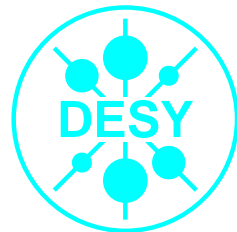


Particle Interpretations of the PVLAS Data

Andreas Ringwald



XII International Workshop on “Neutrino Telescopes”
March 6-9, 2007
Venice, Italy

Plan:

1. Vacuum Magnetic Dichroism and Birefringence

Polarized light propagation through a magnetic field: global data

2. Possible Explanations

Production of new weakly interacting light particles (WILPs)?

3. Crucial Laboratory Tests

Light or dark-current through-a-wall experiments, ...

4. Problems of Particle Interpretations

Astrophysical, cosmological, and other constraints

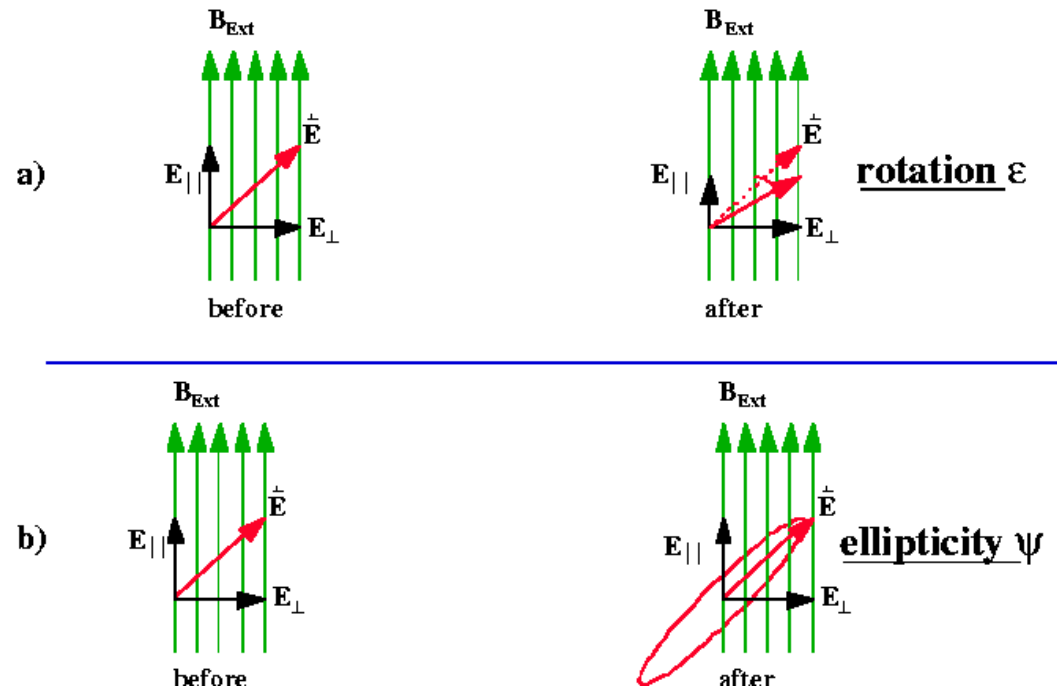
5. WILPs in Models with Light Extra-U(1)s

Light mini-charged particles from gauge kinetic mixing, ...

6. Summary

1. Vacuum Magnetic Dichroism and Birefringence

- Send linearly polarized laser beam through transverse magnetic field \Rightarrow measure changes in polarization state:
 - rotation (dichroism)
 - ellipticity (birefringence)



[Brandi *et al.* '01]

1. Vacuum Magnetic Dichroism and Birefringence

- Send linearly polarized laser beam through transverse magnetic field \Rightarrow measure changes in polarization state:
 - rotation (dichroism)
 - ellipticity (birefringence)

BFRT experiment: [Cameron *et al.* '93]

(Brookhaven, Fermilab, Rochester, Trieste)

$$B \sim 2 \text{ T}, \ell = 8.8 \text{ m}, \omega = 2.4 \text{ eV}, N_{\text{pass}} = 34 - 254$$

PVLAS experiment: [Zavattini *et al.* '06]

$$B = 5 \text{ T}, \ell = 1 \text{ m}, \omega = 1.2 \text{ eV}, N_{\text{pass}} = 44000$$

Q&A experiment: [Chen, Mei, Ni '06]

$$B = 2.3 \text{ T}, \ell = 1 \text{ m}, \omega = 1.2 \text{ eV}, N_{\text{pass}} = 18700$$

1. Vacuum Magnetic Dichroism and Birefringence

- Send linearly polarized laser beam through transverse magnetic field \Rightarrow measure changes in polarization state:
 - rotation (dichroism)
 - ellipticity (birefringence)
- No signal in **BFRT**

BFRT experiment		
Rotation	$(L = 8.8 \text{ m}, \lambda = 514.5 \text{ nm}, \theta = \frac{\pi}{4})$	
N_{pass}	$ \Delta\theta \text{ [nrad]}$	$\Delta\theta_{\text{noise}} \text{ [nrad]}$
254	0.35	0.30
34	0.26	0.11
Ellipticity	$(L = 8.8 \text{ m}, \lambda = 514.5 \text{ nm}, \theta = \frac{\pi}{4})$	
N_{pass}	$ \psi \text{ [nrad]}$	$\psi_{\text{noise}} \text{ [nrad]}$
578	40.0	11.0
34	1.60	0.44
Regen.	$(L = 4.4 \text{ m}, \langle\lambda\rangle = 500 \text{ nm}, N_{\text{pass}} = 200)$	
$\theta \text{ [rad]}$	rate [Hz]	
0	-0.012 ± 0.009	
$\frac{\pi}{2}$	0.013 ± 0.007	

[Cameron et al '93]

1. Vacuum Magnetic Dichroism and Birefringence

- Send linearly polarized laser beam through transverse magnetic field \Rightarrow measure changes in polarization state:
 - rotation (dichroism)
 - ellipticity (birefringence)
- No signal in **BFRT**; signal in **PVLAS**

PVLAS experiment	
Rotation ($L = 1$ m, $N_{\text{pass}} = 44000$, $\theta = \frac{\pi}{4}$)	
λ [nm]	$ \Delta\theta $ [10^{-12} rad/pass]
1064	3.9 ± 0.2
532	6.3 ± 1.0 (preliminary)
Ellipticity ($L = 1$ m, $N_{\text{pass}} = 44000$, $\theta = \frac{\pi}{4}$)	
λ [nm]	ψ [10^{-12} rad/pass]
1064	-3.4 ± 0.3 (preliminary)
532	-6.0 ± 0.6 (preliminary)

[PRL '06; IDM '06]

1. Vacuum Magnetic Dichroism and Birefringence

- Send linearly polarized laser beam through transverse magnetic field \Rightarrow measure changes in polarization state:
 - rotation (dichroism)
 - ellipticity (birefringence)
- No signal in **BFRT**; signal in **PVLAS**; no signal in **Q&A**

Q&A experiment	
Rotation ($L = 1$ m, $\lambda = 1064$ nm, $\theta = \frac{\pi}{4}$)	
N_{pass}	$\Delta\theta$ [nrad]
18700	-0.4 ± 5.3

[Q&A coll. '06]

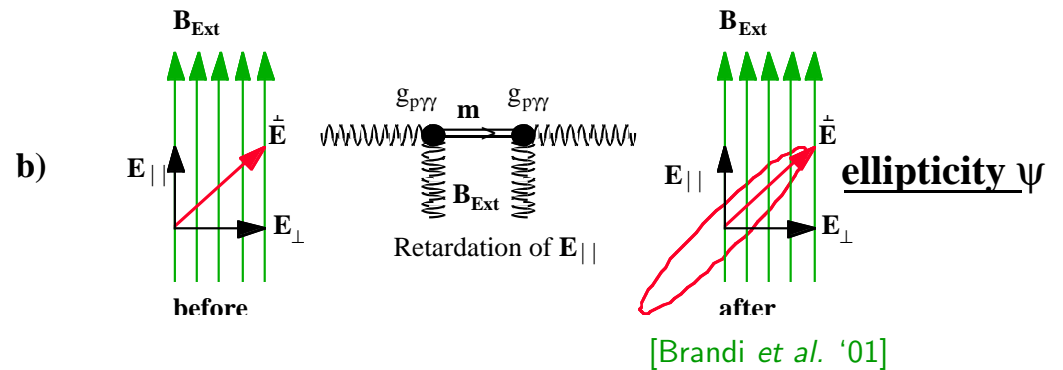
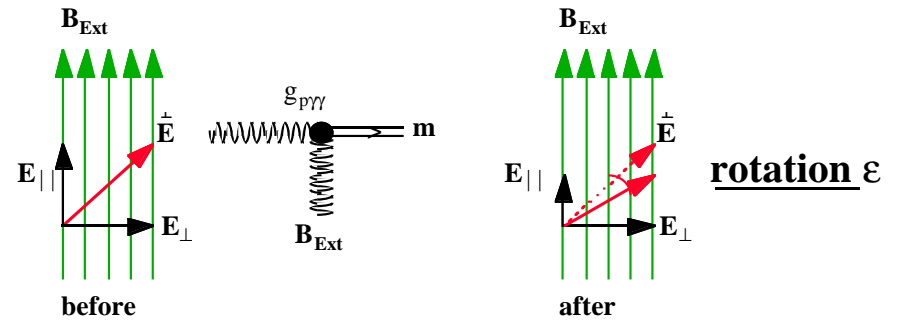
2. Possible Explanations

