

SNO RESULTS, MAJORANA NEUTRINOS, CP-VIOLATION, $(\beta\beta)_{0\nu}$ -DECAY

S. PASCOLI at SISSA (Trieste, Italy)
(and INFN (Trieste - Italy)).

NO-VE: NEUTRINO OSCILLATIONS IN VENICE

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S. BILENKY, S. PASCOLI and S.T. PETCOV [hep-ph/0102265
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to appear in PRD]

- $(\beta\beta)_{0\nu}$ -decay and the effective Majorana mass
- The 3- ν mixing case: i) predictions of $|m_3|$ in the hierarchical, inverted hierarchy and mass degeneracy spectra
ii) information on m_1 from $|m_3|$
iii) information on CPV from $|m_3|$
- The 4- ν mixing case: predictions of $|m_3|$ in the 2+2A, 2+2B, 3+1A, 3+1B, 3+1C neutrino mass spectra.
- Conclusions

References:

- S. T. Petcov and A. Yu. Smirnov, Phys. Lett. B **322**, 109 (1994).
- S.M. Bilenky A. Bottino, C. Giunti and C.W. Kim, Phys. Lett. B **356**, 273 (1995) and Phys. Rev. D **54**, 1881 (1996).
- S.M. Bilenky, C. Giunti, C.W. Kim and S.T. Petcov, Phys. Rev. D **54**, 4432 (1996).
- S.M. Bilenky, C. Giunti, W. Grimus, B. Kayser and S.T. Petcov, Phys. Lett. B**465**, 193 (1999); S.T. Petcov, in *Weak Interactions and Neutrinos*, Proceedings of the 17th International Conference, January 23 - 30, 1999, Cape Town, South Africa (eds. C.A. Dominguez and R.D. Viollier, World Scientific, Singapore, 2000), p. 305; S.M. Bilenky and C. Giunti in *Weak Interactions and Neutrinos*, Proceedings of the 17th International Conference, January 23 - 30, 1999, Cape Town, South Africa (eds. C.A. Dominguez and R.D. Viollier, World Scientific, Singapore, 2000), p. 195.
- C. Giunti Phys. Rev. D**61**, 036002 (2000); T. Fukuyama, K. Matsuda, H. Nishiura and N. Takeda Phys. Rev. D**62**, 93001 (2000) and hep-ph/0007237; M. Czakon, J. Gluza, J. Studnik and M. Zlarek hep-ph/0010077; H. V. Klapdor-Kleingrothaus, H. Pas and A. Yu. Smirnov hep-ph/0003219; S.M. Bilenky, S. P. and S.T. Petcov, Ref. SISSA 13/2001/EP, hep-ph/0104218, hep-ph/0102265.

- Strong indications of NEUTRINO OSCILLATIONS
(from atmospheric, solar- ν data).



If confirmed, they would imply the EXISTENCE OF

NEUTRINO MIXING
and
MASSIVE NEUTRINOS.

- Then the fundamental question to address would be the NATURE of NEUTRINOS:

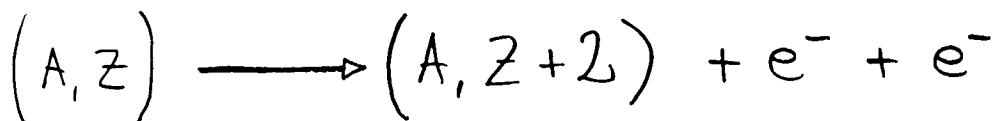
DIRAC PARTICLES
in theories with
L conserved



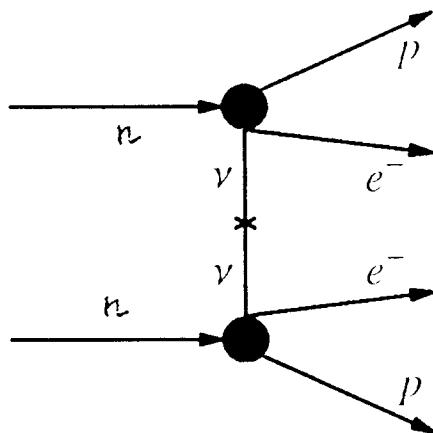
MAJORANA PARTICLES
in theories with
L violated

- The question of massive neutrinos is one of the fundamental questions related to the problem of ν -mixing.
- The nature of massive neutrinos is related to the symmetries of elementary particle interactions.
- The Majorana nature of neutrinos can manifest itself in $\Delta L=2$ processes. Now the most sensitive one is $(\beta\beta)_{0\nu}$ - decay.

NEUTRINOLESS DOUBLE β -DECAY



$\Delta L = 2$
Majorana
neutrinos



The probability amplitude is sensitive to the
EFFECTIVE MAJORANA MASS PARAMETER $|\langle m \rangle|$:

$$|\langle m \rangle| = \left| |U_{e1}|^2 m_1 + |U_{e2}|^2 m_2 e^{i\alpha_{21}} + |U_{e3}|^2 m_3 e^{i\alpha_{31}} \right|$$

in the hypothesis of $3 \rightarrow 2$ mixing.

- Experimentally $(\beta\beta)_{0\nu}$ not yet detected:

$$|\langle m \rangle| \leq (0.35 \div 1.05) \text{ eV} \quad (\text{*Ge Heidelberg-Moscow})$$

$$|\langle m \rangle| \leq (0.33 \div 1.35) \text{ eV} \quad (\text{IGEX Collaboration})$$

- Higher sensitivity in the upcoming experiments:

$$|\langle m \rangle| \approx 0.1 \text{ eV}$$

$$|\langle m \rangle| \approx 0.01 \text{ eV}$$

RELEVANT PARAMETERS IN $|km|$

(3- ν mixing case)

- Oscillation parameters: i) Δm_0^2 , Δm_{atm}^2
ii) θ_0 , δ_{CP}

They can be measured with good accuracy.
 θ_{atm} and δ do not play any role.

