Brick filling is in progress, > 110000 bricks have already been inserted**  
**75% target: SM1 full, SM2 at 50%**



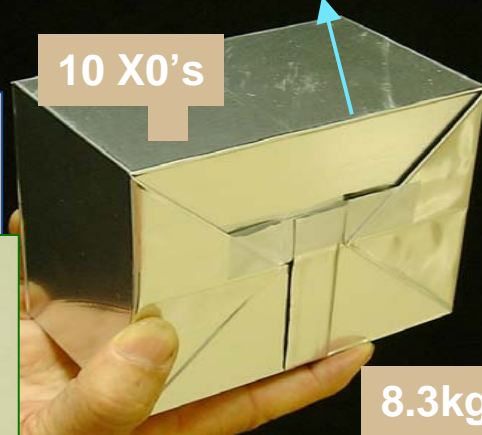
# Brick Target

- Micro-metric space resolution (emulsion) + target mass (lead)
- Compact and modular structure

## **Brick (basic unit)**

56 Pb (1mm) plates + 57 emulsion films

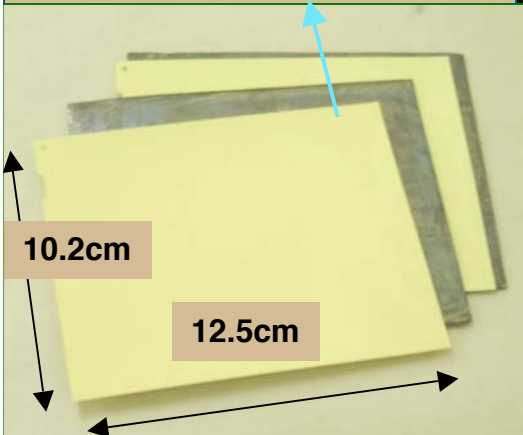
10 X0's



8.3kg

## **Emulsion film**

2 emulsion layers (44  $\mu\text{m}$  thick)  
poured onto a 200  $\mu\text{m}$  plastic

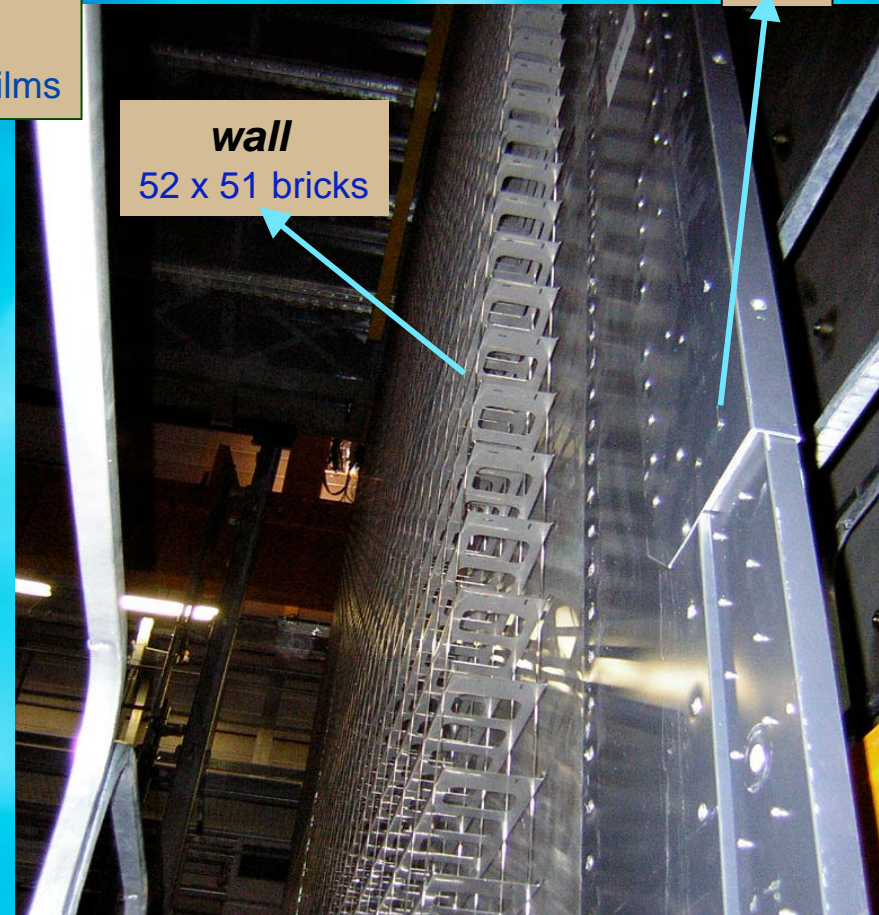


10.2cm

12.5cm

## **CS doublet (connection T T- brick)**

two double refreshed emulsion films, vacuum packed and glued onto the bottom of each brick



wall  
52 x 51 bricks

TT

**Total target mass : ~ 1350 t**  
(~150000 bricks, 9M emulsion plates)



# Brick Assembly Machine

Automatic lead/emulsion piling in a dark room



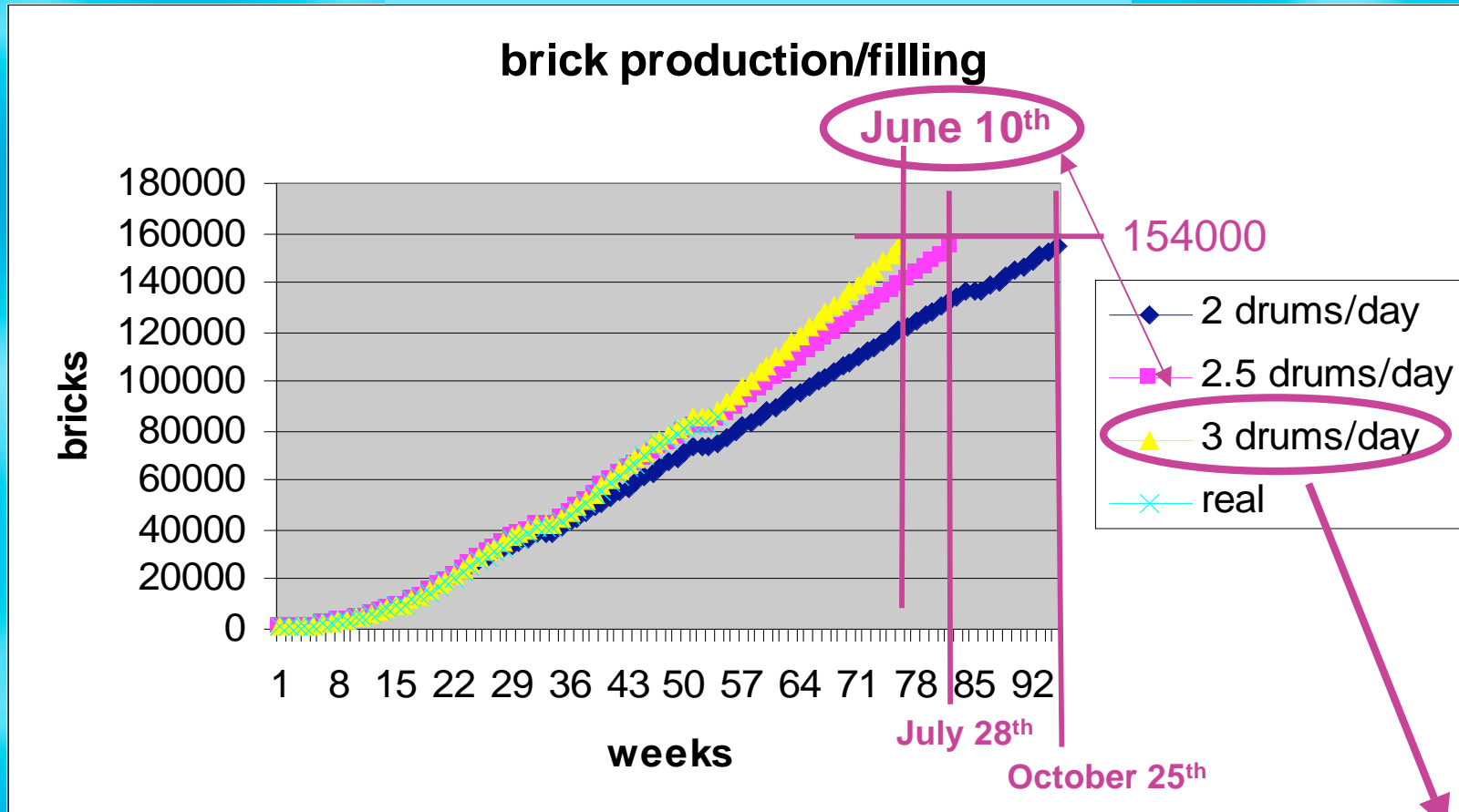
**Robots piling up bricks  
at a rate of ~ 700 bricks/day**





# Detector filling

Including 15% contingency (~3 weeks)

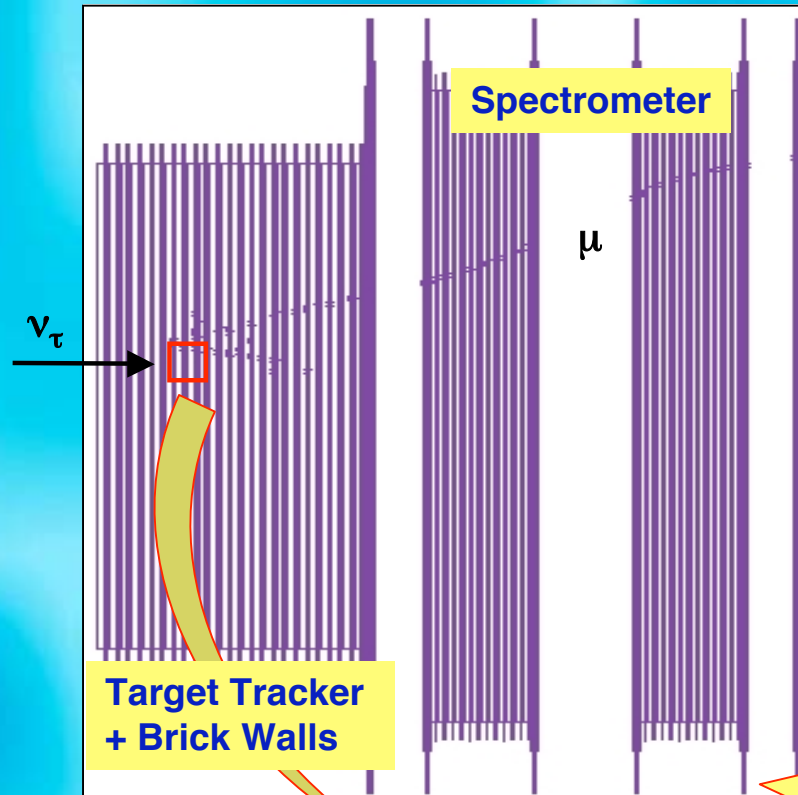


1 drum = 230 bricks

Achieved

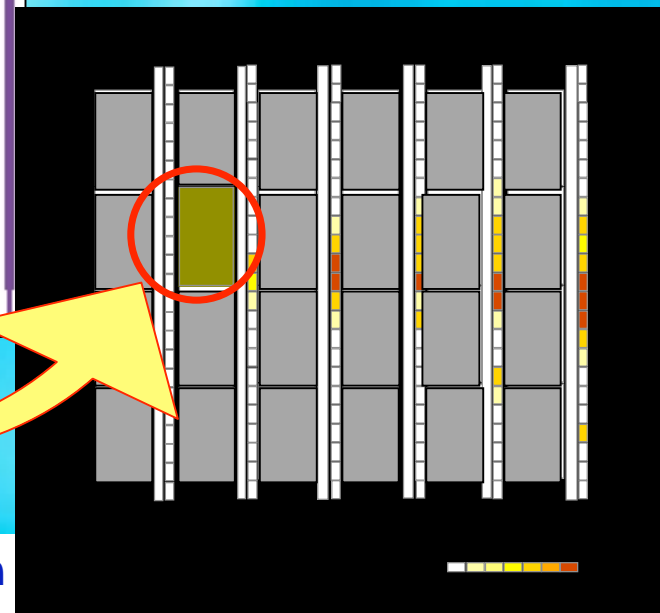


# OPERA running



On-line analysis of electronic data  
Brick finding algorithm

↓  
selection of the interaction  
brick provided by time correlation  
of electronic detectors