- Viable explanation in terms of real and virtual production of
 - light neutral spin-zero boson (Axion-Like Particle (ALP)),

$$(g/4) \phi^{(-)} F_{\mu\nu} \tilde{F}^{\mu\nu} \quad \left(\phi^{(+)} F_{\mu\nu} F^{\mu\nu} \right) \mathbf{a)}$$

Effects of Nearly Massless, Spin Zero Particles on Light Propagation in a Magnetic Field

[Maiani, Petronzio, Zavattini '86]

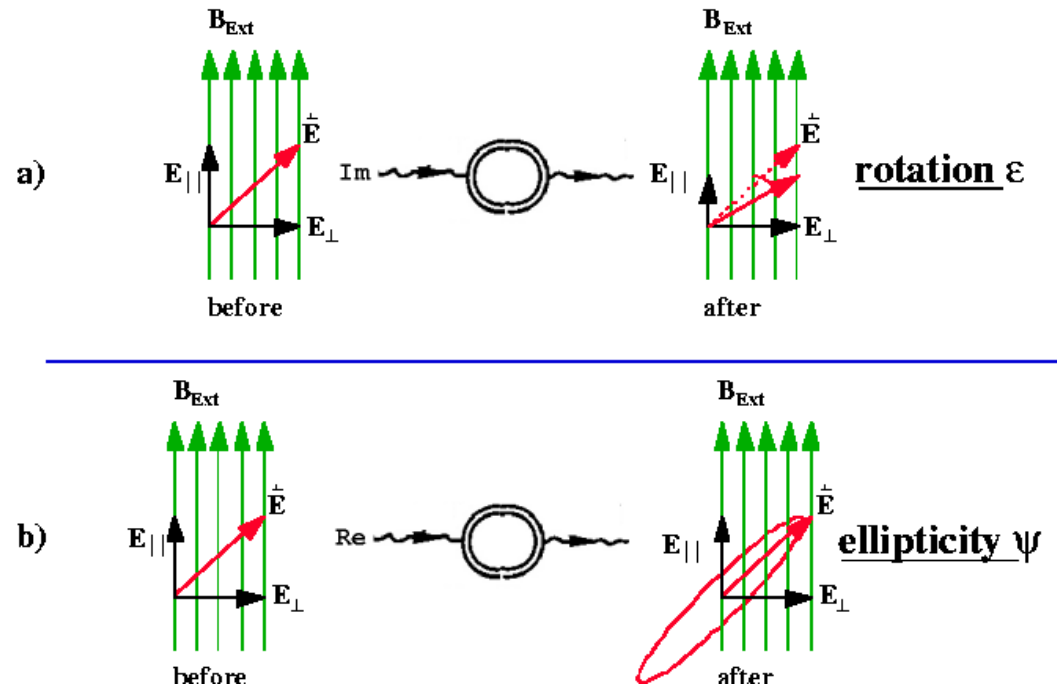


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- Viable explanation in terms of real and virtual production of
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 - anti-particle pair,

$$\partial_\mu \rightarrow \partial_\mu - i\epsilon e A_\mu$$

Polarized Light Propagating in a Magnetic Field as a Probe for Millicharged Fermions [Gies, Jaeckel, AR '06]



In analogy to theoretically well-studied e^+e^- real and virtual production

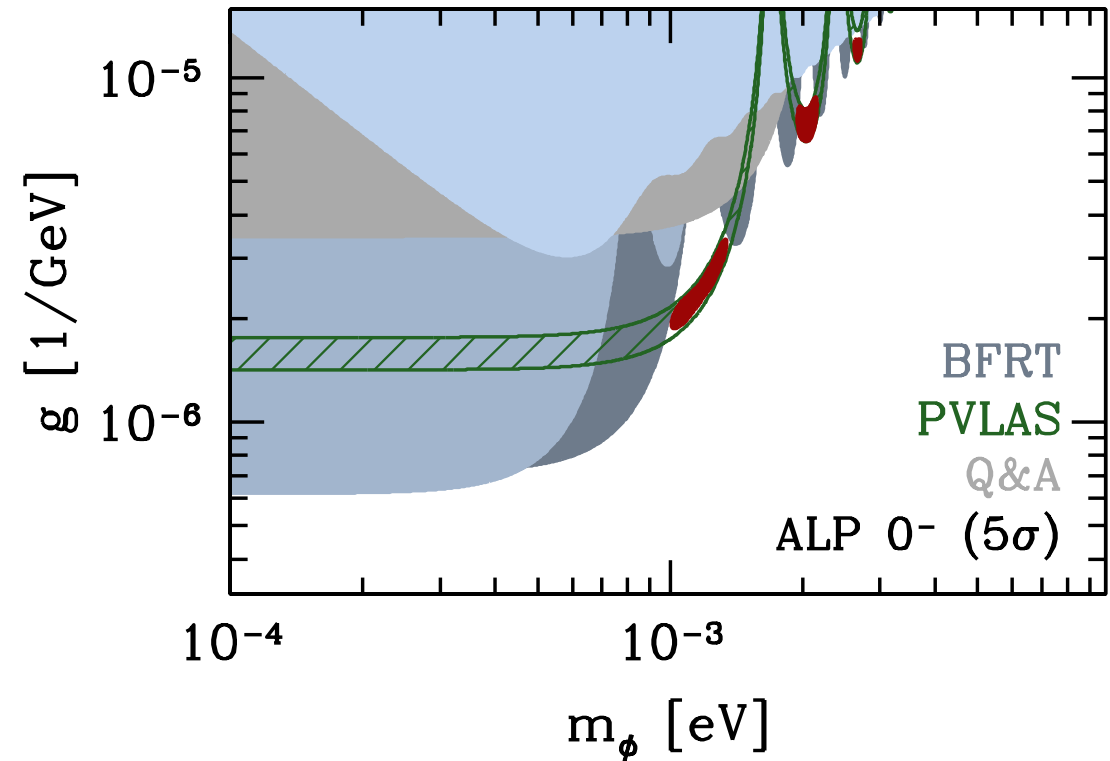
[...;Toll '52;...;Adler '71;...;Tsai,Erber '74,'75;...]

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If interpreted in terms of ALP:



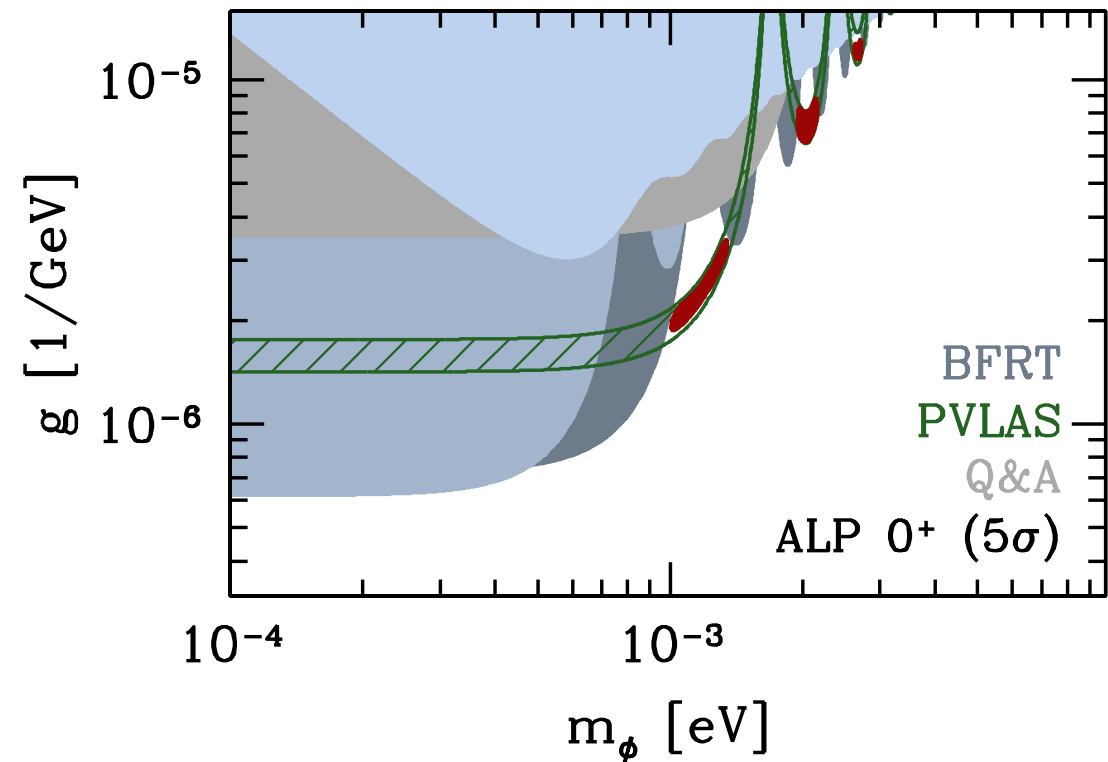
[Ahlers, Gies, Jaeckel, AR '06]