! Crucial: to determine the correct ν_0 -anomaly solution:

SNO	LMA	$\tan^2 \theta_0 \approx 0.2 \div 1.0$	$\Delta m_0^2 \approx 10^{-5} \div 10^{-4} \text{ eV}^2$
	SMA	$\tan^2 \theta_0 \approx 10^{-4} \div 10^{-3}$	$\Delta m_0^2 \approx 10^{-9} \div 10^{-7} \text{ eV}^2$
SNO	LOW	$\tan^2 \theta_0 \approx 0.4 \div 3.0$	$\Delta m_0^2 \approx 10^{-6} \div 10^{-5} \text{ eV}^2$

- Non-oscillation parameters: i) m_1 , mass of the highest neutrino

ii) α_{21}, α_{31}

Majorana CPV phases

If CP is conserved:

$$\alpha_{21}, \alpha_{31} = 0, \pm \pi.$$

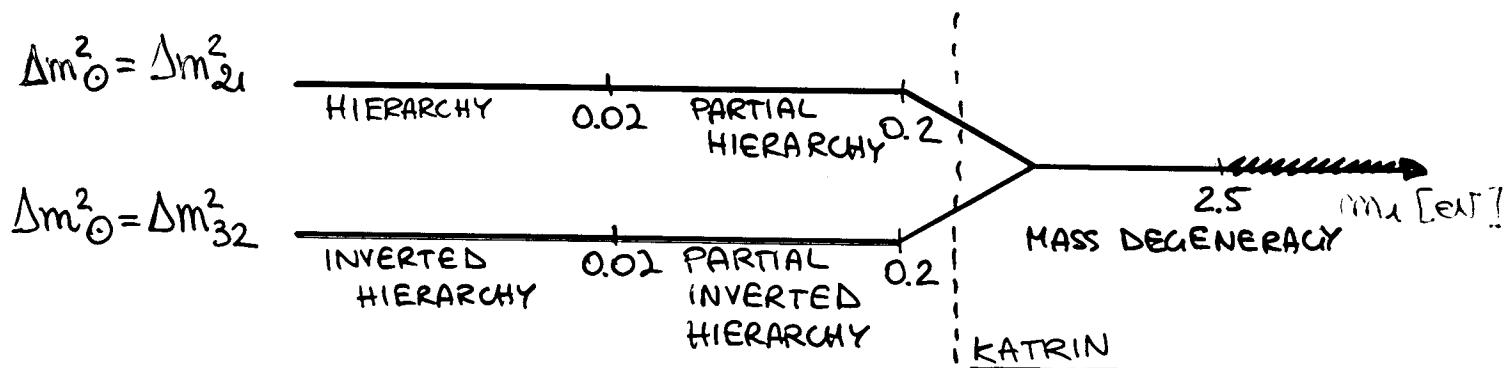
These parameters are not constrained so far.
 $(m_1 \leq 2.2 \div 2.5 \text{ eV from } {}^3\text{H}-\beta \text{ decay}).$

- Lightest ν mass m_1

$$m_{\nu_1}, \quad m_2 = \sqrt{m_1^2 + \Delta m_{21}^2}, \quad m_3 = \sqrt{m_1^2 + \Delta m_{31}^2}$$

All the 3 ν -masses are determined by m_1 and two mass-squared differences which have to be identified with Δm_{31}^2 , Δm_{21}^2 :

$$\Delta m_{\text{atm}}^2 \equiv \Delta m_{31}^2, \quad \Delta m_{\odot}^2 = \Delta m_{21}^2, \Delta m_{32}^2$$



! Crucial: to determine the correct mass spectrum fixing the value of m_1 and resolving the ambiguity $\Delta m_{\odot}^2 = \Delta m_{21}^2, \Delta m_{32}^2$.

- CPV phases α_{21}, α_{31}

With mixing of 3 massive Majorana neutrinos, 3 CPV phases are relevant: $S, \alpha_{21}, \alpha_{31}$.

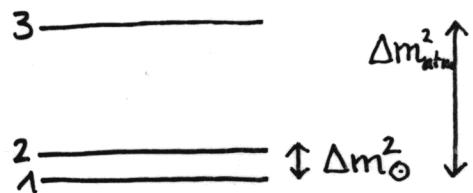
$\tilde{\alpha}$ universal CPV phase. It does not enter in $(\beta\beta)_{\text{ov}}$ -decay.

α_{21}, α_{31} Majorana CPV phases, they are present only if ν are Majorana particles. They enter in $(\beta\beta)_{\text{ov}}$ -decay and other $\Delta L=2$ processes. CP conservation implies $\alpha_{21}, \alpha_{31} = 0, \pm \pi$.

HIERARCHICAL MASS SPECTRUM

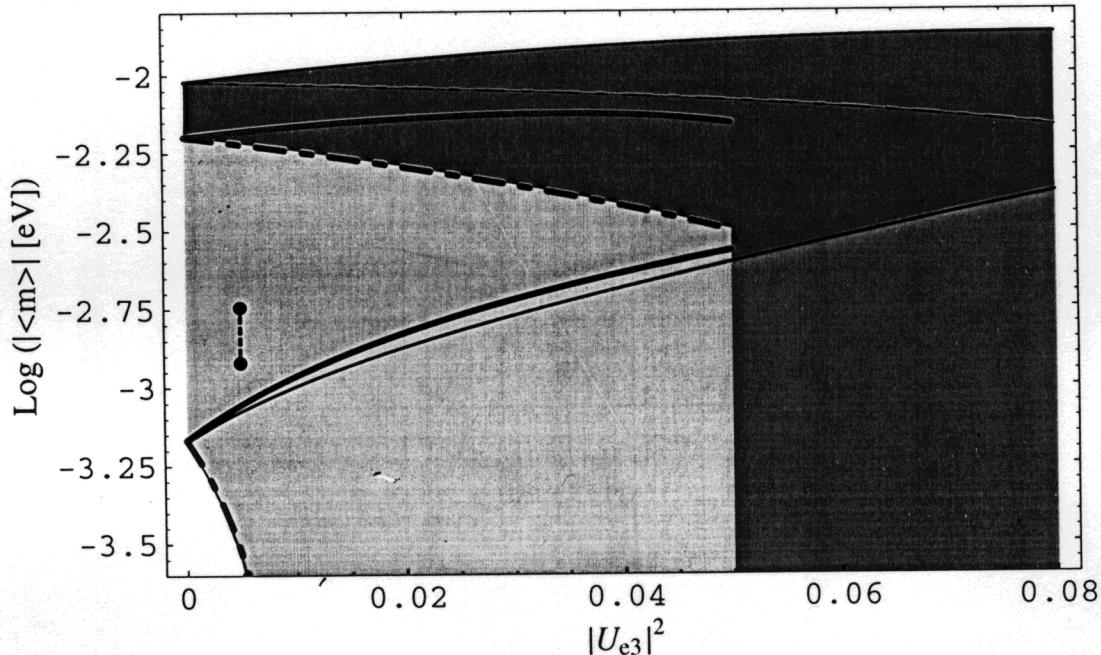
characterized by: $m_1 \ll m_2 \ll m_3$

$$\Delta m_{\odot}^2 = \Delta m_{21}^2$$



$$|\langle m \rangle| = \left| \sqrt{\Delta m_{\odot}^2} (1 - |U_{e3}|^2) \sin^2 \theta_{\odot} + \sqrt{\Delta m_{\text{atm}}^2} |U_{e3}|^2 e^{i\alpha_{32}} \right|$$

$\log(|\langle m \rangle| [\text{eV}])$ vs $|U_{e3}|^2$ in the LMA case:



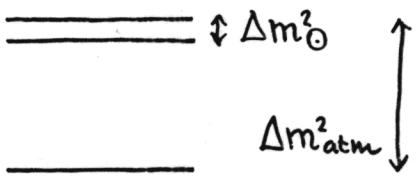
$$(0 \div 0) \text{ eV} \leq |\langle m \rangle| \leq (7.6 \div 14) \times 10^{-3} \text{ eV}$$

- $|\langle m \rangle|$ can provide information on α_{32}
- however, being $|\langle m \rangle| \lesssim 0.01 \text{ eV}$, $|\langle m \rangle|$ is beyond the sensitivity of the upcoming $(\beta\beta)_{0\nu}$ -decay experiments.