The selected brick is removed from the target. If an interaction is detected in the CS, the brick is exposed to cosmic rays (alignment) and the emulsions are developed and sent to scanning stations/labs

Event location  
Decay search  
Particle ID and kinematics  
Event classification





# LNGS: Development Facility



Line for the automated emulsion development

5 development chains ready



Maximum rate:  
~ 50 bricks/day  
(16 hours)



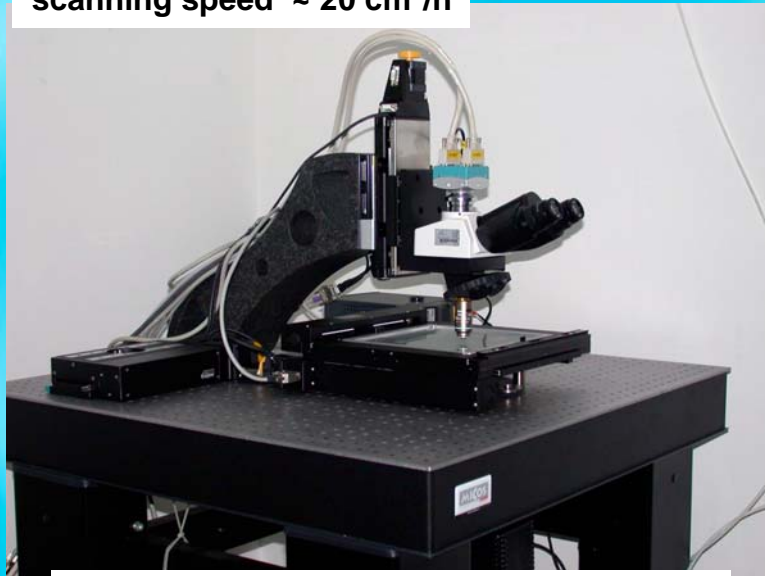
# Automated microscopes for nuclear emulsions

20 ~ 30 bricks will be daily extracted from the target and analyzed using high-speed automated systems

European Scanning System

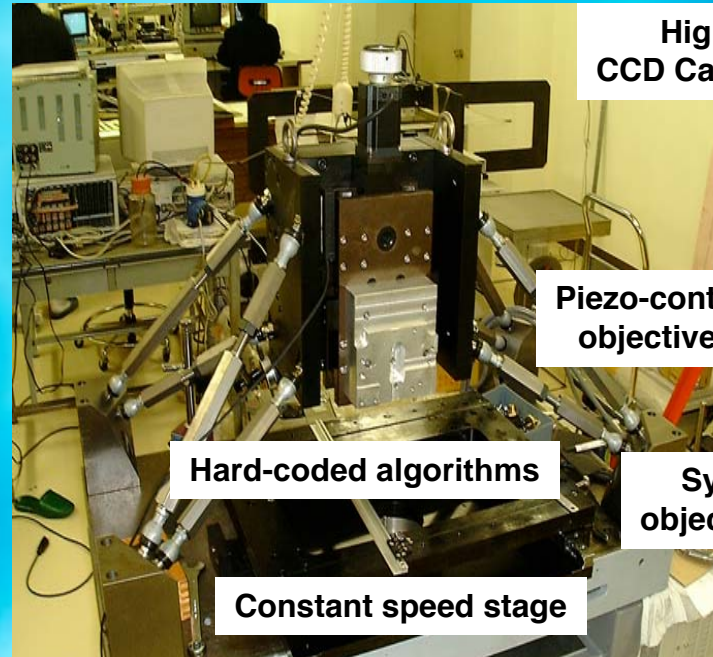
S-UTS (Japan)

scanning speed ~ 20 cm<sup>2</sup>/h



Customized commercial optics and mechanics + asynchronous DAQ software

MTM



High speed  
CCD Camera (3 kHz)