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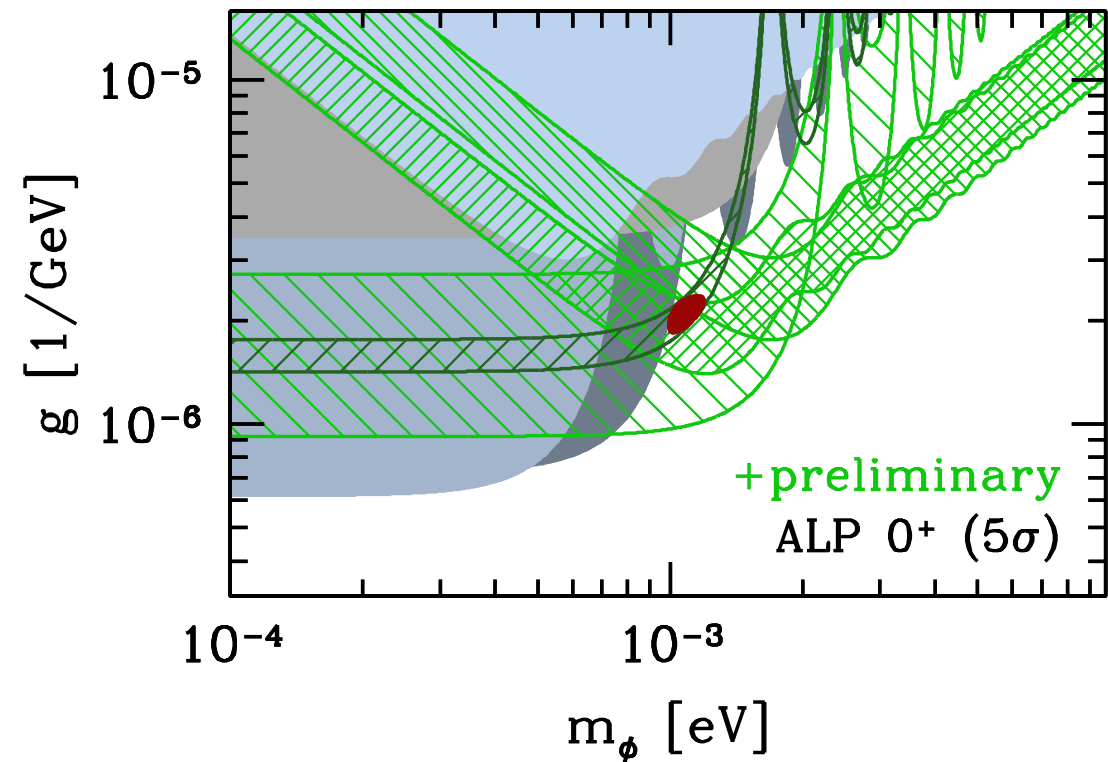
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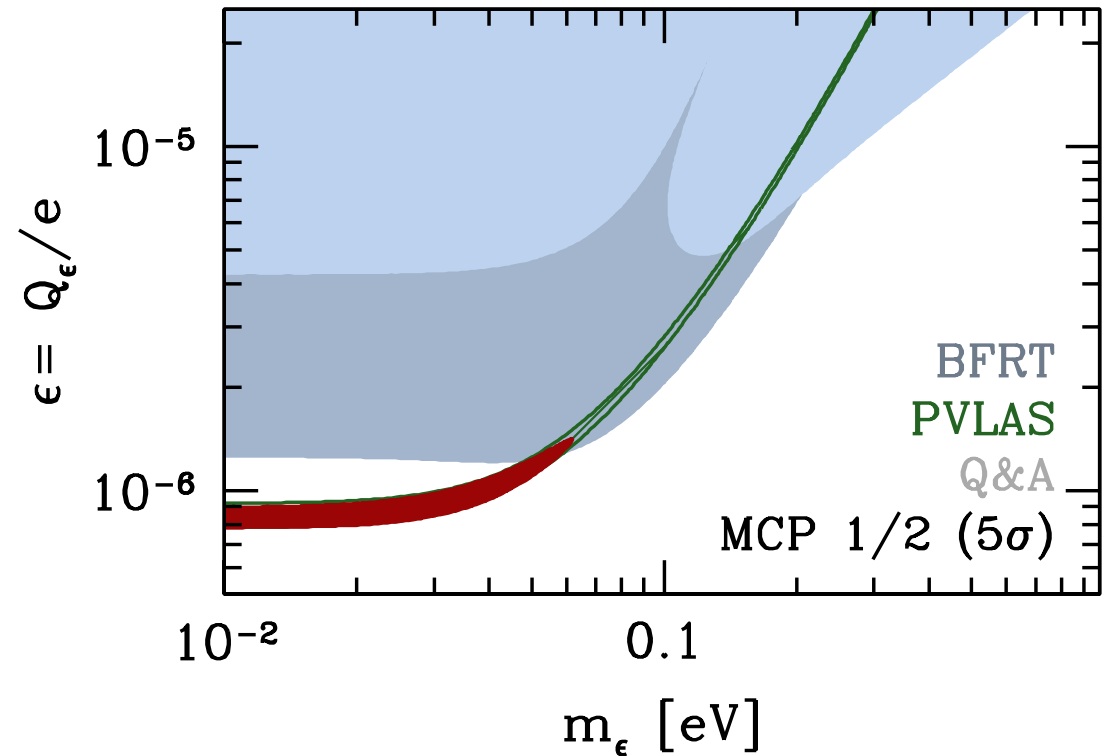
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If interpreted in terms of MCP:



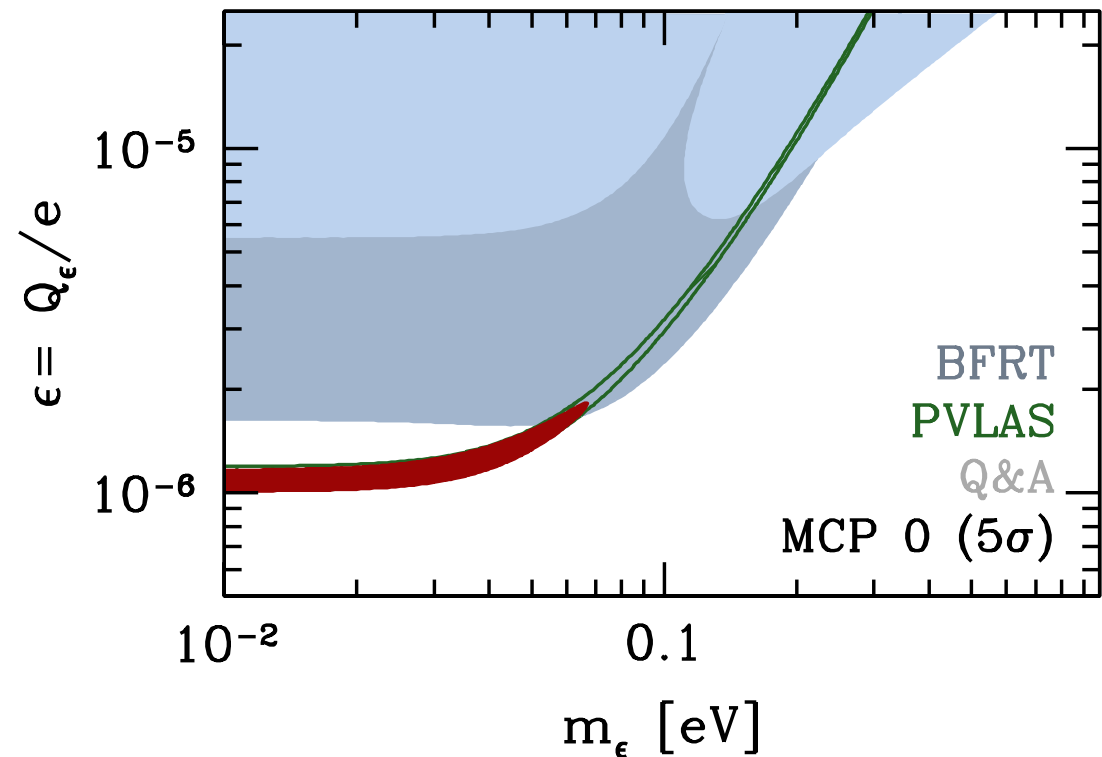
[Ahlers, Gies, Jaeckel, AR '06]