INVERTED HIERARCHY SPECTRUM

characterized by: $m_1 \ll m_2 \simeq m_3$

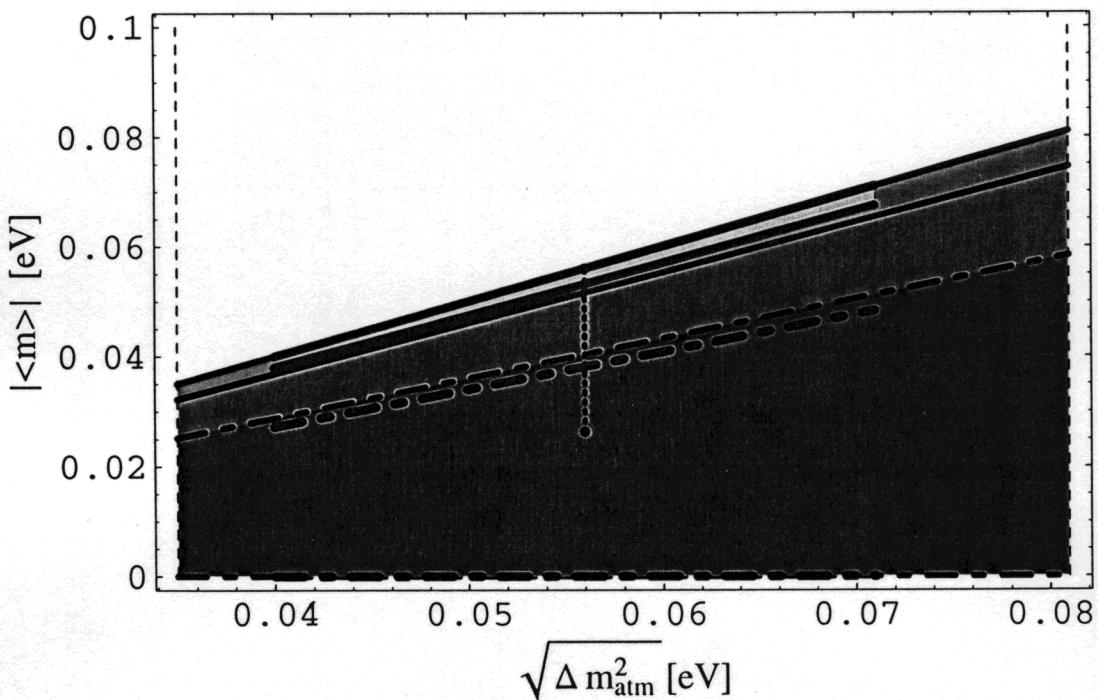
$$\Delta m_0^2 = \Delta m_{32}^2$$



$$|\langle m \rangle| = (1 - |U_{e1}|^2) \sqrt{1 - \sin^2 2\theta_0 \sin^2 \frac{\alpha_{32}}{2}} \sqrt{\Delta m_{\text{atm}}^2}$$

$|\langle m \rangle|$ can provide information on α_{32} .

$|\langle m \rangle|$ vs $\sqrt{\Delta m_{\text{atm}}^2}$ for the LMA MSW case:



$$(7.3 \times 10^{-3} \div 0) \text{ eV} \leq |\langle m \rangle| \leq (8.9 \div 8.1) \times 10^{-2} \text{ eV}$$

! $|\langle m \rangle|$ in the range of sensitivity of the upcoming $(\beta/\beta)_\text{or}$ -decay experiments.

! The red colour denotes the just CPV region which can be spanned only in presence of CP violation.

MASS QUASI-DEGENERACY

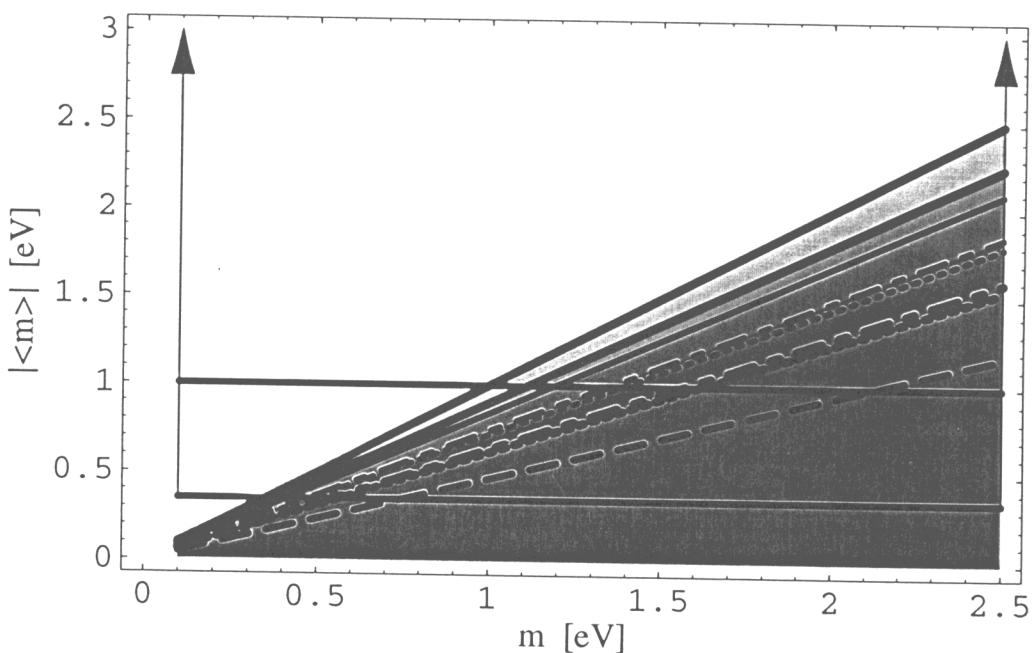
characterized by: $m_1 \approx m_2 \approx m_3 \equiv m$



$$|\langle m \rangle| = m_1 \left| (\cos^2 \theta_0 + \sin^2 \theta_0 e^{i\alpha_{21}})(1 - |U_{e3}|^2) + |U_{e3}|^2 e^{i\alpha_{31}} \right|$$

$|\langle m \rangle|$ can provide information on m_1 and α_{21}, α_{31} .