Piezo-controlled  
objective lens

Hard-coded algorithms

Synchronization of  
objective lens and stage

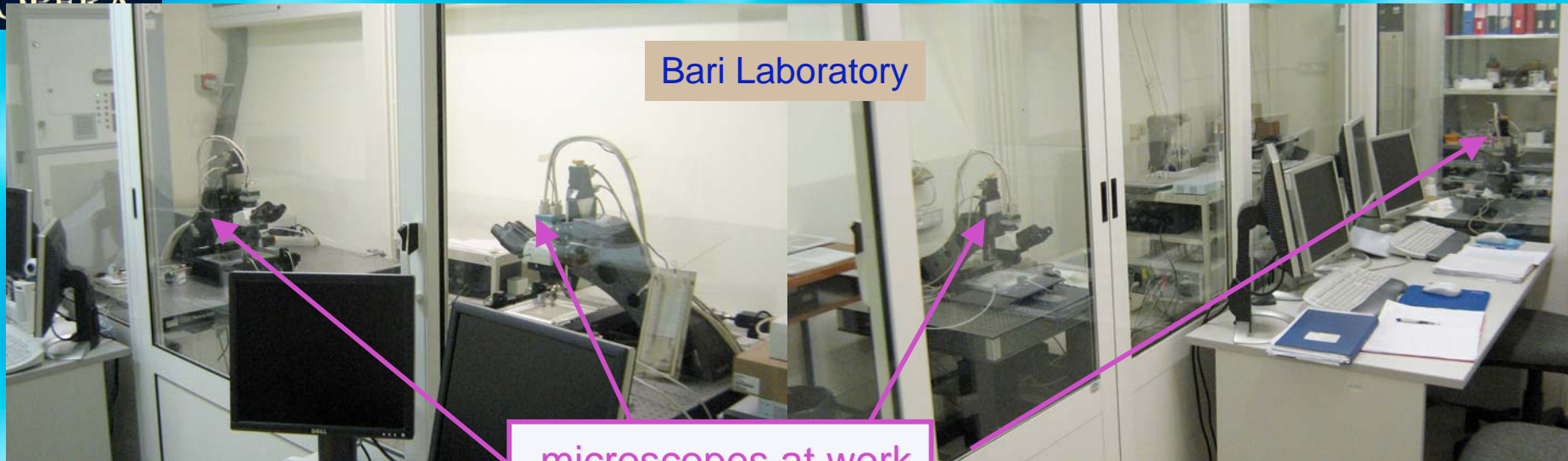
Constant speed stage

Venice, April 17, 2008

18



# Scanning Laboratories



Bari Laboratory

microscopes at work



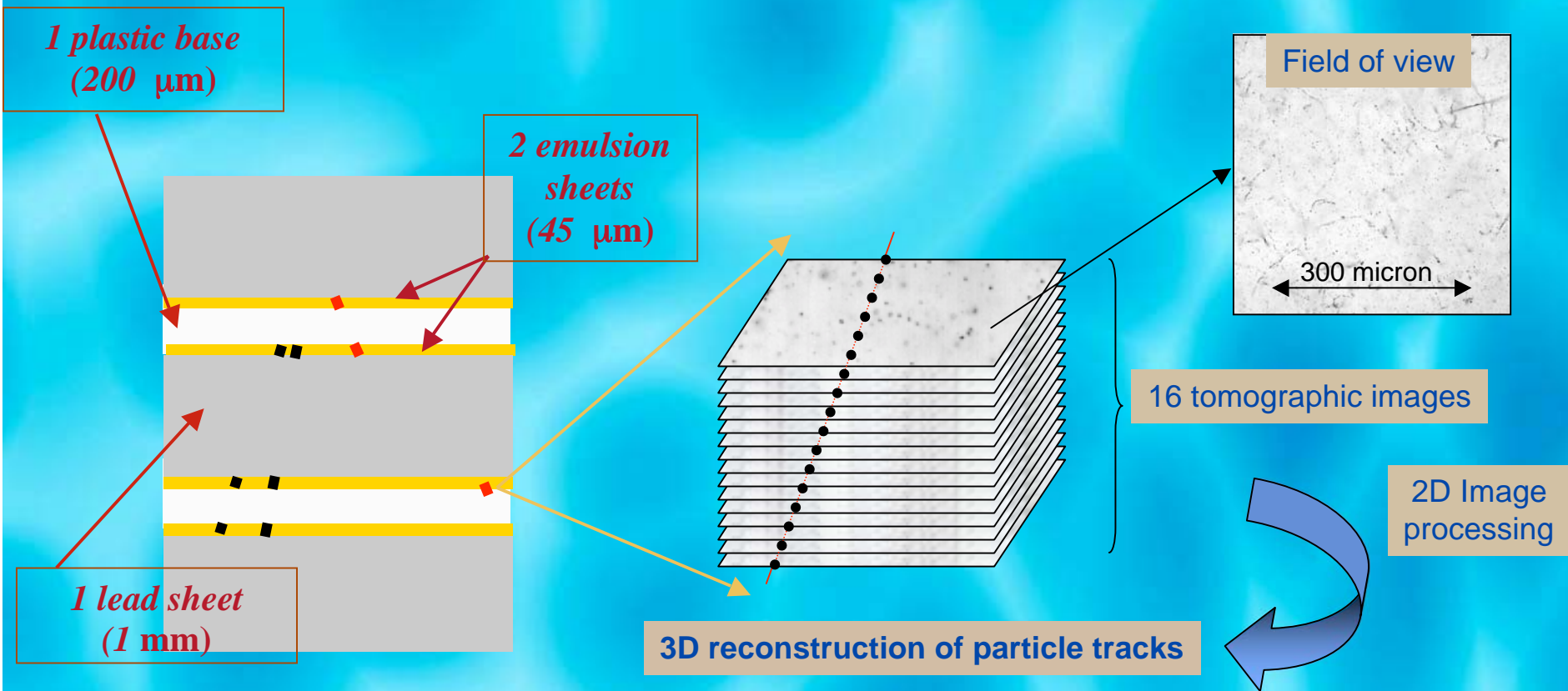
Bern Laboratory



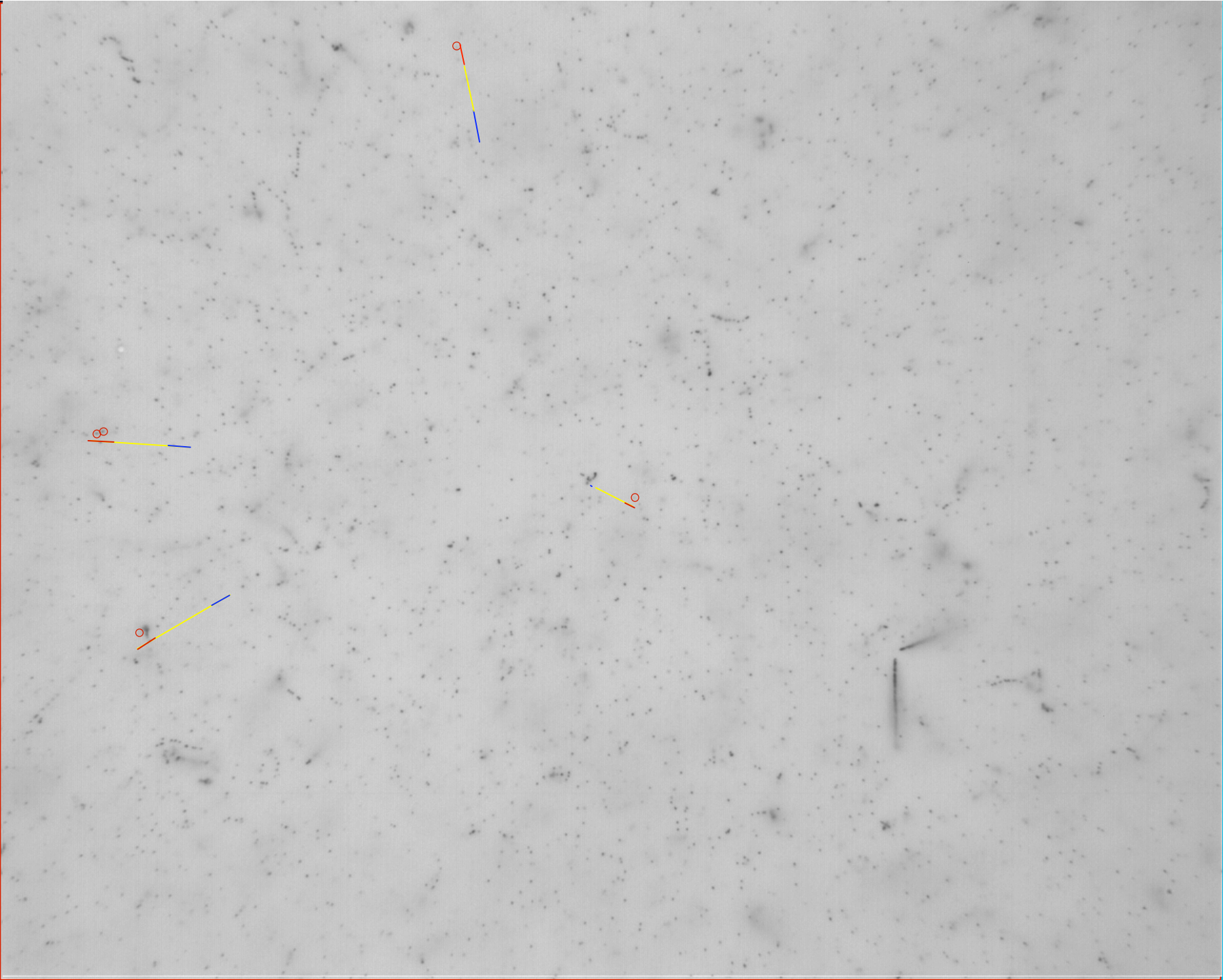
# Scanning

Fully automated scanning at high speed ( $> 20 \text{ cm}^2/\text{h}$ )

Spatial resolution  $< 1 \mu\text{m}$   
Angular resolution  $< 2 \text{ mrad}$



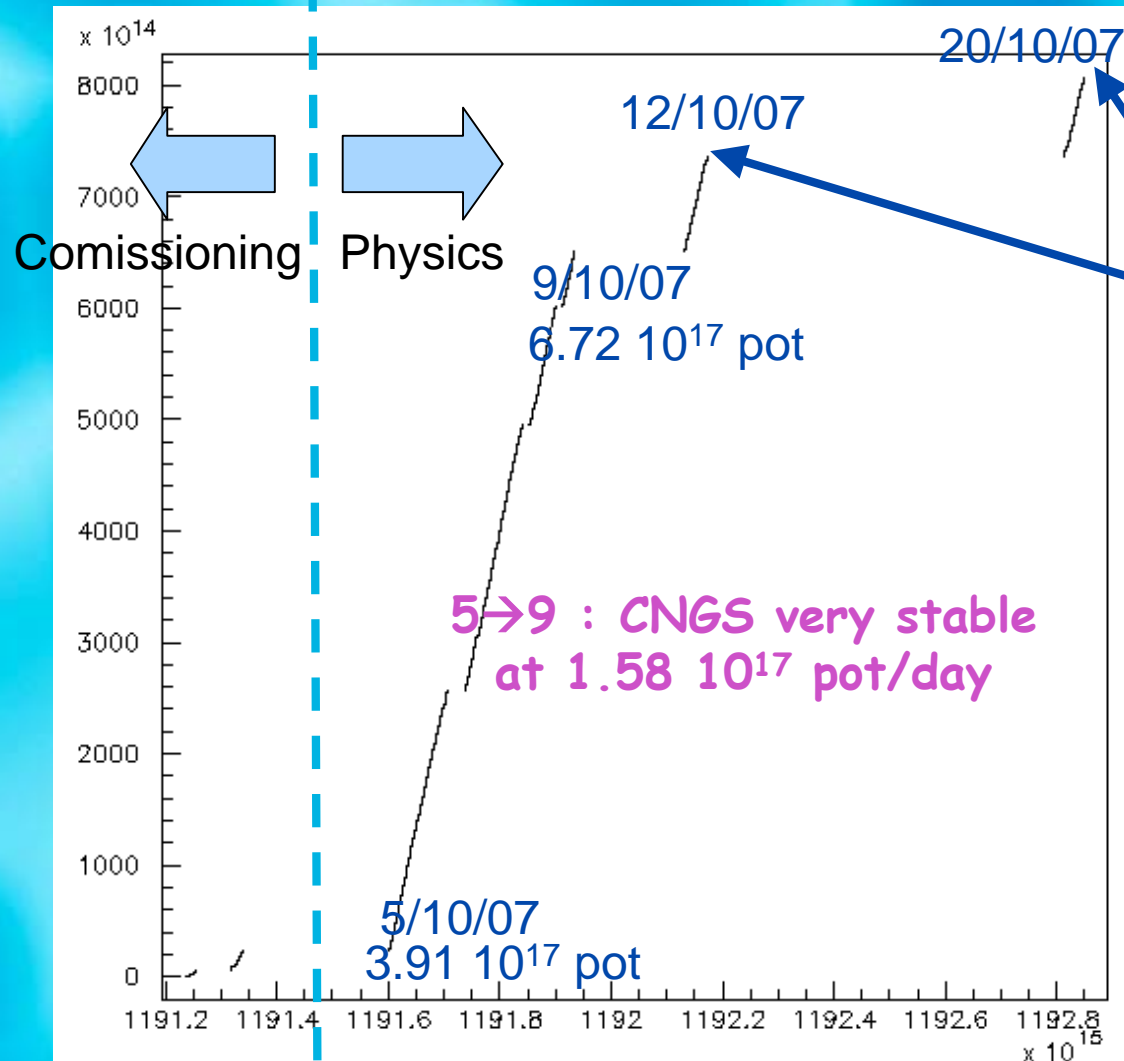






# CNGS 2007 RUN

CNGS integrated intensity from Sept. 24<sup>th</sup> to Oct 20<sup>th</sup> 2007:  $8.24 \cdot 10^{17}$



Problems in ventilation control units of the proton target

Nominal intensity :  
 $4.5 \cdot 10^{19}$  pot /yr (1yr=200 days)

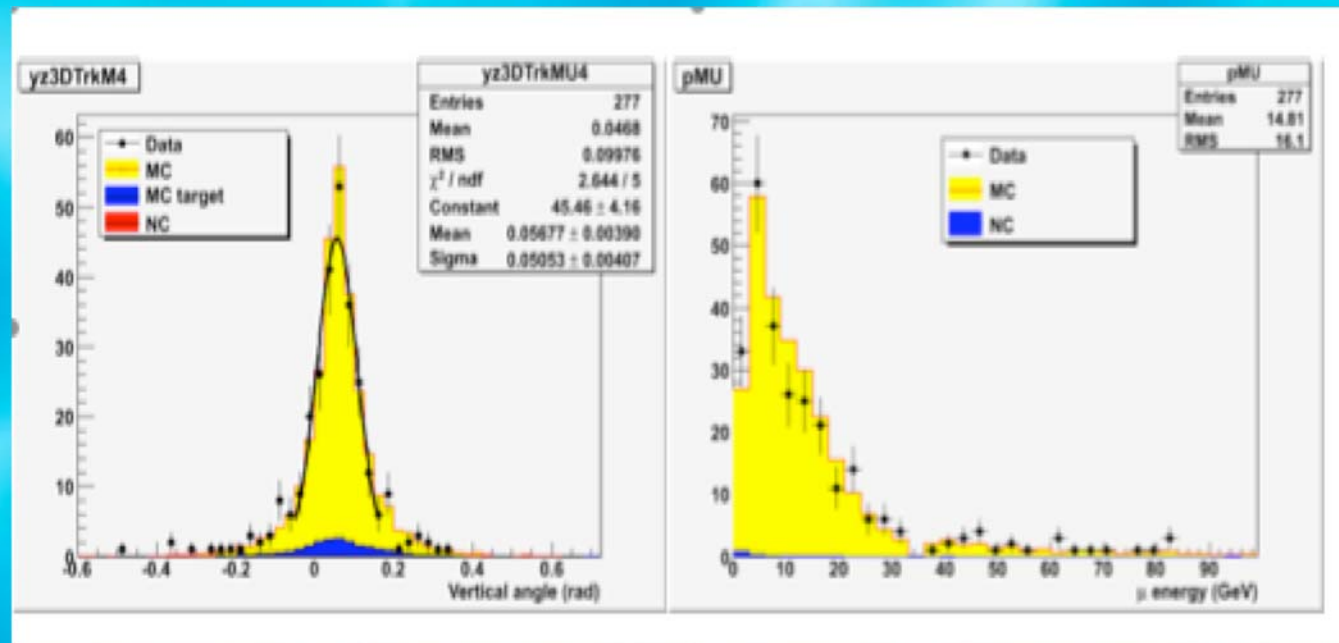


# 2007 run: Interactions outside the Target

331 events passed the analysis cut  
303 had been expected:  
ratio similar to the one observed in 2006



CNGS operated  
in the same conditions as in  
2006 run



Muon vertical angle

Energy of the muon (RPC)





## 2007 run : $\nu$ Interactions in the Target

$8.24 \cdot 10^{17}$  p.o.t.