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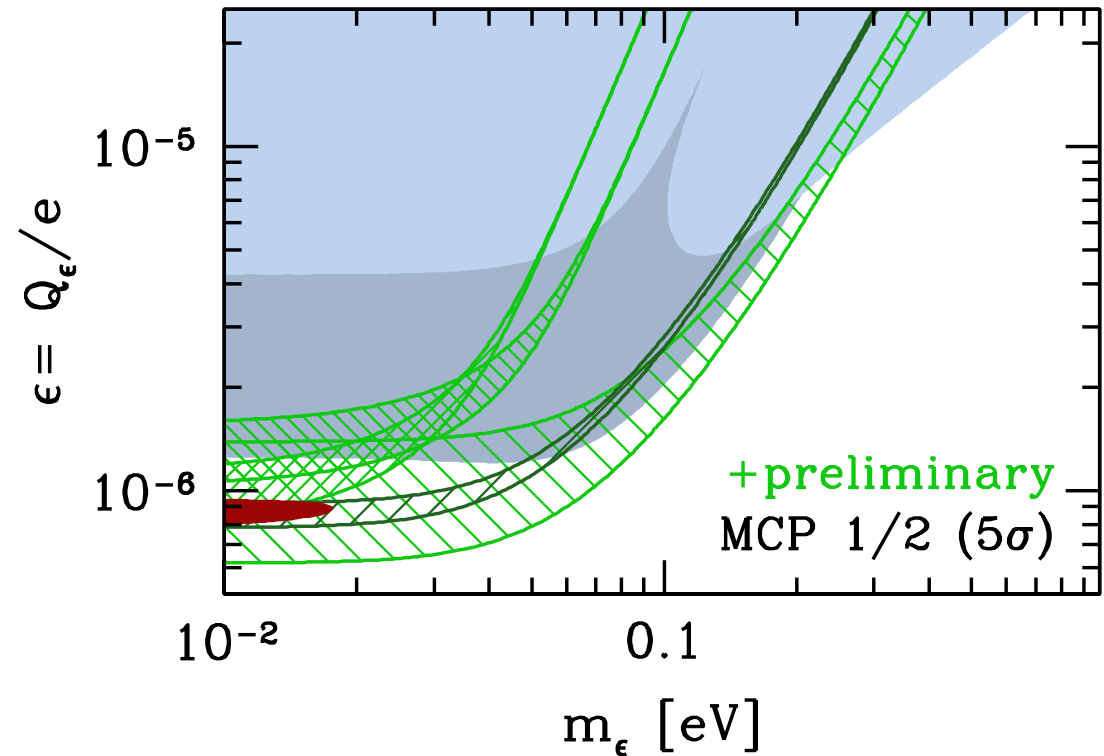
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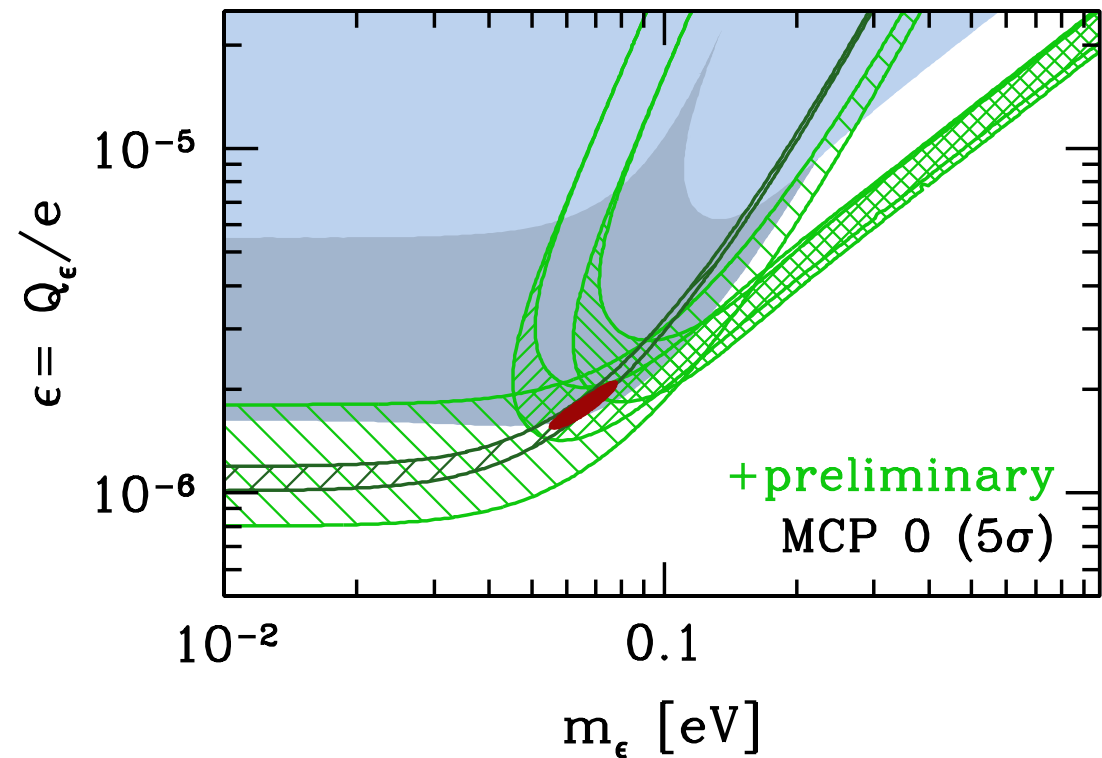
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If interpreted in terms of MCP:



[Ahlers, Gies, Jaeckel, AR '06]

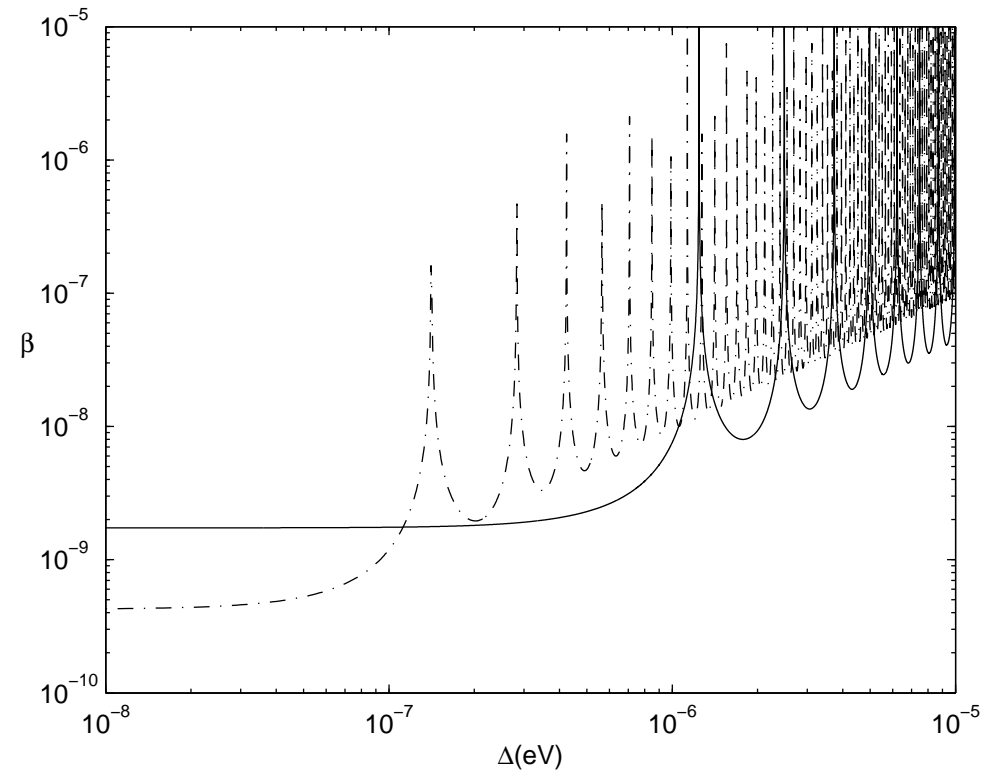
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$$\partial_\mu \rightarrow \partial_\mu - i\epsilon e A_\mu$$

- Explanation in terms of real and virtual production of Second Photon (SP) in fusion model of photon: [de Broglie '32]
 - Photon: $S = 1$ bound state of spin 1/2 particle-antiparticle pair
 - SP: $S = 0$ bound stateParameters: Mass splitting Δ and magnetic moment $\beta\mu_B$ of particle

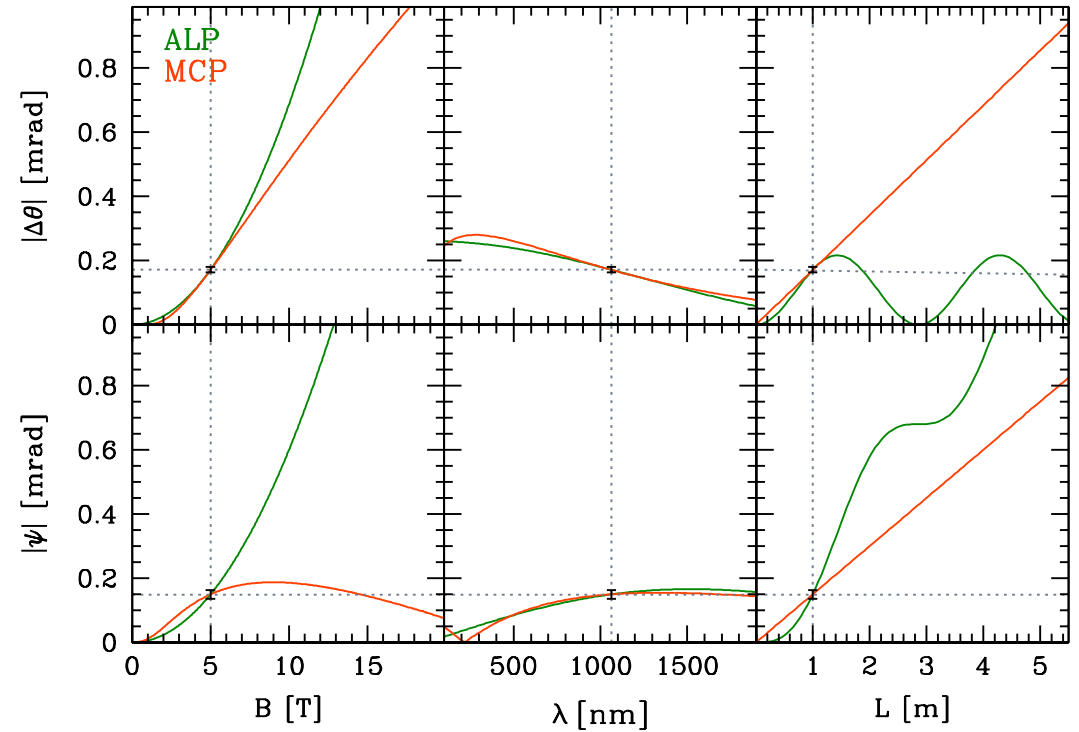
If interpreted in terms of SP:



[Beswick, Rizzo '07]

3. Crucial Laboratory Tests

- Laser polarization experiments at higher magnetic fields



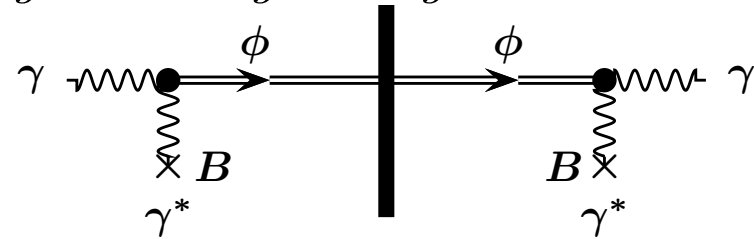
[Ahlers, Gies, Jaeckel, AR '06]

BMV (Toulouse): 11 T pulsed magnet

3. Crucial Laboratory Tests

- Laser polarization experiments at higher magnetic fields
- Light shining through a wall

“Light shining through a wall”



[Sikivie '83; Ansel'm '85; Van Bibber *et al.* '87]

Name	Laboratory	Magnets	$P_{\gamma\phi\gamma} g_{\text{PVLAS}}$
ALPS	DESY/D	$B_1 = B_2 = 5 \text{ T}$ $l_1 = l_2 = 4.21 \text{ m}$	$\sim 10^{-19}$
BMV	LULI/F	$B_1 = B_2 = 11 \text{ T}$ $l_1 = l_2 = 0.25 \text{ m}$	$\sim 10^{-21}$
LIPSS	Jlab/USA	$B_1 = B_2 = 1.7 \text{ T}$ $l_1 = l_2 = 1 \text{ m}$	$\sim 10^{-23.5}$
OSQAR	CERN/CH	$B_1 = B_2 = 11 \text{ T}$ $l_1 = l_2 = 7 \text{ m}$	$\sim 10^{-17}$
PVLAS	Legnaro/I	$B_1 = 5 \text{ T}$ $l_1 = 1 \text{ m}$ $B_2 = 2.2 \text{ T}$ $l_2 = 0.5 \text{ m}$	$\sim 10^{-23}$

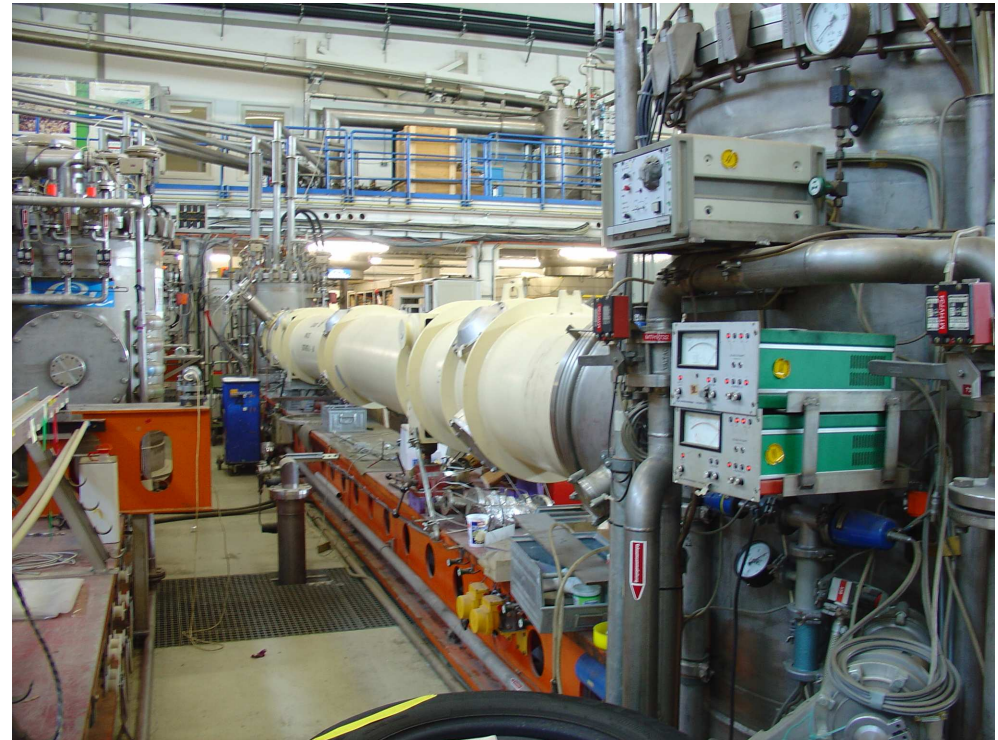
[AR '06]

3. Crucial Laboratory Tests

- Laser polarization experiments at higher magnetic fields
- Light shining through a wall

Axion-Like Particle Search:

[DESY, Laser Zentrum Hannover, Sternwarte Bergedorf]



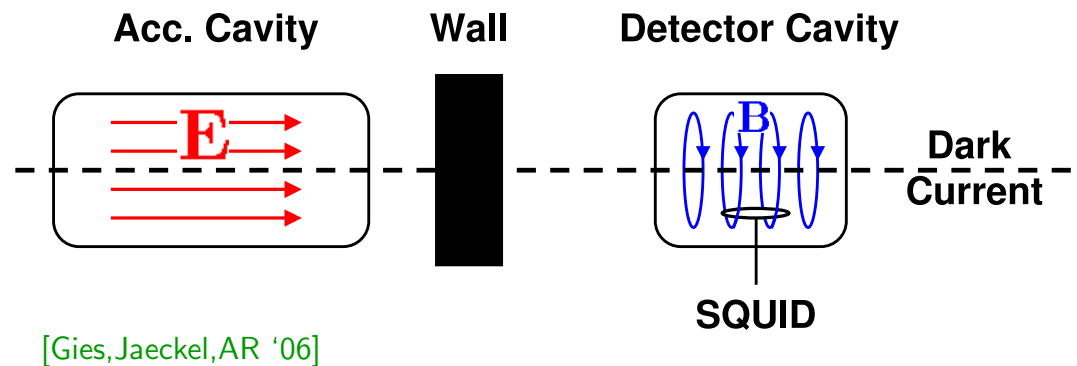
$$B = 5 \text{ T}, \ell = 4.2 \text{ m}, \underbrace{\langle P \rangle = 0.2 \text{ kW}, \omega = 1.2 \text{ eV}}_{\dot{N}_0 \sim 1 \times 10^{21} / \text{s}}, N_r = 0$$

Test of ALP interpretation of PVLAS in
summer 2007

Venice, March 2007

3. Crucial Laboratory Tests

- Laser polarization experiments at higher magnetic fields
- Light shining through a wall
- Dark current through a wall

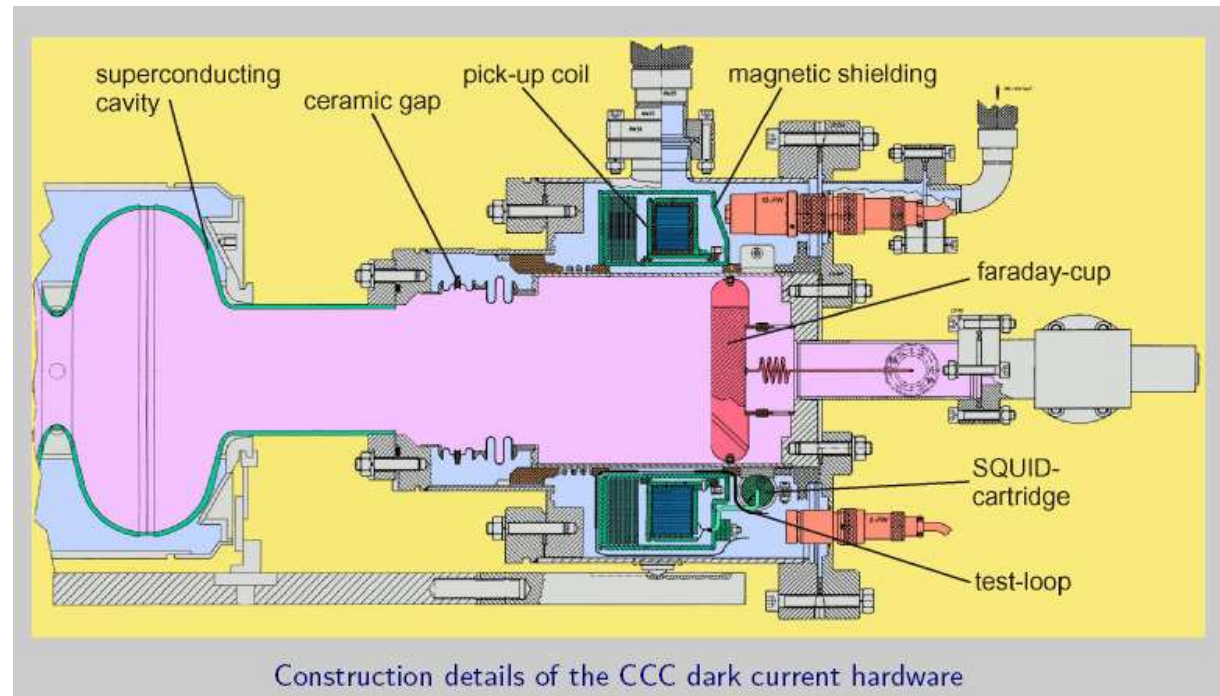


3. Crucial Laboratory Tests

- Laser polarization experiments at higher magnetic fields
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Cryogenic Current Comparator:

[DESY, GSI, Universität Jena]

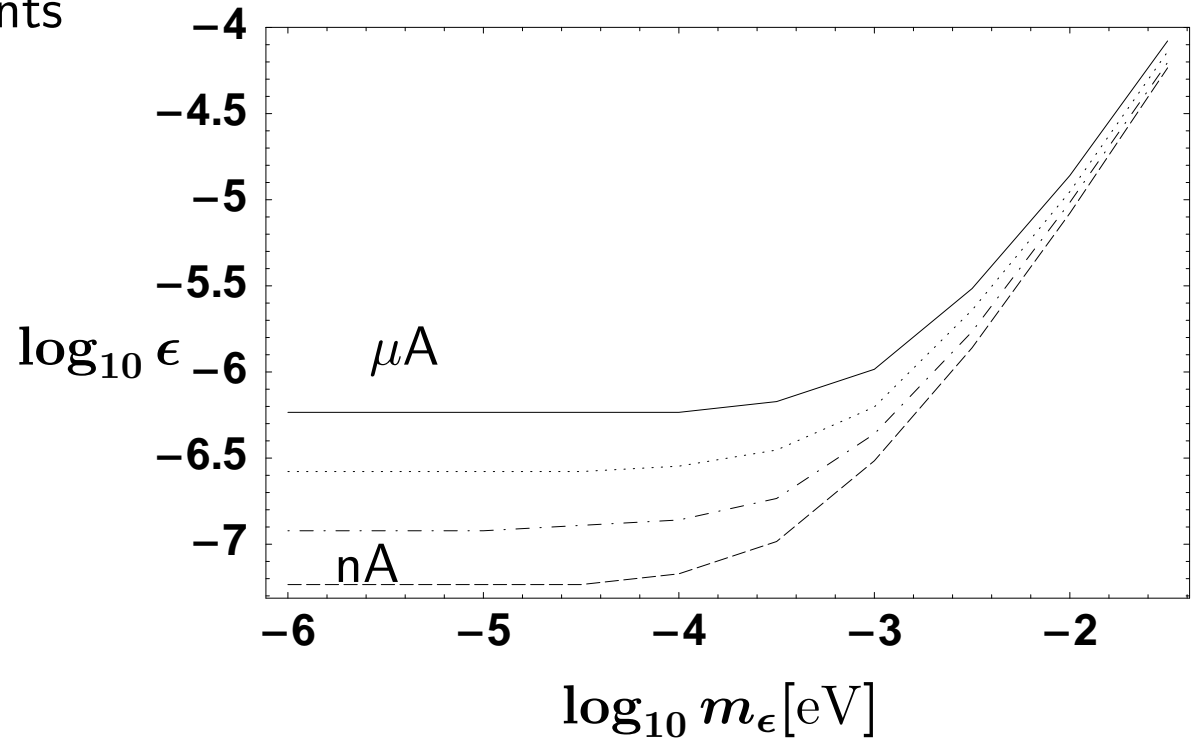


[M. Wendt TESLA2004]

3. Crucial Laboratory Tests

- Laser polarization experiments at higher magnetic fields
- Light shining through a wall
- Dark current through a wall

TESLA accelerator cavity | CCC:



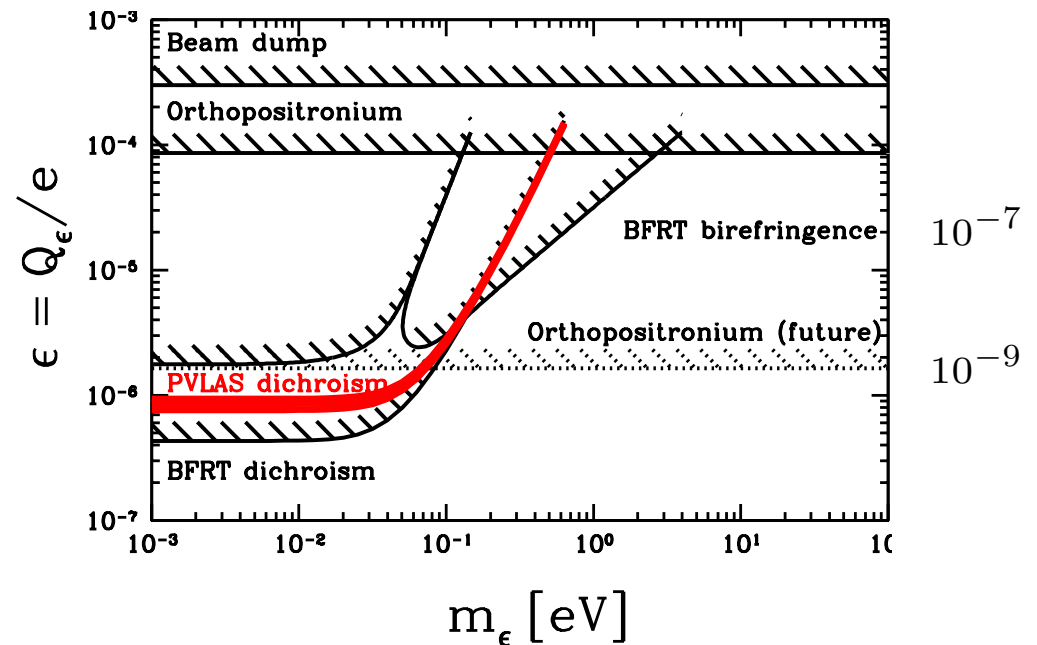
[Gies, Jaeckel, AR unpubl.]

3. Crucial Laboratory Tests

- Laser polarization experiments at higher magnetic fields
- Light shining through a wall
- Dark current through a wall
- Invisible Orthopositronium decay

“Search for Invisible Orthopositronium Decay” [Dobroliubov, Ignatiev '89]

$$\text{BR}(\text{OP} \rightarrow \epsilon^+ \epsilon^-) \simeq \frac{3\pi\epsilon^2}{4\alpha(\pi^2 - 9)} \simeq 371 \epsilon^2$$



[Mitsui *et al.* '93]: $\text{BR}(\text{OP} \rightarrow \text{inv.}) < 2.8 \times 10^{-6}$

[Badertscher *et al.* '06]: $\text{BR}(\text{OP} \rightarrow \text{inv.}) < 4.2 \times 10^{-7}$

3. Crucial Laboratory Tests

- Laser polarization experiments at higher magnetic fields
- Light shining through a wall
- Dark current through a wall
- Invisible Orthopositronium decay
- Searches for excess e^- from elastic $\epsilon^\pm e^-$ scattering in detector near nuclear reactor