$|\langle m \rangle|$ vs $m_1 \approx m$ for the LMA MSW case.



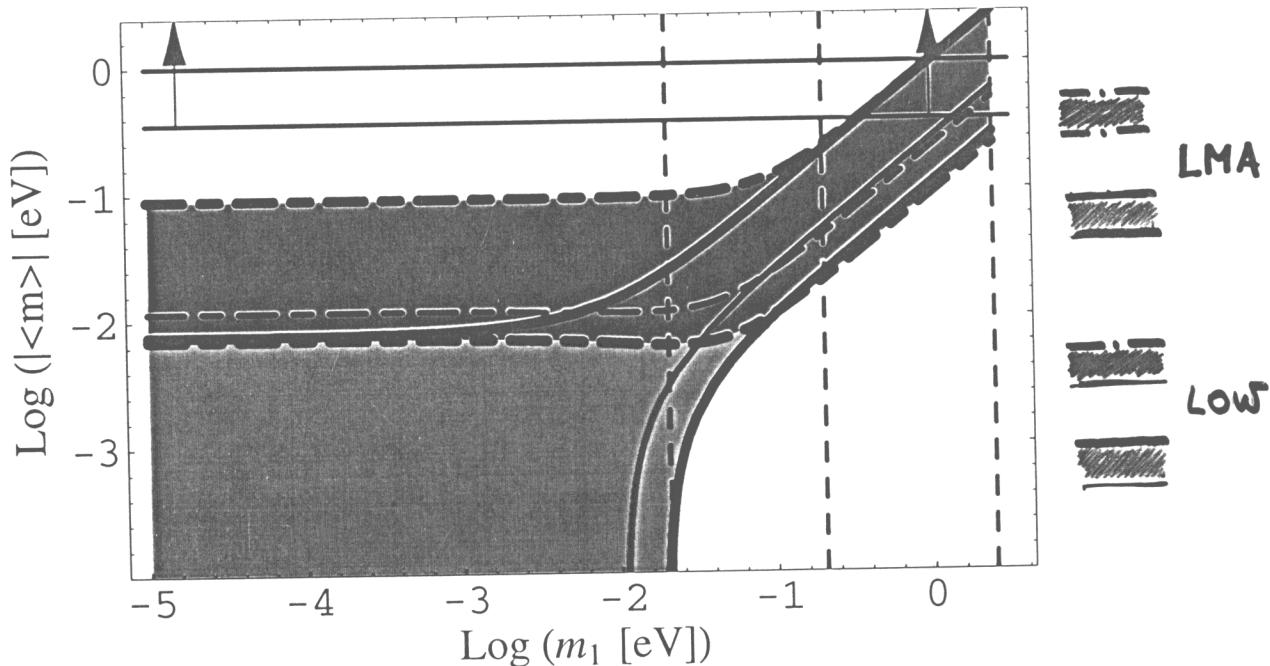
$$0 \text{ eV} \leq |\langle m \rangle| \leq 2.5 \text{ eV}$$

- part of the allowed range is already excluded by the Heidelberg - Moscow bound.
- ! • most of the allowed region is in the range of sensitivity of present and upcoming $(\beta\beta)_{0\nu}$ -decay experiments.

!

The value of $|\langle m \rangle|$ can help to resolve the ambiguity $\Delta m^2_0 = \Delta m^2_{21}$, Δm^2_{31} and to get information on m_1 .

$\log(|\langle m \rangle| [\text{eV}])$ vs $\log(m_1 [\text{eV}])$ for $\Delta m^2_0 = \Delta m^2_{21}$ (green) and $\Delta m^2_0 = \Delta m^2_{32}$ (blue):



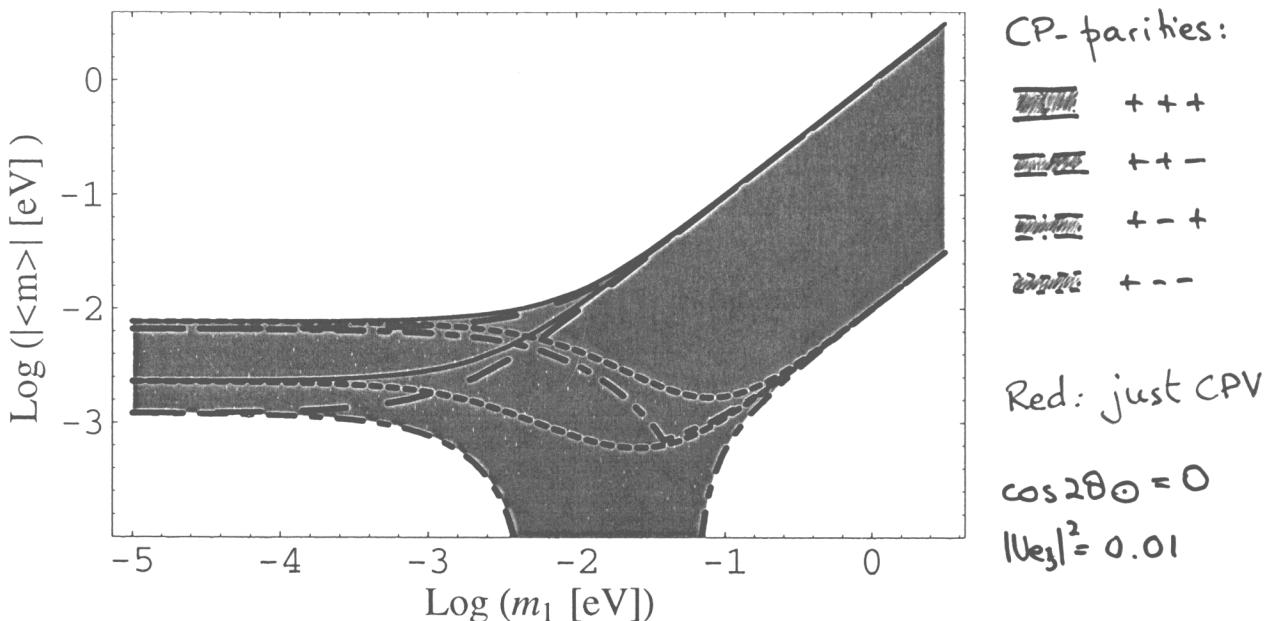
- 1) If $|\langle m \rangle| \geq 0.01 \text{ eV}$, the hierarchical spectrum is ruled out.
- 2) If $|\langle m \rangle| \geq 0.1 \text{ eV}$, the inverted hierarchical spectrum is disfavoured and the mass quasi-degeneracy case is favoured.