$32 \pm 6$  expected events in bricks

**38** events registered during the  
2007 CNGS run :

29 CC    9 NC

**Brick handling, Film Processing, Scanning : first test on real  
neutrino interactions**



## Summary of the CS measurements

LNGS- European Scanning Station	All 19	CC 15	NC 4
No candidate on CS	1	1*	0
Good candidate on CS	18	14	4

\*fiducial volume acceptance

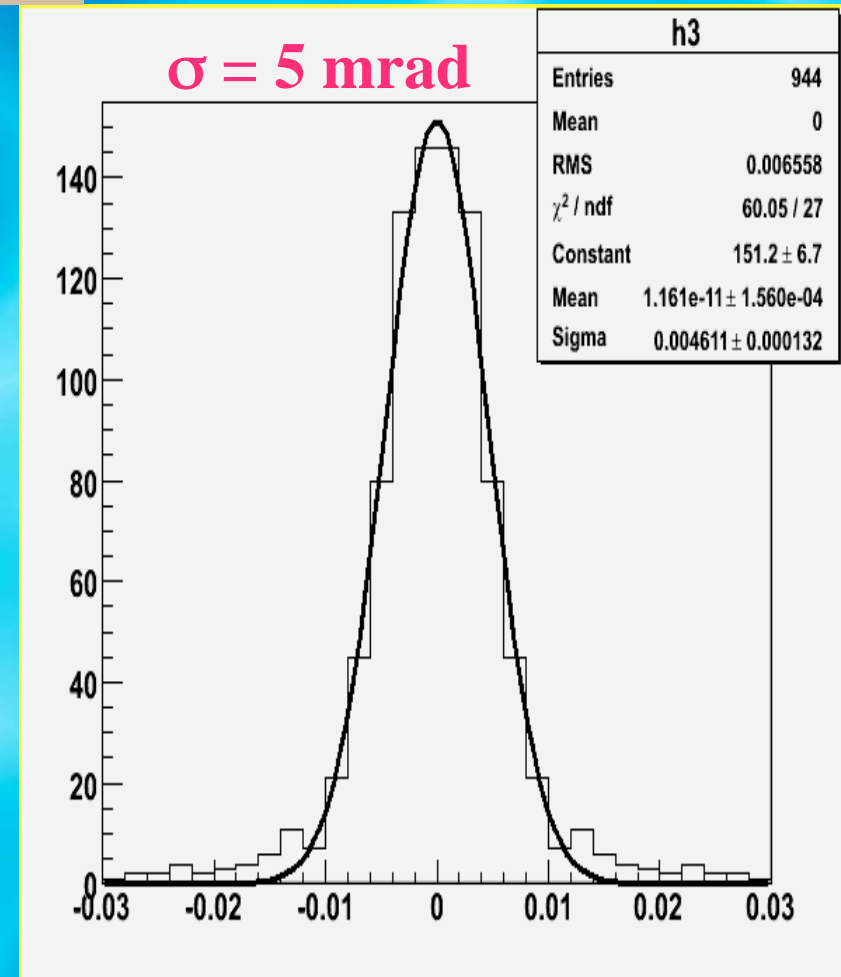
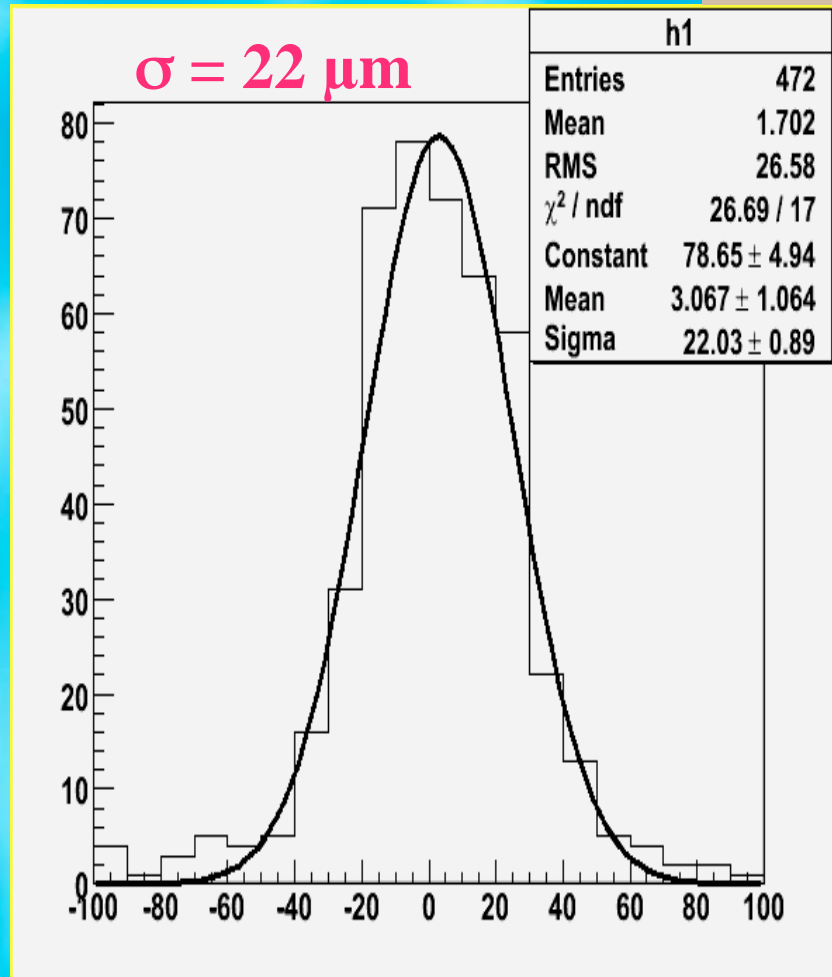
Japan - Nagoya Laboratory	All 19	CC 14	NC 5
No candidate on CS	1	0	1
Good candidate on CS	18	14	4

36 out of the 38 events have a good CS tagging  
most of the predictions were found in ~ 2-3 centimeters



# CS to brick: design accuracy in position and angle

## Test-beam

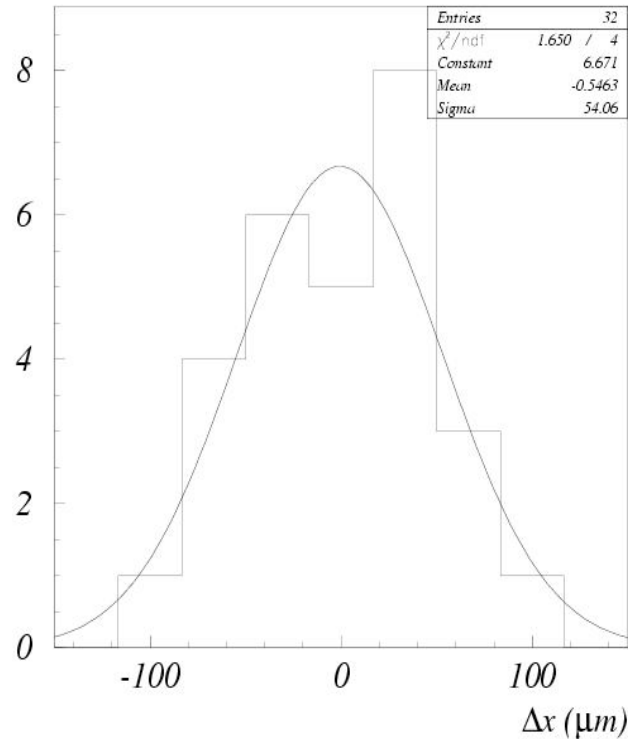




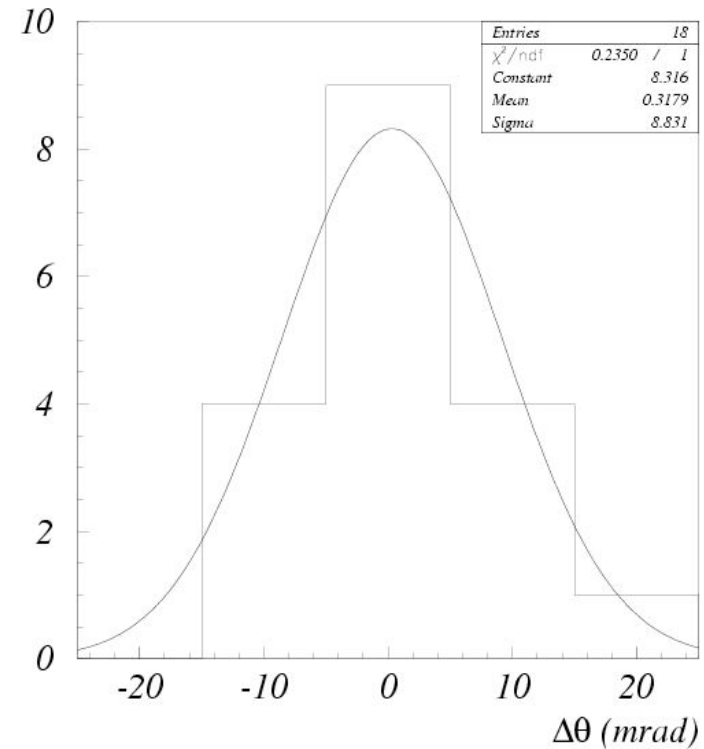
# CS to brick connection for neutrino interactions

Data quality looks fine

$\sigma = 54 \mu\text{m}$



$\sigma = 9 \text{ mrad}$



Systematical uncertainties (gap, marks, ...) dominate



## Event Location & Data analysis

	Europe	Japan
All events	19 (15 CC, 4 NC)	19 (14 CC, 5 NC)
Events confirmed in the CSd	18	18
Events located in the bricks	14	10
Interactions in dead material	1	1
CS-Brick mismatch	1	1
Analysis in progress	2	6

For **28** events out of 36,  
analysis is almost completed  
On this small sample