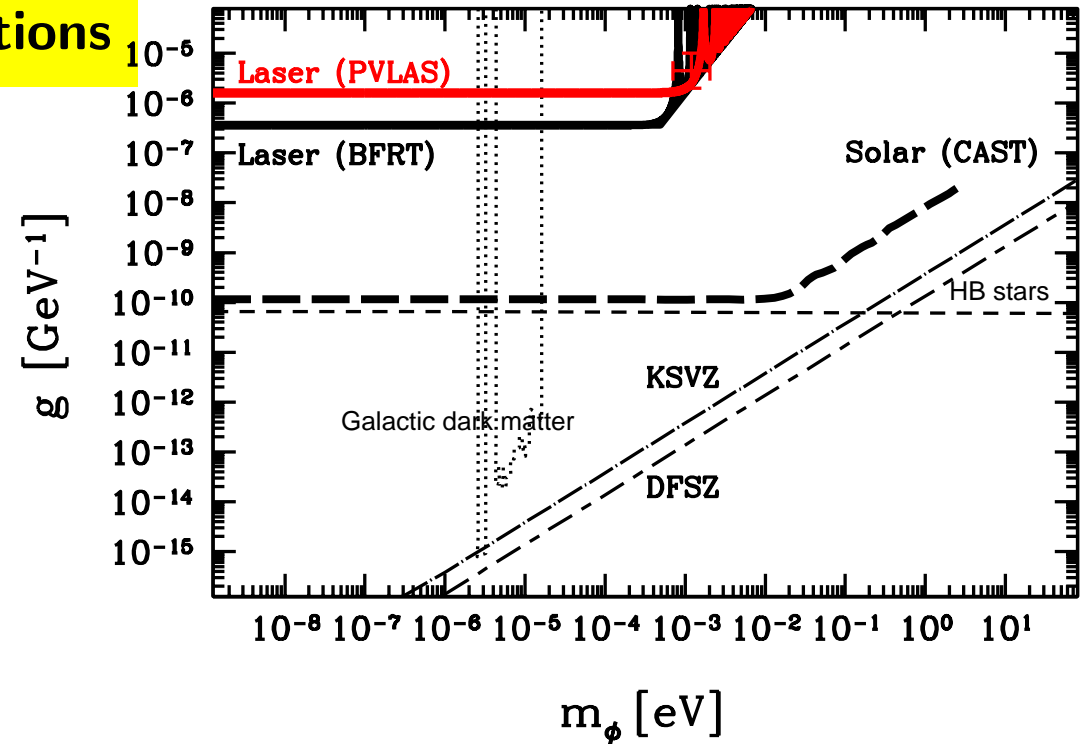
- Nuclear power reactors with $P > 2 \text{ GW}^{24}$ emit more than $10^{20} \gamma/\text{s}$
 - These γ s may convert within reactor into $\epsilon^+ \epsilon^-$ pairs
 - A small fraction of these particles could lead to an observable excess of electrons from $\epsilon^\pm e^-$ scattering in a detector
 - Recent results from the **TEXONO** experiment set up at the Kuo-Sheng Nuclear Power Station (2.8 GW) in Taiwan probing for $\mu_{\bar{\nu}_e}$ by searching for an excess of events from νe^- magnetic scattering [TEXONO Coll. '03]
- ⇒ Bound on fractional electric charge,

$$\epsilon \lesssim 10^{-5}, \text{ for } m_\epsilon \lesssim 1 \text{ keV}$$

- May be improved in near future with massive liquid argon detector

4. Problems of Particle Interpretations

- Energy loss of stars:
 - ALPs: Primakoff $\gamma Z \rightarrow \phi Z$



Way out: $g|_{\text{star}} \ll g|_{\text{vacuum}}$

[Masso, Redondo '05; Jaeckel *et al.* '06]

⇒ Even more sub-eV particles and fields,

e.g. [Masso, Redondo '06; Mohapatra, Nasri '06]

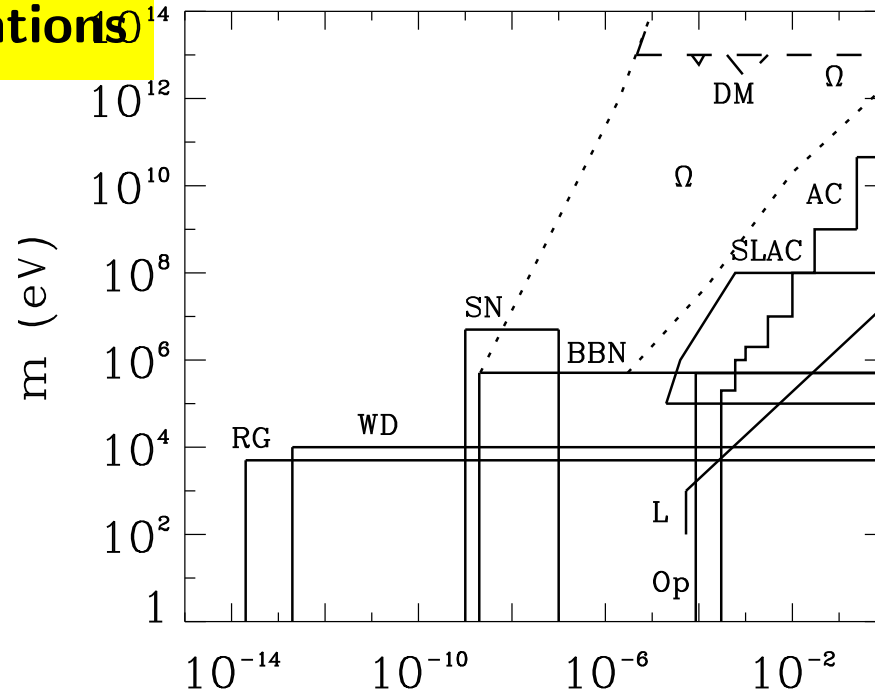
– U(1) bosons mixing with photon

– scalar field with low scale phase transition

Venice, March 2007

4. Problems of Particle Interpretations

- Energy loss of stars:
 - ALPs: Primakoff $\gamma Z \rightarrow \phi Z$
 - MCPs: plasmon decay $\gamma^* \rightarrow \epsilon^+ \epsilon^-$



[Davidson, Hannestad, Raffelt '00] ϵ

Way out: $\epsilon_{\text{plasma}} \ll \epsilon_{\text{vacuum}}$

[Masso, Redondo '06]

\Rightarrow Even more sub-eV particles and fields,
e.g. U(1) bosons mixing with photon

[Abel *et al.* '06; Foot, Kobakhidze '07]

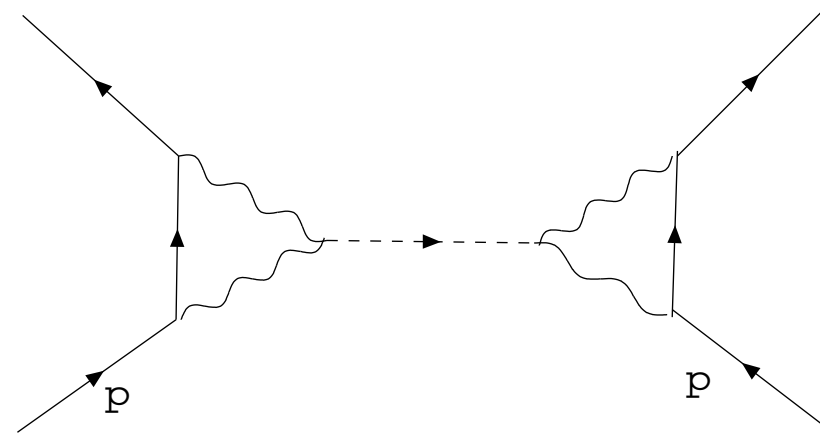
4. Problems of Particle Interpretations

- Energy loss of stars:
 - ALPs: Primakoff $\gamma Z \rightarrow \phi Z$
 - MCPs: plasmon decay $\gamma^* \rightarrow \epsilon^+ \epsilon^-$
 - SPs: no problem
- ALP 0^+ : Non-Newtonian force,

$$V(r) = G \frac{m_1 m_2}{r} + \frac{y^2}{4\pi} \frac{n_1 n_2}{r} e^{-m_\phi r}$$

from Yukawa coupling

$$\begin{aligned} \mathcal{L}_{\phi pp} &= y \phi \bar{\Psi}_p \Psi_p \\ y &\simeq \frac{3\alpha}{2\pi} (g m_p) \log \frac{\Lambda}{m_p} \end{aligned}$$



[Adelberger et al. '06]

From torsion-balance experiment:

$$g < 4 \times 10^{-17} \text{ GeV}^{-1},$$

for $m_\phi = 1 \text{ meV}$; $\Lambda \gg m_p$

[Dupays et al. '06; Adelberger et al. '06]

Way out: ALP 0^+ couples only to additional light U(1) bosons mixing with photon

5. WILPs in Models with Light Extra-U(1)s

28

- Consider extension of SM with additional “hidden sector” U(1)’s \Rightarrow in general gauge kinetic mixing with “visible” U(1), e.g.

$$\mathcal{L} = -\frac{1}{4} F^T \mathcal{K}_F F + \frac{1}{2} A^T \mathcal{M}_A^2 A + e_j A,$$

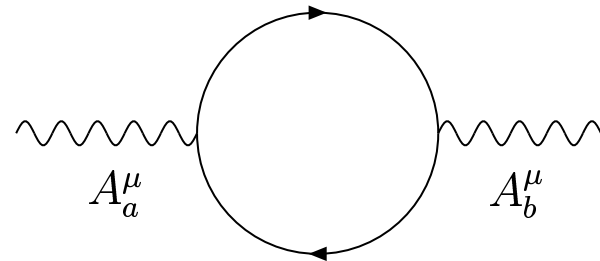
with special mixing

$$\mathcal{K}_F = \begin{pmatrix} 1 & \chi & \chi \\ \chi & 1 & 0 \\ \chi & 0 & 1 \end{pmatrix}$$

Mixing may arise from integrating out heavy particles:

e.g. threshold effect from two species of fermions, (e_a, e_b) and $(e_a, -e_b)$, with masses m and m' :

$$\chi \simeq \frac{e_a e_b}{6\pi^2} \log \left(\frac{m'}{m} \right)$$



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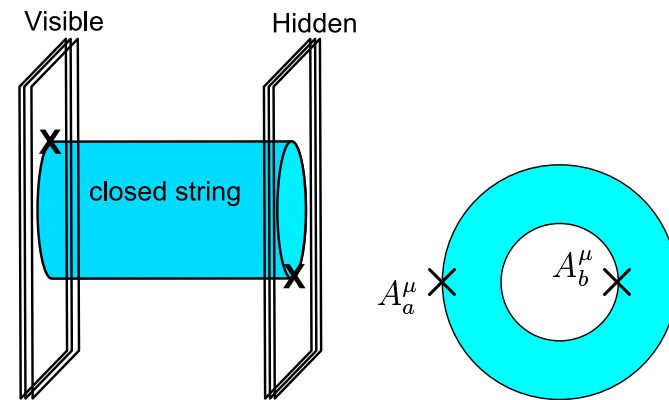
with special mixing