INFORMATION ON m_1

CASE A.

$$\Delta m^2_{\odot} = \Delta m^2_{21}$$

From the determination of the value of $|km| > 1$, we could get information on m_1 and CPV.



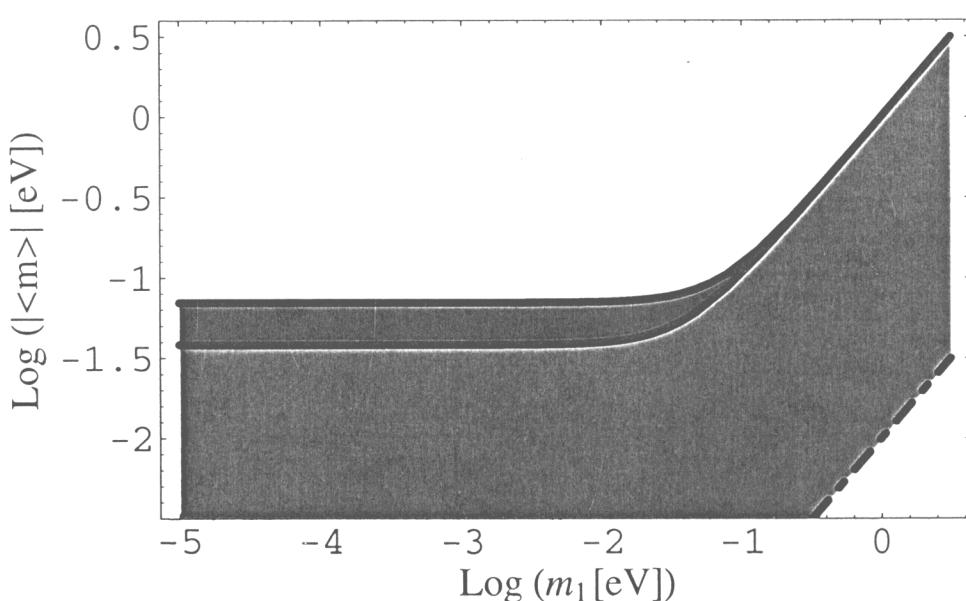
Log ($|km|$ [eV]) vs Log (m_1 [eV]) for LMA.

- If $|km| > 10^{-2}$ eV, the value of m_1 is constrained in a range and in particular has a lower bound.
- If 10^{-3} eV $< |km| < 10^{-2}$ eV, the value of m_1 has only an upper bound.

To infer information on CPV we need a constraint on m_1 . From the values of m_1 (KATRIN) and $|km|$ ($(\beta\beta)_{0\nu}$ -decay) it will be possible to establish if CP-parity is violated in the leptonic sector through Majorana CPV phases.

CASE B

$$\Delta m_{\odot}^2 = \Delta m_{32}^2$$



$$|\langle m \rangle|^2 = 0.01$$

$$\cos 2\theta_0 = 0$$

(LMA)

CP-parities:

	$\pm ++$
	$\pm +-$

Red: just CPV

- i) If $|\langle m \rangle| > 10^{-1}$ eV, m_1 is constrained to lie in a range of values and in particular has a lower bound. Only the $\pm ++$ CP-parities patterns are allowed and the value of m_1 is fixed (if CP conservation) or we have CP-violation.
- ii) If $0.04 \text{ eV} < |\langle m \rangle| < 0.1 \text{ eV}$, m_1 has an upper limit from ${}^3\text{H}-\beta$ decay experiments. Only the $\pm ++$ CP-parities patterns are allowed but m_1 is not constrained (CP-conservation) or CPV.
- iii) If $|\langle m \rangle| < 0.04$ eV, m_1 has only an upper bound. In the case of CP conservation only the $\pm +-$ CP-parity patterns are allowed with fixed value of m_1 .
If $m_1 < 0.5$ eV and $|\langle m \rangle| > 0.005$ eV, then ! we have CP-violation.

INFORMATION ON CP-VIOLATION

From a measurement of $|\langle m \rangle|$, it might be possible to get information on the CPV phases:

It depends on:

i) the solution of the solar-ν problem:

- for the LMA (LOW) solution, currently favoured, it will be possible to constrain at least one of the CPV phases.
- for the SMA solution, it will be very difficult to get any information on α_{21} and α_{31} as the dependence of $|\langle m \rangle|$ on them is suppressed by the values of the mixing angles.

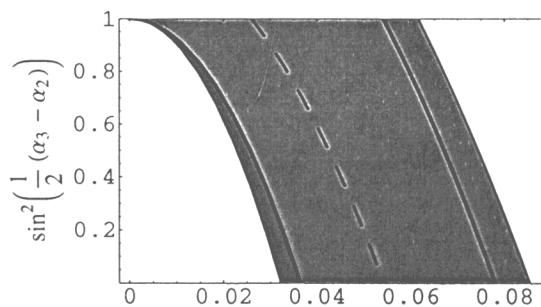
ii) the spectrum considered:

- $\Delta m^2_{\odot} = \Delta m^2_{21}$, m_3 negligible:
 $|\langle m \rangle|$ is beyond the sensitivity of present and upcoming experiments.
- $\Delta m^2_{\odot} = \Delta m^2_{32}$, m_1 negligible:
 $|\langle m \rangle|$ can fix the value of $\alpha_{31} - \alpha_{21}$.
- $\Delta m^2_{\odot} = \Delta m^2_{32}, \Delta m^2_{21}$, $m_1 > 0.2 \text{ eV}$:
 $|\langle m \rangle|$ can constrain α_{21} and α_{31} .

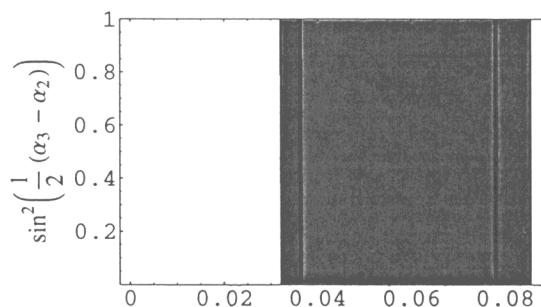
$$\Delta m_{\odot}^2 = \Delta m_{32}^2, m_3 \text{ negligible}$$

From the measurement of $|\langle m \rangle|$, it might be possible to infer the value of the CPV phase α_{32} :