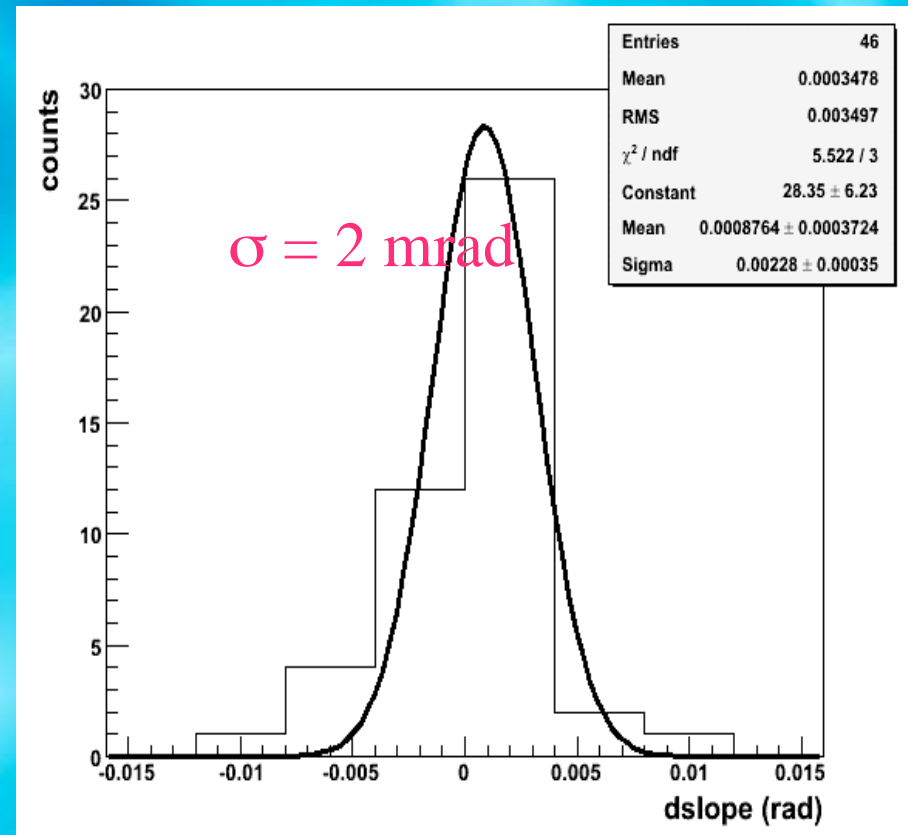
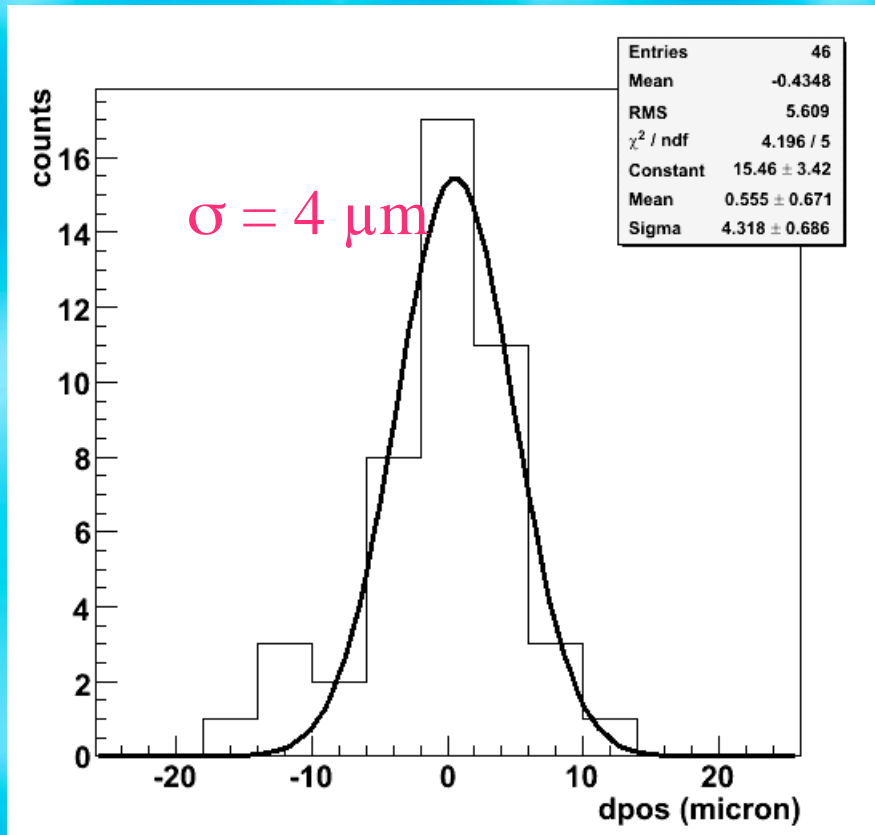


Wall Finding : > 95%  
Brick Finding : 80% +- 7%

**more data needed for a complete tuning**



# Scan-back accuracy



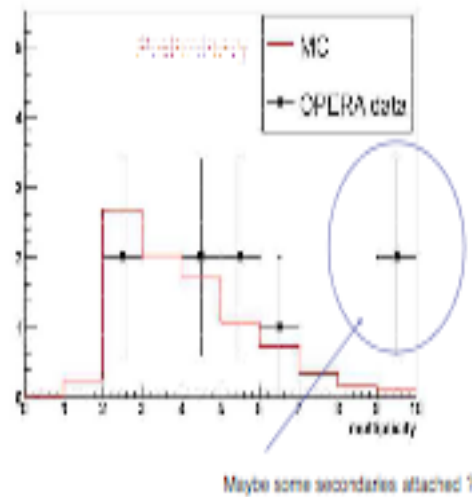


# Vertex analysis

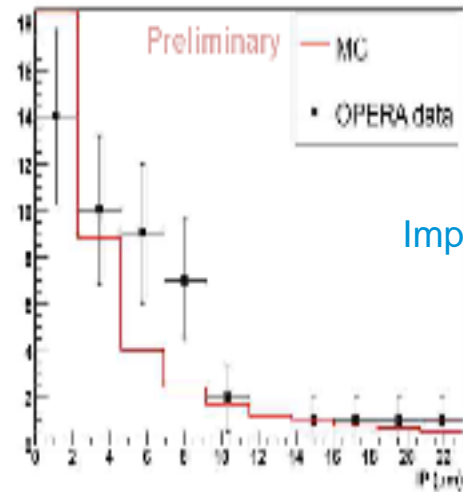
PRELIMINARY

## Kinematical reconstructed variables

Multiplicity



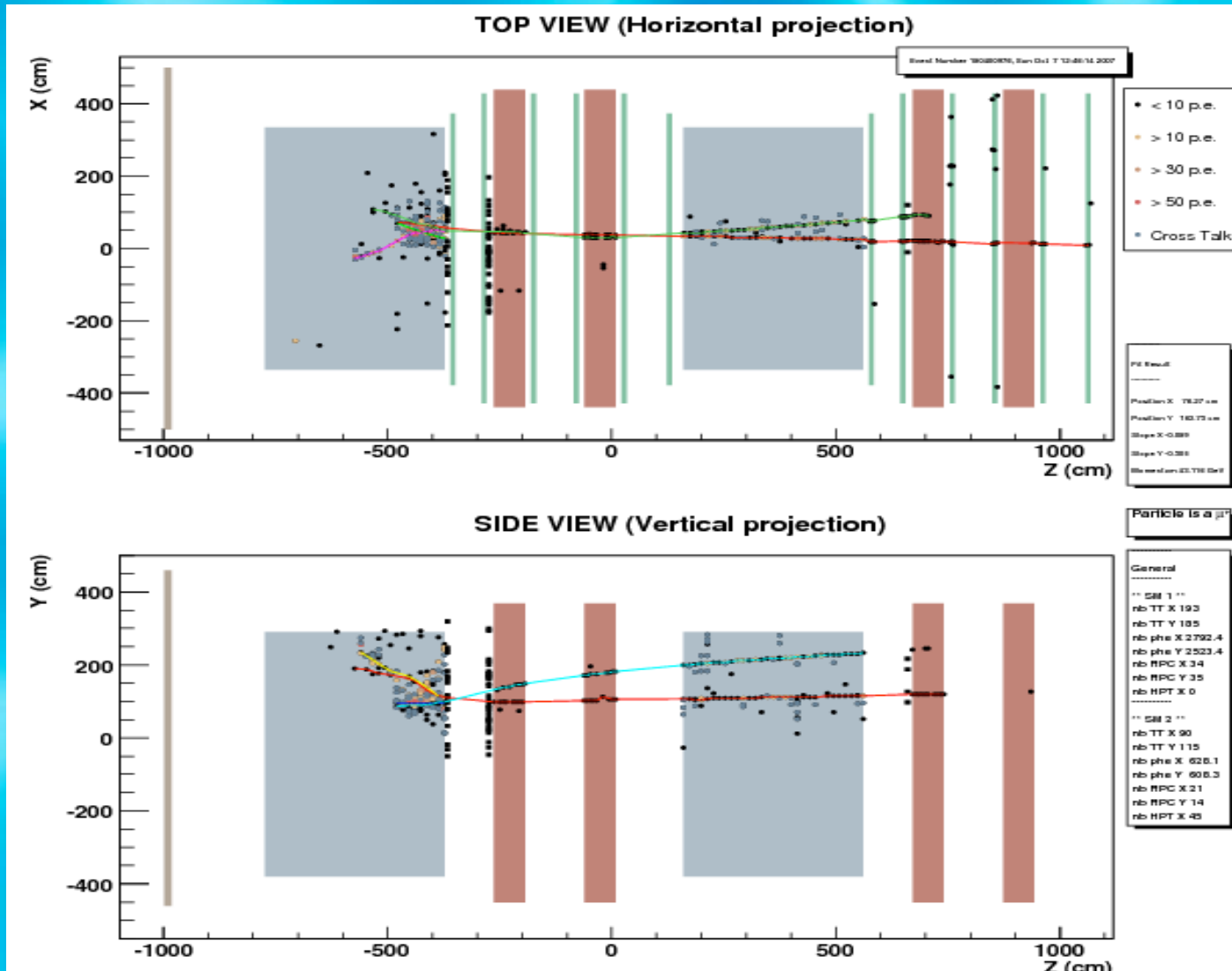
Impact parameter





# Event 180400976

CC interaction / 9 prongs







# Event 180400976

**PRELIMINARY**



Tra cks	TX	TY	P GeV	IP ( $\mu$ m)
1	0.026	-0.097	-	5.7
2	-0.024	0.019	>20	1.2
3	0.045	0.197	2.3 <sup>-0.9</sup> + 3.2	1.7
4	0.024	0.203	4.5 <sup>-2.1</sup> + 22.8	2.6
5	0.100	0.051	4.5 <sup>-2.1</sup> + 22.8	7.9
6	-0.047	0.199	2.5 <sup>-1.1</sup> + 7.0	2.9
7	0.073	0.152	3.0 <sup>-1.2</sup> + 7.9	8.8
8	0.085	0.254	1.6 <sup>-0.7</sup> + 4.3	10.3
9	0.074	0.201	1.8 <sup>-0.8</sup> + 6.0	4.6
10	0.114	0.010		9.0
11	0.132	-0.001		18.1