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Mixing may arise from integrating out heavy particles:

e.g. closed string exchange between visible sector D-branes and hidden sector anti-D-branes,

$$\chi \sim g_a g_b \left(\frac{2^{(8-p)/2} M_s}{\alpha_p M_P} \right)^{\frac{2(5-p)}{6-p}} \left(\frac{R}{r} \right)^{\frac{d-p+3}{6-p}}$$



[Abel, Jaeckel, Khoze, AR '06]

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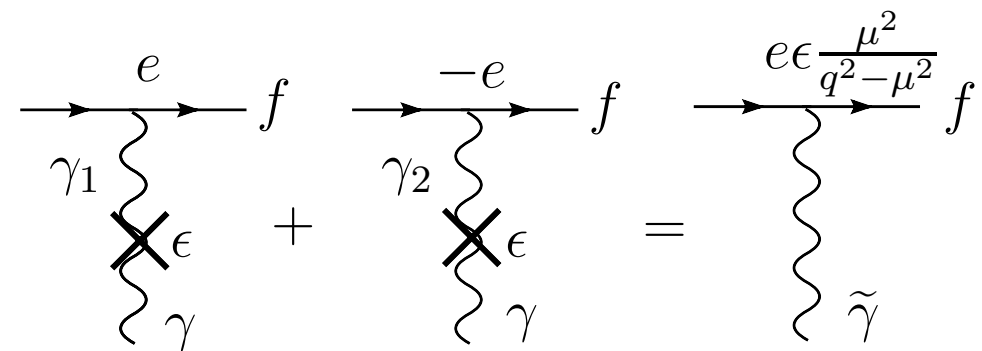
with special mixing and mass pattern

$$\mathcal{K}_F = \begin{pmatrix} 1 & \chi & \chi \\ \chi & 1 & 0 \\ \chi & 0 & 1 \end{pmatrix}, \mathcal{M}_A^2 = \begin{pmatrix} m_\gamma^2 & 0 & 0 \\ 0 & \mu^2 & 0 \\ 0 & 0 & 0 \end{pmatrix}$$

- Hidden sector CP, with special charge assignment (0, e, -e), acquires visible mini-charge

$$\epsilon \simeq \begin{cases} -\chi & \text{for } m_\gamma = 0 \\ (\mu^2/\omega_p^2) \chi & \text{for } m_\gamma = \omega_p \gg \mu \end{cases}$$

\Rightarrow Viable MCP explanation of PVLAS ($m_\gamma = 0$)
for $\chi \sim 10^{-6}$, $m_\epsilon \lesssim 0.1$ eV;



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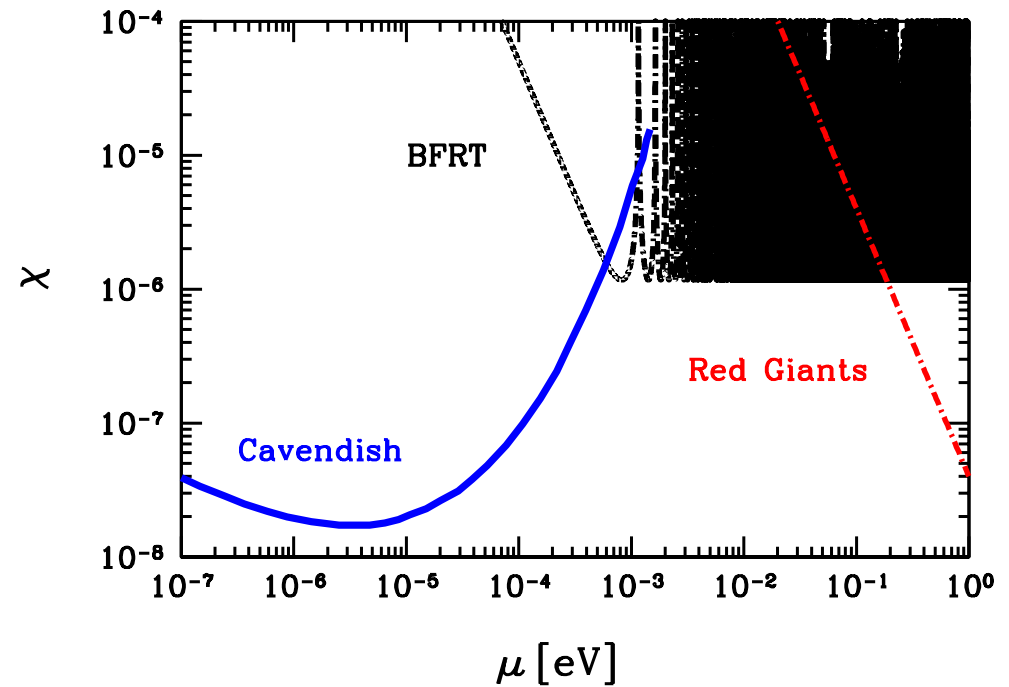
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$$\epsilon \simeq \begin{cases} -\chi & \text{for } m_\gamma = 0 \\ (\mu^2/\omega_p^2) \chi & \text{for } m_\gamma = \omega_p \gg \mu \end{cases}$$

\Rightarrow Viable MCP explanation of PVLAS ($m_\gamma = 0$) for $\chi \sim 10^{-6}$, $m_\epsilon \lesssim 0.1$ eV; in accordance with the life time of stars ($m_\gamma = \omega_p \sim \text{keV}$), if $\mu \lesssim 0.1$ eV [Abel et al. '06]



[AR unpubl]

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with special mixing and mass pattern

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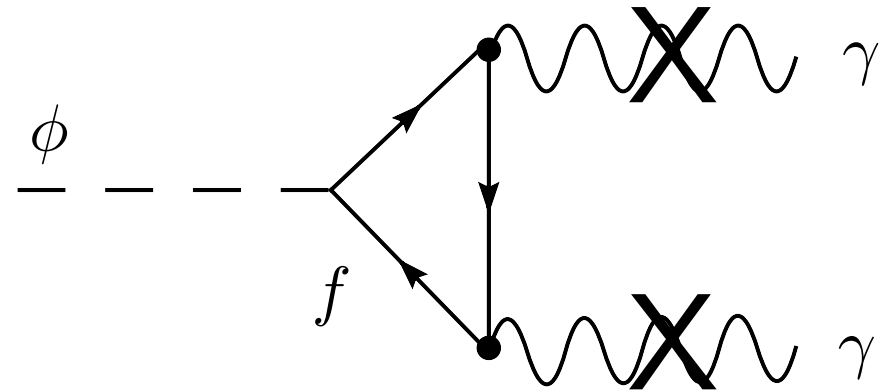
- Hidden sector CP, with special charge assignment (0, e, -e), acquires visible mini-charge

$$\epsilon \simeq \begin{cases} -\chi & \text{for } m_\gamma = 0 \\ (\mu^2/\omega_p^2) \chi & \text{for } m_\gamma = \omega_p \gg \mu \end{cases}$$

\Rightarrow Viable MCP explanation of PVLAS ($m_\gamma = 0$) for $\chi \sim 10^{-6}$, $m_\epsilon \lesssim 0.1$ eV; in accordance with the life time of stars ($m_\gamma = \omega_p \sim \text{keV}$), if $\mu \lesssim 0.1$ eV [Abel et al. '06]

\Rightarrow Extend minimal model by hidden sector scalar, $m_\phi \sim \text{meV}$, coupled to hidden sector CPs \Rightarrow

Viable ALP explanation [Masso, Redondo '06]
A. Ringwald (DESY)



- Coupling to two photons can be arranged to be in PVLAS range,

$$g \sim \frac{\alpha}{2\pi} \chi^2 \frac{y_f}{m_f} \sim 2 \times 10^{-6} \text{ GeV}^{-1} \left(\frac{\chi}{10^{-6}} \right)^2 \left(\frac{y_f \text{ eV}}{m_f} \right)$$

- Yukawa coupling to proton,

$$y \sim \frac{\alpha}{\pi} \chi^2 \frac{\mu}{m_f},$$

suppressed by factor μ/m_p compared to case without kinetic mixing \Rightarrow no problem with non-Newtonian forces, if $\mu \lesssim \text{meV}$

Venice, March 2007

6. Summary

- The evidence for a vacuum magnetic dichroism and birefringence by **PVLAS** has triggered a lot of theoretical and experimental activities:
 - Particle interpretations alternative to **ALP** interpretation: e.g. **MCP**
 - Models, which evade strong astrophysical and cosmological bounds on such particles, have been found. Require typically even more **WILPs** than just the ones introduced for the solution of the **PVLAS** puzzle
 - Decisive laboratory based tests of particle interpretation of **PVLAS** anomaly in very near future. More generally, experiments will dig into previously unconstrained parameter space of above mentioned models
- Experiments exploiting low energy photons may give information about fundamental particle physics complementary to the one obtained at high energy colliders