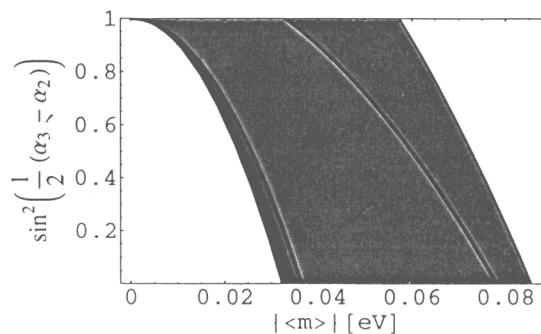
$$\sin^2 \frac{\alpha_{32}}{2} = \left(1 - \frac{|\langle m \rangle|^2}{\Delta m_{\text{atm}}^2 (1 - |\langle U_{e1} \rangle|^2)} \right) \frac{1}{\sin^2 2\theta_{\odot}} \quad (\text{S. Bilenky et al '96})$$



LMA



SMA

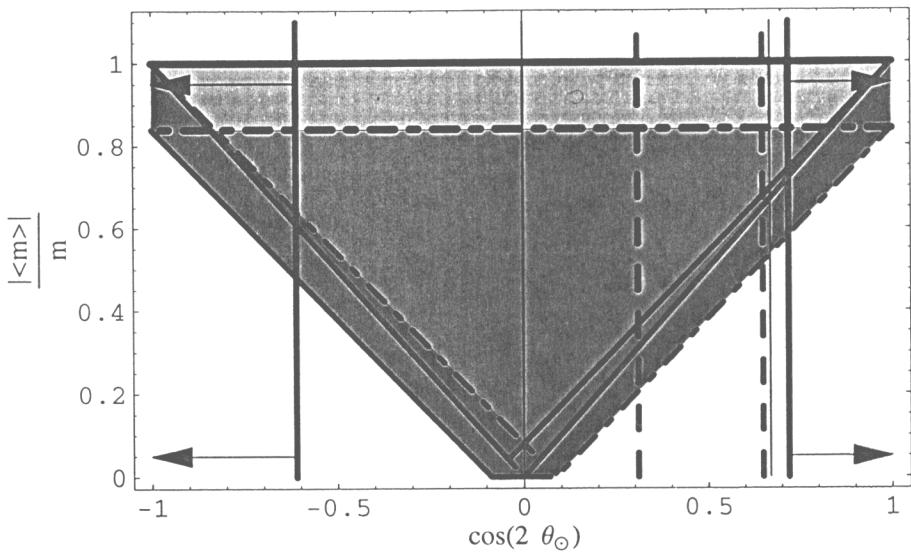


LOW

$\sin^2 \frac{\alpha_{32}}{2}$ vs $|\langle m \rangle|$ according to the different solutions of the θ_{\odot} -problem.

In the case of the SMA solution, we have no constraints on α_{32}

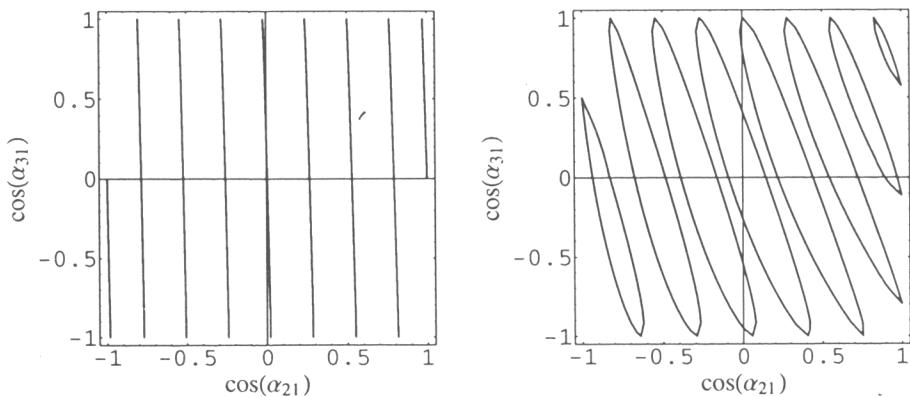
$$\Delta m_0^2 = \Delta m_{21}^2, \Delta m_{31}^2, m_1 > 0.2 \text{ eV}$$



The red colour denotes the just CPV region.

Two CPV phases are relevant: α_{21} and α_{31} .

If $|m|$ is determined, we cannot fix the values of both phases but we can constrain their ranges. We plot $\cos \alpha_{31}$ vs $\cos \alpha_{21}$ for fixed values of the parameters and different values of $|m|/m$.



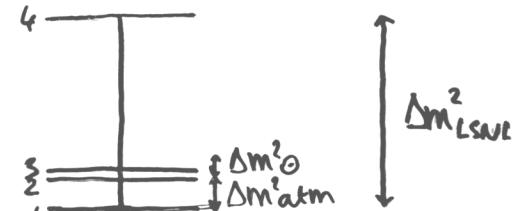
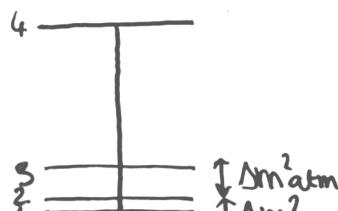
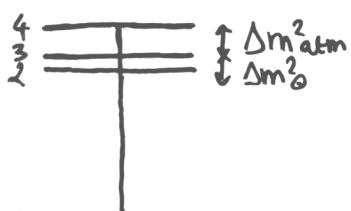
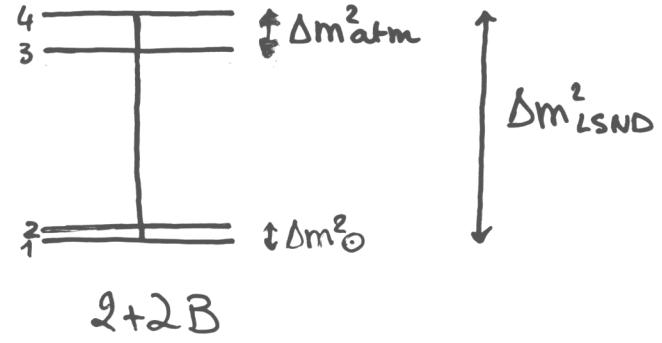
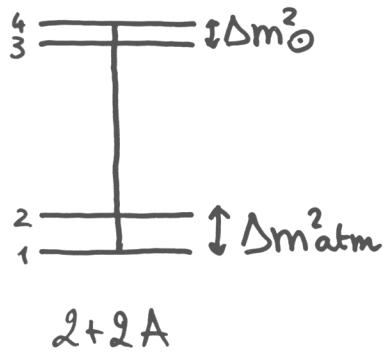
$$\begin{aligned} \frac{|m|^2}{m^2} &= (\cos^4 \theta_0 (1 - |U_{e3}|^2)^2 + \sin^2 \theta_0 (1 - |U_{e3}|^2)^2 + |U_{e3}|^4) \\ &= 2 \cos^2 \theta_0 |U_{e3}|^2 (1 - |U_{e3}|^2) \cos \alpha_{31} + 2 \sin^2 \theta_0 \cos^2 \theta_0 (1 - |U_{e3}|^2)^2 \cos \alpha_{31} \\ &\quad + 2 \sin^2 \theta_0 |U_{e3}|^2 (1 - |U_{e3}|^2) \cos(\alpha_{21} - \alpha_{31}) \end{aligned}$$