kink

mu

e-pair

### Secondary Vertex

Daughter momentum = 1.9<sup>+0.7</sup><sub>-0.4</sub>

$\theta_{\text{kink}} = 0.031$  rad

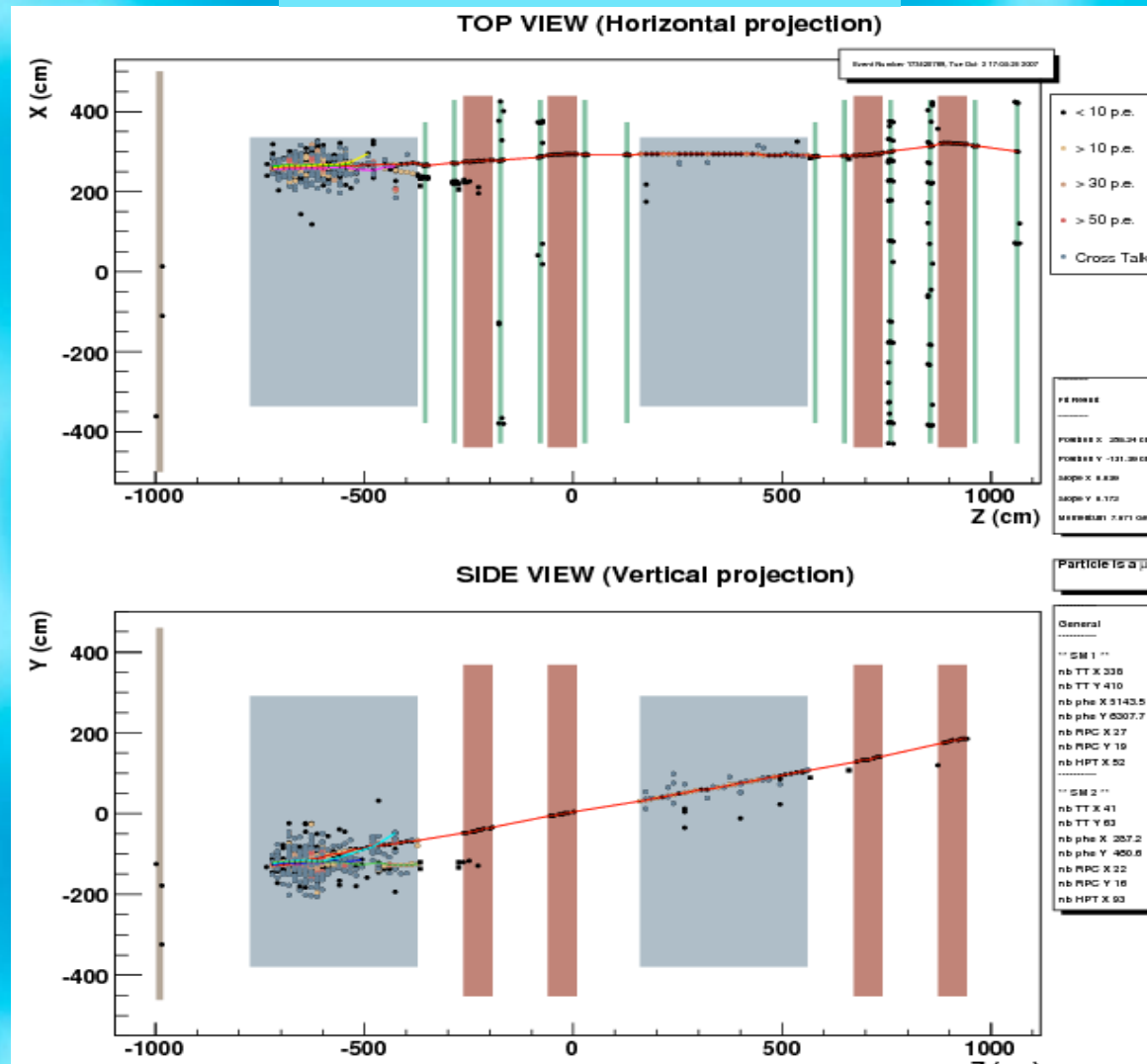
Flight length = 5901  $\mu\text{m}$

$P_T = 59$  MeV



# Event 173520769

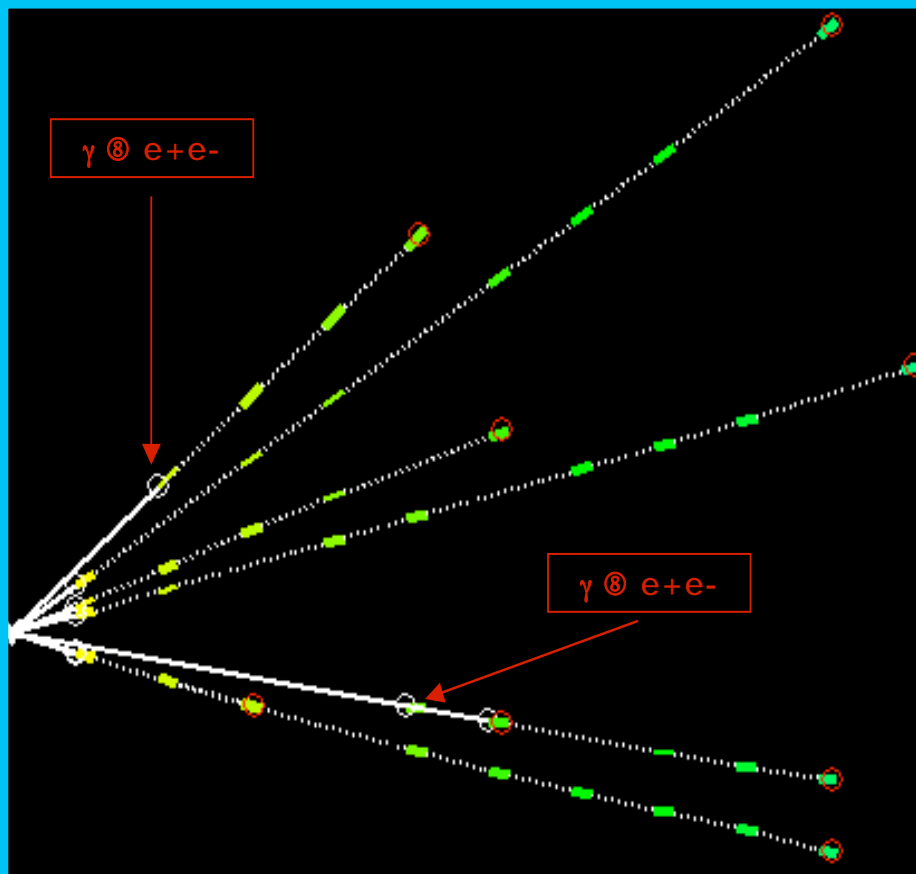
## CC interaction / 4 prongs





# Event 173520769

PRELIMINARY



Tra cks	TX	TY	IP ( $\mu$ m)
1	0.012	0.160	1.9
2	-0.739	0.095	4.7
3	-0.033	-0.060	6.3
4	-0.015	0.065	4.1

$E_{\gamma_1} : \sim 510 \text{ MeV}$

$E_{\gamma_2} : \sim 260 \text{ MeV}$

$\theta_{\gamma\gamma} : 300 \pm 20 \text{ mrad}$

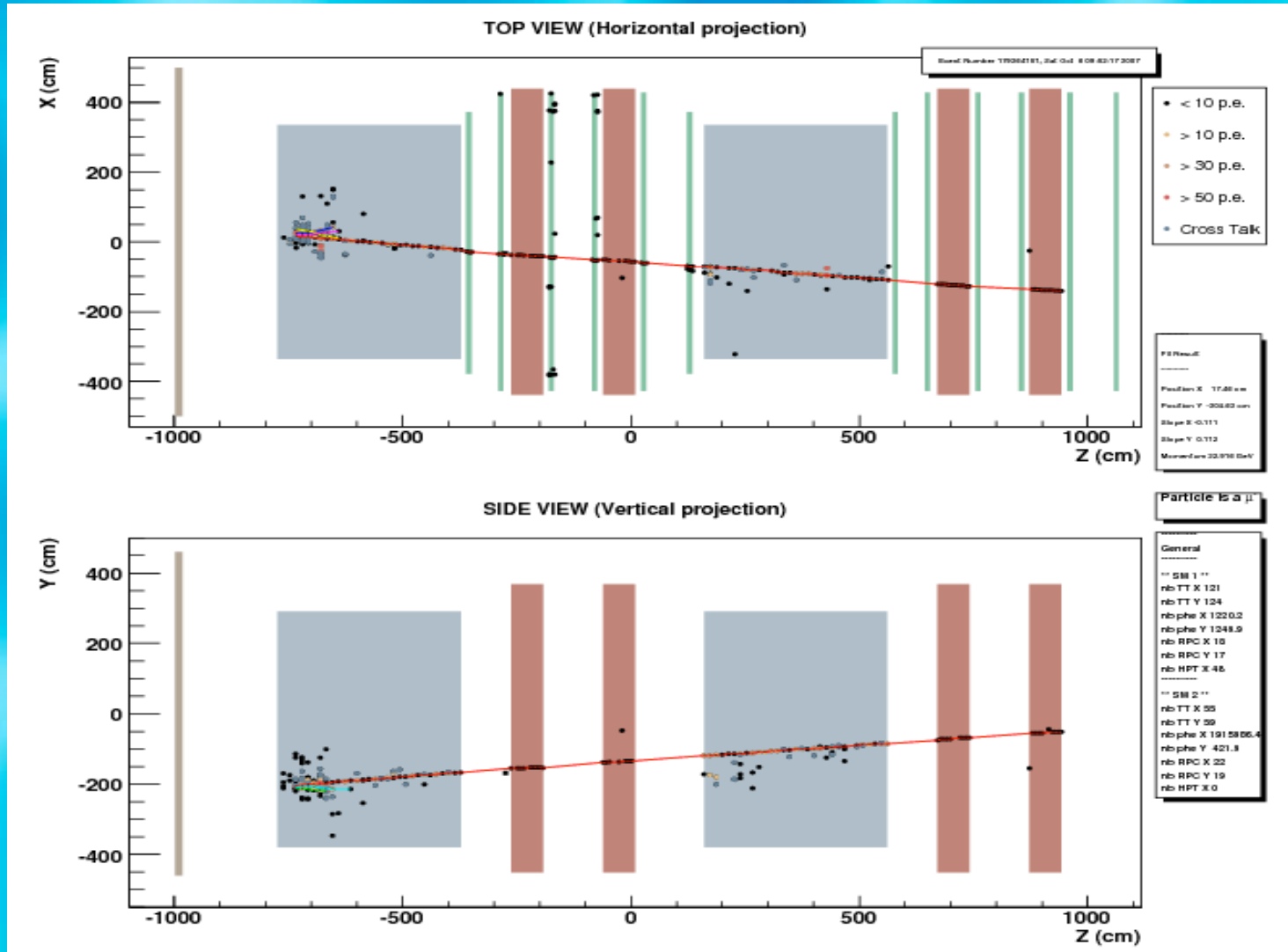
compatible with  $\pi^0 \rightarrow 2 \gamma$  :