$|\langle m \rangle|$ IN THE 4-D MIXING CASE

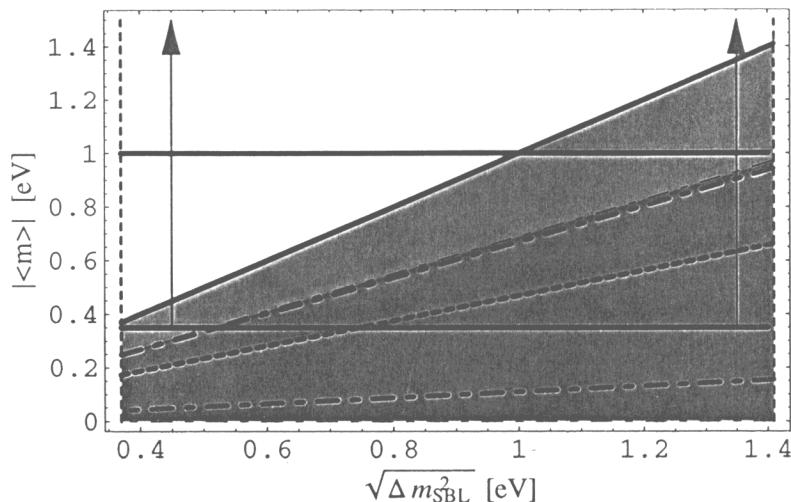
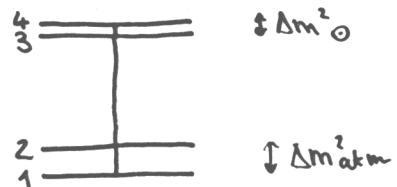
If LSND results will be confirmed, then to explain atmospheric, solar and LSND results in terms of oscillations we need to consider the mixing of 4 neutrinos.

The predictions on $|\langle m \rangle|$, using the oscillations parameters and varying freely the non-oscillation ones $m_1, \alpha_{21}, \alpha_{31}, \alpha_{41}$ depend strongly on the spectrum considered.

We consider the following possibilities:



2+2A

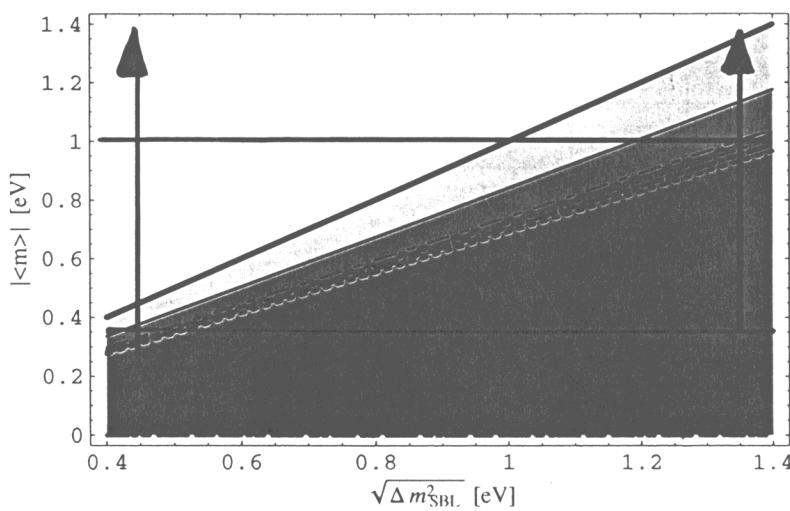
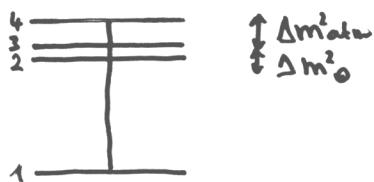


(LMA)

$$m_\beta = \sqrt{m_1^2 + \Delta m_{SBL}^2} \gtrsim 0.4 \text{ eV}$$

completely tested
by KATRIN

3+1A

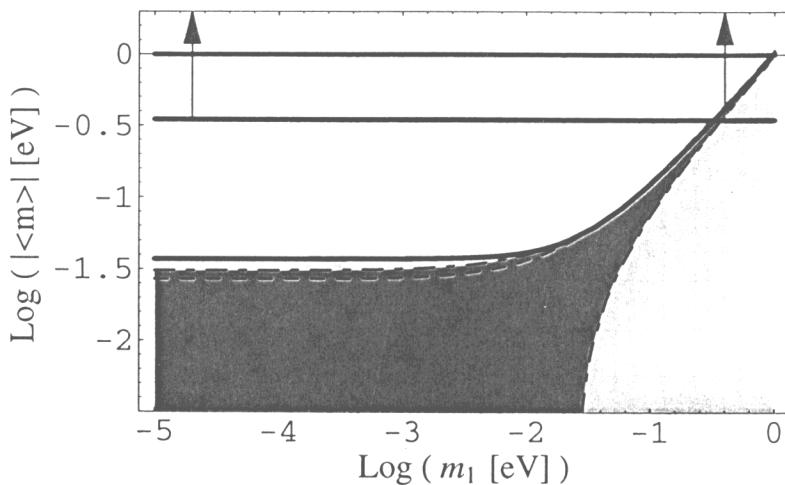
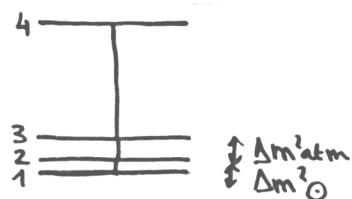


(LMA)

In both cases, most of the range of $|⟨m⟩|$ is in the region of sensitivity of present and upcoming experiments.

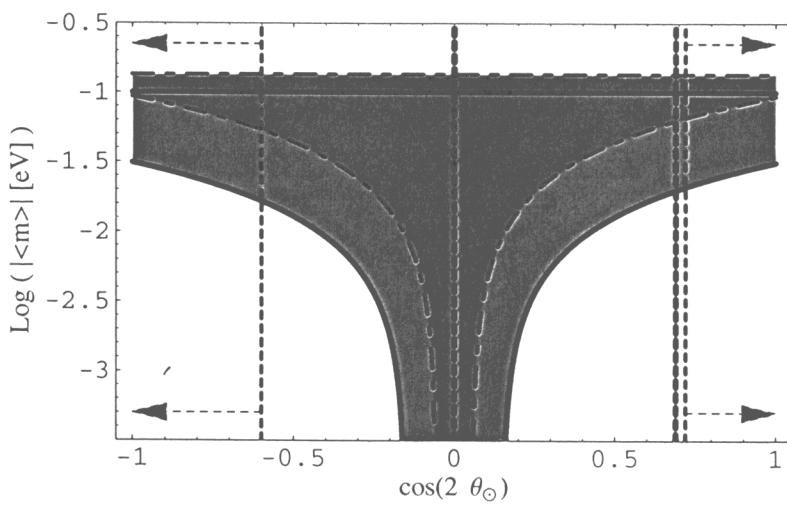
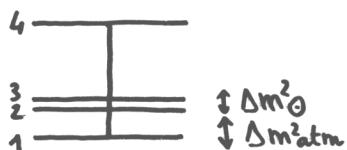
The lower bound on $|⟨m⟩|$ depends strongly on $\cos 2\theta$.

3+1 B



Blue: SMA
Yellow + Blue:
LMA, LOW.

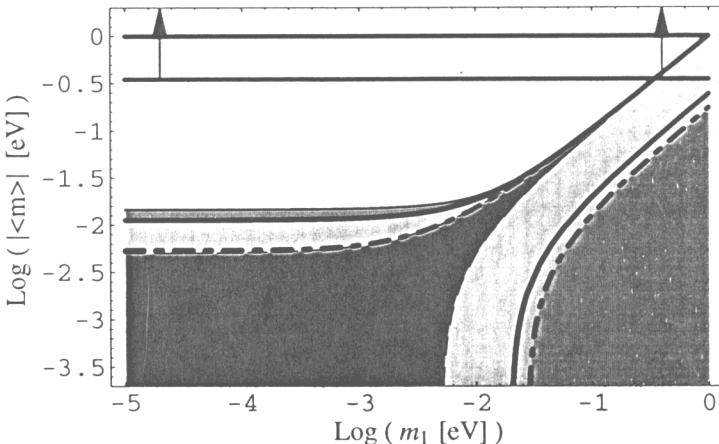
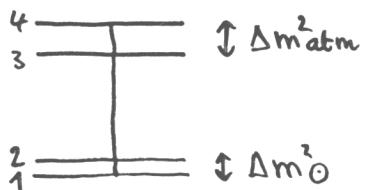
3+1C



(LMA)

Unless $m_1 \gtrsim 0.1$ eV, $|⟨m⟩|$ is much lower than in the two previous cases and only upcoming experiments will be sensitive to its value.

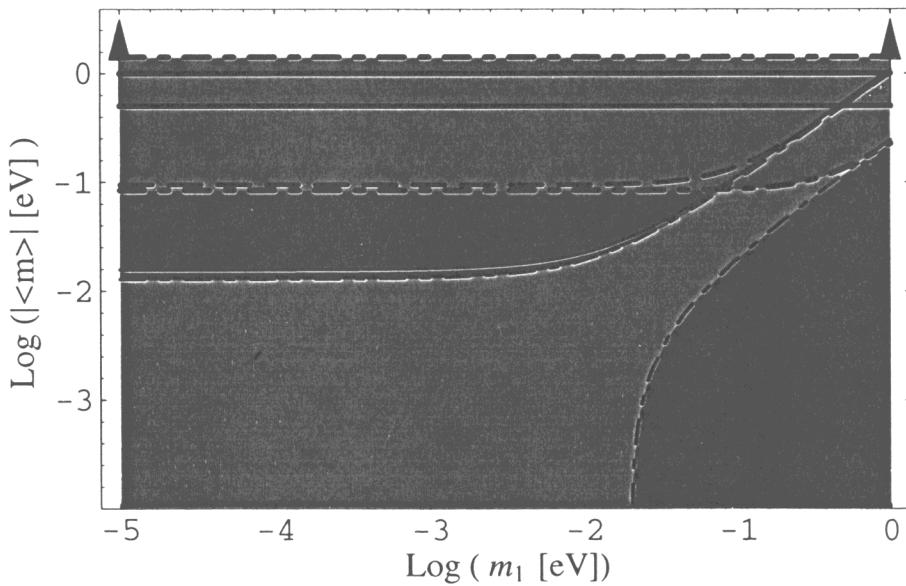
2+2B



Blue: SMA
Yellow + Blue:
LHA, LOW

Unless $m_1 \gtrsim 0.1 \text{ eV}$, $|⟨m⟩| < 0.02 \text{ eV}$, far beyond the sensitivity of upcoming experiments.

We can summarize the predictions on $|⟨m⟩|$ as:



If $|⟨m⟩| > 0.1 \text{ eV}$, the allowed spectra are 2+2A and 3+1A while the 2+2B, 3+1B and 3+1C are ruled out.
If $|⟨m⟩| < 0.1 \text{ eV}$, the 2+2A spectrum is disfavoured.

CONCLUSIONS

If $(\beta\beta)_{0\nu}$ -decay is detected, then we could conclude that neutrinos are massive Majorana particles and the total lepton charge L is not conserved ($\Delta L = 2$).

• 3-ν MIXING CASE

With an improvement of the precision of the oscillation parameters, Δm^2_0 , Δm^2_{atm} , θ_0 , θ_{atm} , the determination of the value of $|km| \geq 0.01$ eV would lead to obtain information on m_1 and then on the type of mass spectrum.

In particular, in the 3-ν mixing case:

- i) for the LMA and LOW solutions of the ν_0 -anomaly, we could obtain either a range for m_1 (if $\Delta m^2_0 = \Delta m^2_{21}$ or $\Delta m^2_0 = \Delta m^2_{32}$, $m_1 > 10^{-2}$ eV), or an upper bound on m_1 (if $\Delta m^2_0 = \Delta m^2_{32}$, $m_1 < 10^{-2}$ eV). With additional information on m_1 (KATRIN) we could establish CPV in the lepton sector and either the value of CPV phases α_{21}, α_{31} (if CPV) or the CP-parity patterns of the ν (if CPC).
(* There are exceptional cases where $|km|$ alone \Rightarrow CPV).
- ii) for the SMA solution, we could obtain either the value of m_1 ($\Delta m^2_0 = \Delta m^2_{21}$ or $\Delta m^2_{32} = \Delta m^2_0$, $m_1 > 10^{-2}$ eV) or an upper bound on m_1 ($\Delta m^2_0 = \Delta m^2_{32}$, $m_1 < 10^{-2}$ eV). No information on CP can be inferred.

- 4- ν MIXING CASE

The LSND results will be tested within the next few years by the MiniBooNE experiment.

For 4 of the allowed spectra (2+2A, 3+1A, 3+1B, 3+1C), large parts of the ranges of $|km\rangle$ are in the range of sensitivity of present and upcoming ($\beta\beta$) experiments.

The 2+2A and 3+1A spectra can be completely tested by the KATRIN experiments.

For the 2+2B spectrum, $|km\rangle$ is beyond the sensitivity of the upcoming experiments.