estimated  $\pi^0$  mass:  $110 \pm 30 \text{ MeV}$



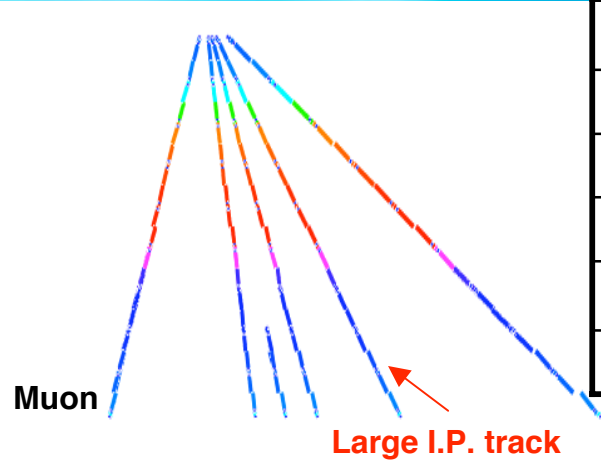
# Event 179264151

CC interaction / 5 prongs



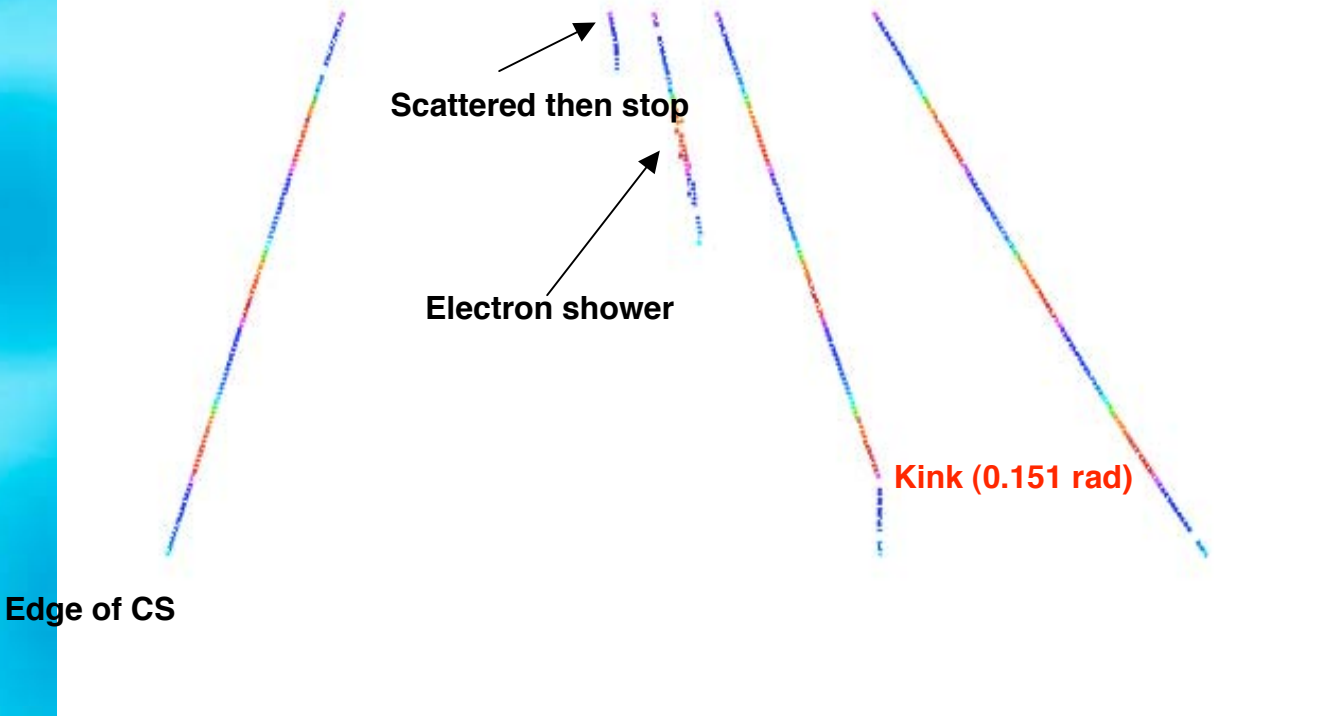


**PRELIMINARY**



Track	TX	TY	I.P. (micron)
1	0.218	0.110	0.6
2	0.049	0.002	7.0
3	-0.109	0.114	muon
4	0.443	-0.005	0.2
5	0.119	-0.059	22.6

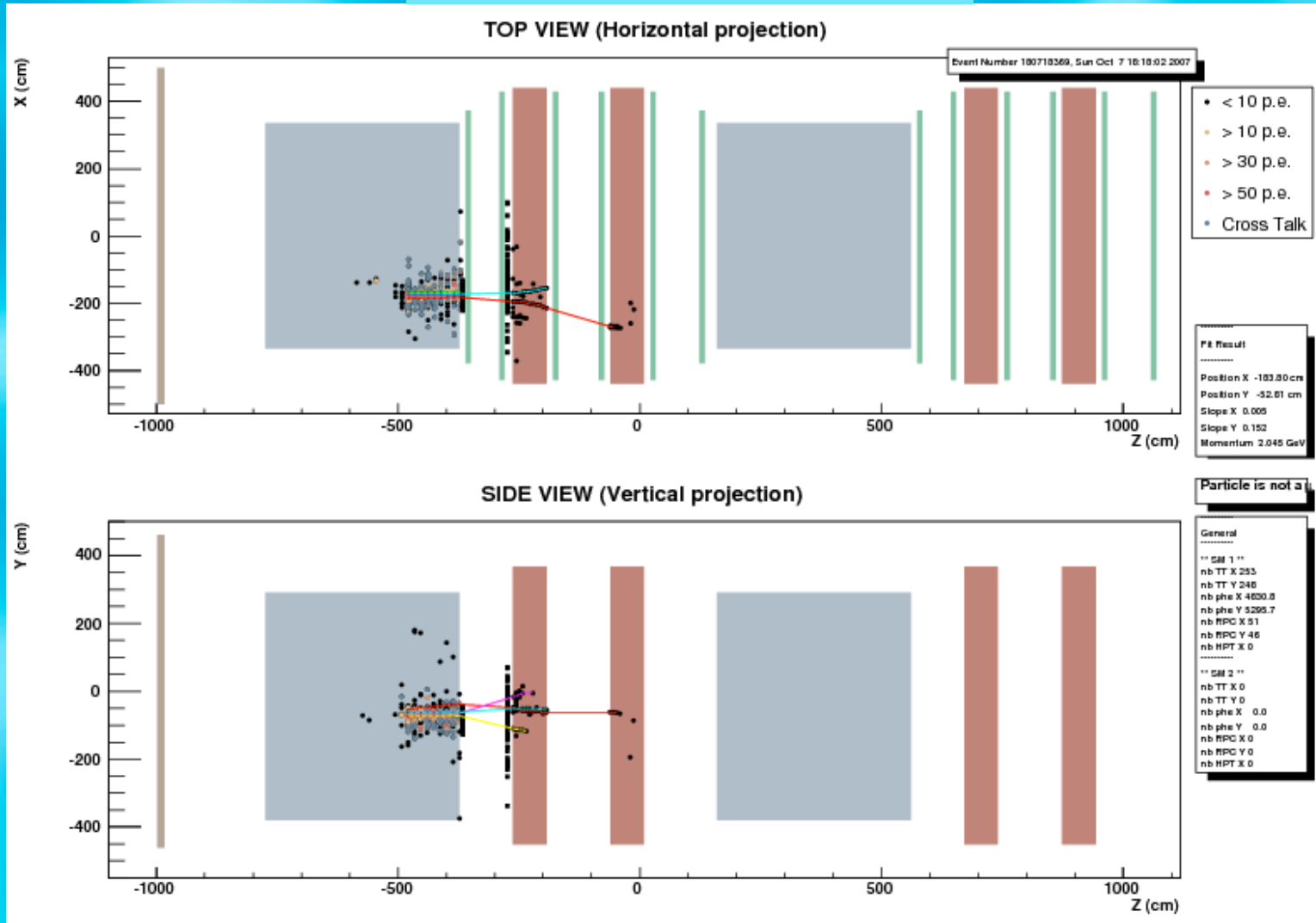
Brick to brick connection: distance 55500micron





# Event 180718369

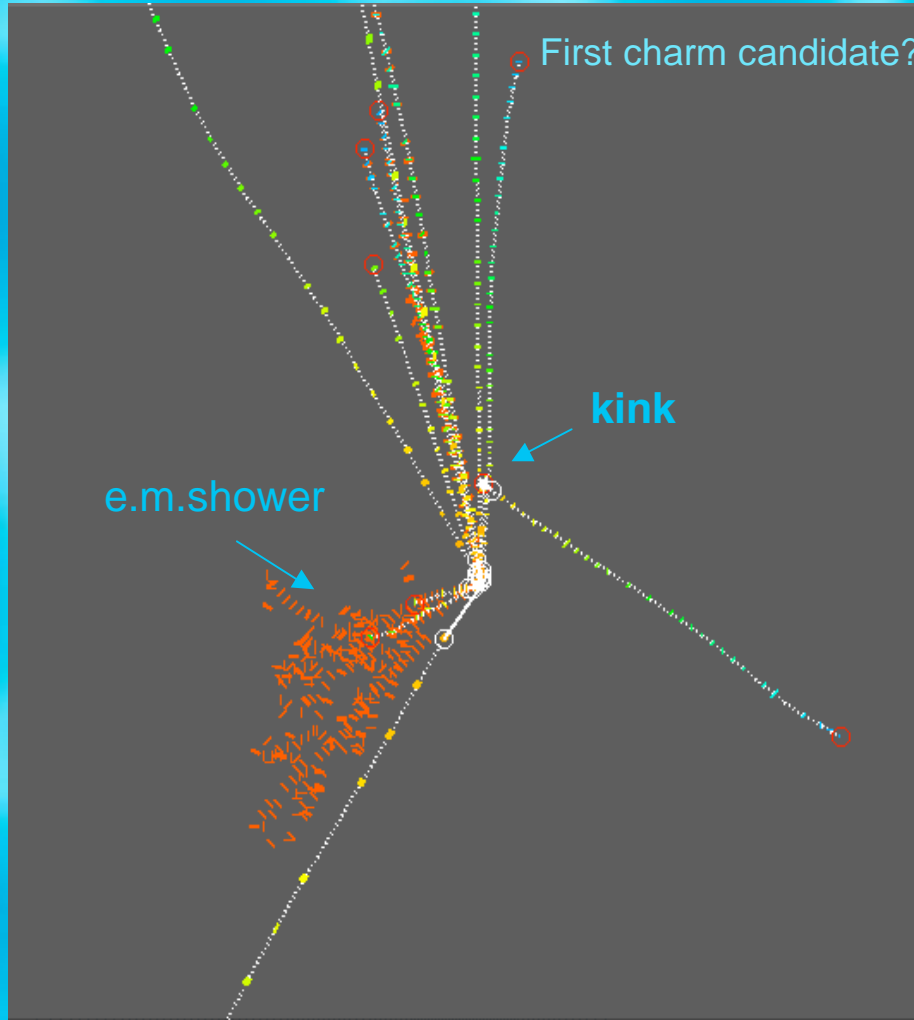
## CC interaction / 10 prongs





# Event 180718369

**PRELIMINARY**



Tracks	TX	TY	IP (μm)	Momentum (GeV)
1	0.005	0.036	3.3	$1.7^{+0.5}_{-0.3}$
2	0.005	0.139	1.0	parent
3	0.002	0.064	6.6	>20.0
4	-0.021	0.064	7.1	$2.1^{+0.7}_{-0.4}$
5	-0.029	0.046	2.8	>8.4
6	-0.031	0.064	7.3	$2.4^{+0.8}_{-0.5}$
7	-0.076	0.068	4.2	$1.8^{+1.6}_{-0.6}$
8	-0.089	0.141	6.9	$2.5^{+1.4}_{-0.7}$
9	-0.183	0.106	5.4	$0.7^{+0.2}_{-0.1}$
10	-0.297	-0.143	19.2	$0.7^{+0.3}_{-0.1}$
11	-0.067	0.008	7.3	$3.5^{+3.6}_{-1.2}$ e.pair
12	-0.069	0.005	16.8	$2.0^{+3.1}_{-0.8}$ e.pair

### Secondary Vertex

Daughter momentum =  $3.9^{+1.7}_{-0.9}$

$\theta_{\text{kink}} = 0.204$  rad

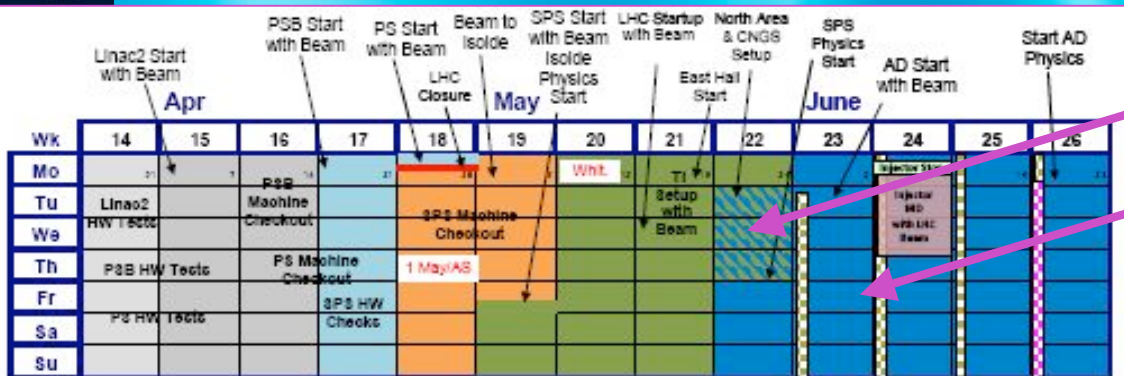
Flight length = 3247 μm

$P_T = 796$  MeV

$P_T^{\text{MIN}} = 606$  MeV (90% C.L.)



# 2008 run : SPS schedule



May 29<sup>th</sup>

CNGS: ~June 20<sup>th</sup>  
Completed repairs



130 days max for the CNGS  
Spread over 20 weeks



Nominal CNGS year : 200days



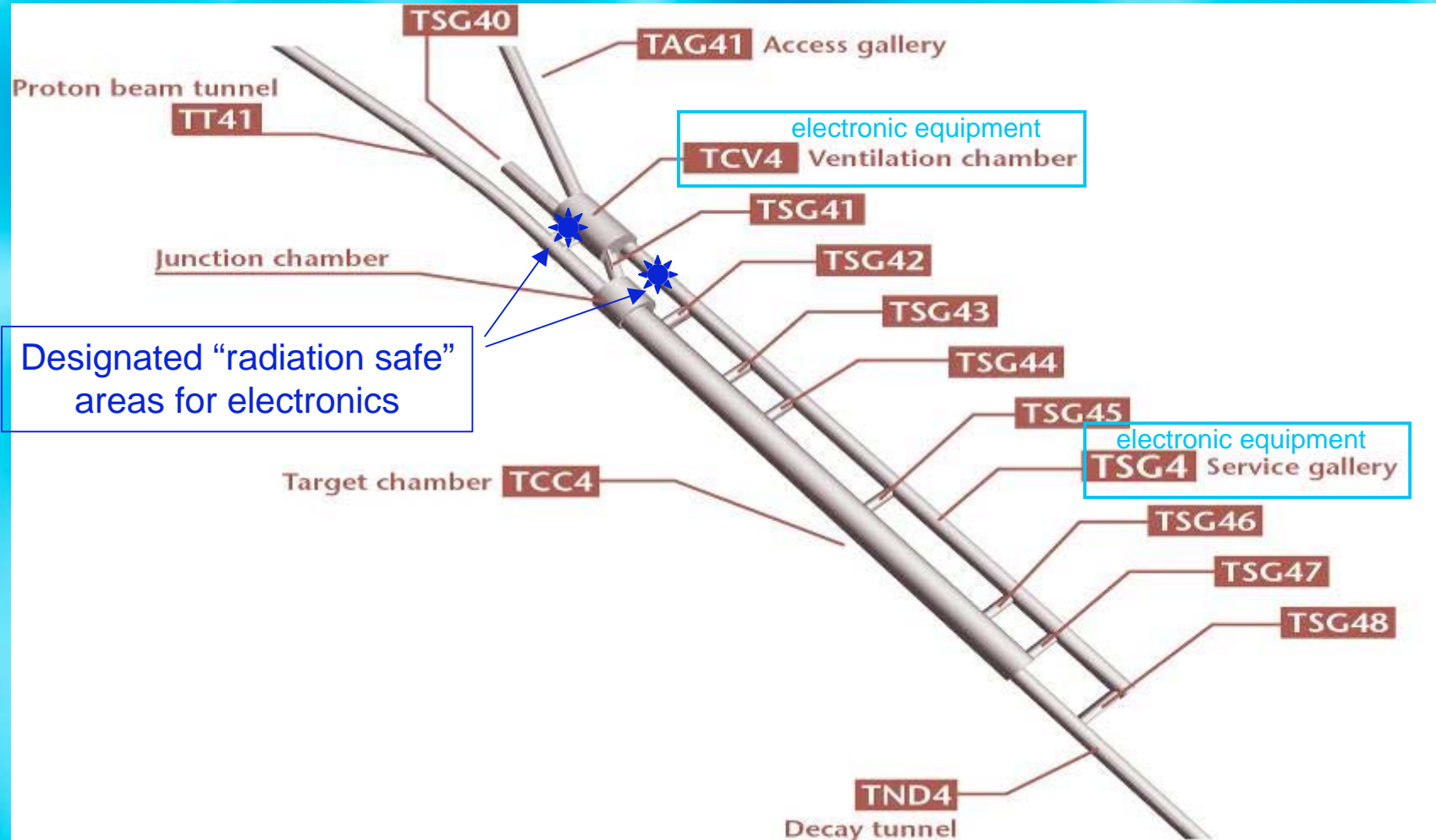
November 10<sup>th</sup>





# CNGS Electronics - underground

## Layout of the ventilation control





# Radiation effects on electronics

The failed ventilation electronics were installed along the TSG4 (service gallery), next to the ventilation ducts, and in the TCV4 (ventilation chamber).

In both areas the radiation levels, as predicted by the FLUKA simulations and confirmed by the radiation protection measurements, are far too high for COTS components → the electronics should have failed and they did



During the 2007/2008 shutdown, work is organized to remedy the problem and assure nominal running of the facility for 2008 and beyond.



# Actions for 2008 shutdown

## **General guideline:**

Replace damaged electronics

Move the electronics out of the CNGS tunnels as possible

## **For the equipment which must stay in the area:**

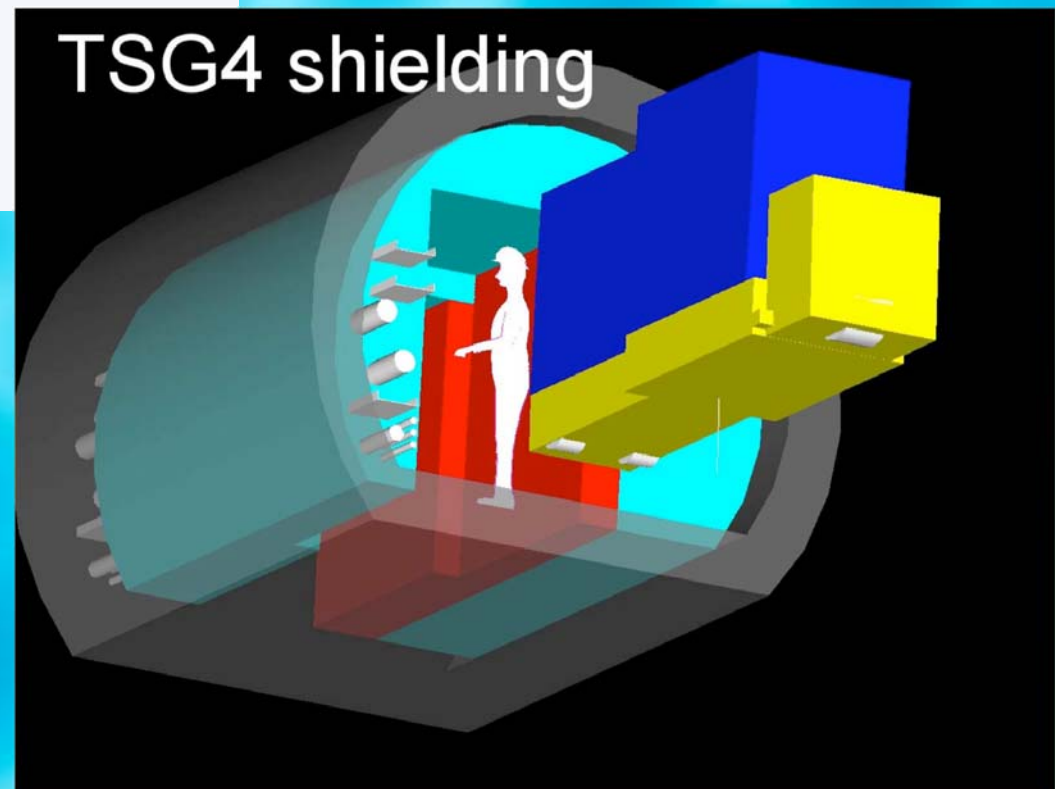
- 1) Create a “radiation-safe” area by adding adequate shielding and move all the electronics into this area
- 2) Address the sensitivity to radiation of the installed electronics and investigate upgrade possibilities
- 3) Install a radiation monitoring system for electronics as in LHC



# CNGS repair state

## Radiation damages repair and protection:

the work is fully funded  
electronics repair and reshuffling has started  
shielding plugs defined, civil engineering starting  
completion of the work expected by the end  
of week 23 (June 6th)  
stable beam during week 25  
(~June 20th)





# 2008 run : expected statistics

	Expected 2008 run	Nominal CNGS
Number of days	130	200
Super Cycle (sec.)	48	27.6
Circulating proton	$2 \times 2 \cdot 10^{13}$	$4.8 \cdot 10^{13}$
efficiency	80%	55%
Integrated pot	$2.1 \cdot 10^{19}$	$4.5 \cdot 10^{19}$

Minimal granted by CERN  
If no delay in the CNGS repair

Total number of interactions	$\sim 2000$
Tau candidate (@ $2.5 \cdot 10^{-3} \text{ eV}^2$ )	$\sim 1$

$\sim 110 \text{ evts/week}$

no MTE running (only commissioning) in 2008



# Conclusions

Neutrino events have been observed in the OPERA bricks during the short CNGS run of 2007:  
With the collected small sample, several issues were faced and solved

Brick candidate extraction, brick handling, CS scanning , vertex location: tested the full chain

Analysis and kinematical reconstruction of the neutrino interactions in emulsion are in progress

**Next CNGS run with full target expected by june 2008 : First OPERA year !**



The detector is ready for the next phase

...waiting for the first  
 $\nu_\tau$  interaction  
in emulsion.....

