

# Introduzione alle masse e ai mescolamenti dei neutrini (I)



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The lectures are intended for a broad audience of students or researchers from different fields in particle physics

The goal is to “get you interested” in neutrino physics, by recalling basic neutrino properties and phenomena, which will be further discussed in more specialized lectures

Some simple exercises are also proposed (with solutions)

People interested in further reading can usefully browse the “Neutrino Unbound” website: [www.nu.to.infn.it](http://www.nu.to.infn.it) , or just mail me for advice about specific topics: [eligio.lisi@ba.infn.it](mailto:eligio.lisi@ba.infn.it)

**Feel free to stop me and ask questions at any time!**

# Outline:

Pedagogical Introduction  
Neutrino masses and spinor fields  
Neutrinoless double beta decay  
2 $\nu$  & 3 $\nu$  oscillations in vacuum  
[Homework]

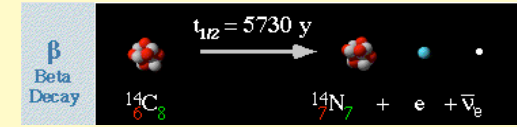
I

Recap  
2 $\nu$  oscillations in matter  
Solar and KamLAND oscillations  
Absolute neutrino masses  
[Homework]

II

# 2010: the 80<sup>th</sup> Neutrino Birthday!

The neutrino was invented in 1930 by Wolfgang Pauli as a "desperate remedy" to explain the continuous  $\beta$ -ray spectrum via a 3-body decay, e.g.,



*Original: Photograph of Pauli 1933*  
Abschrift/15.12.56 PM

Offener Brief an die Gruppe der Radioaktiven bei der  
Gauvereins-Tagung zu Tübingen.

Abschrift  
Physikalisches Institut  
der Eidg. Technischen Hochschule  
Zürich

Zürich, 4. Dec. 1930  
Oliverastrasse

Liebe Radioaktive Damen und Herren,

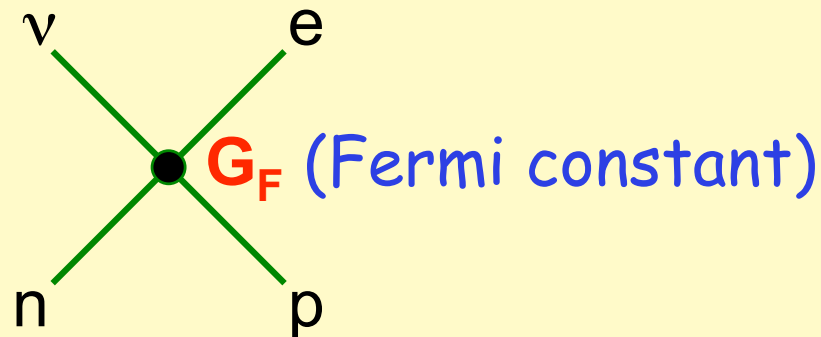
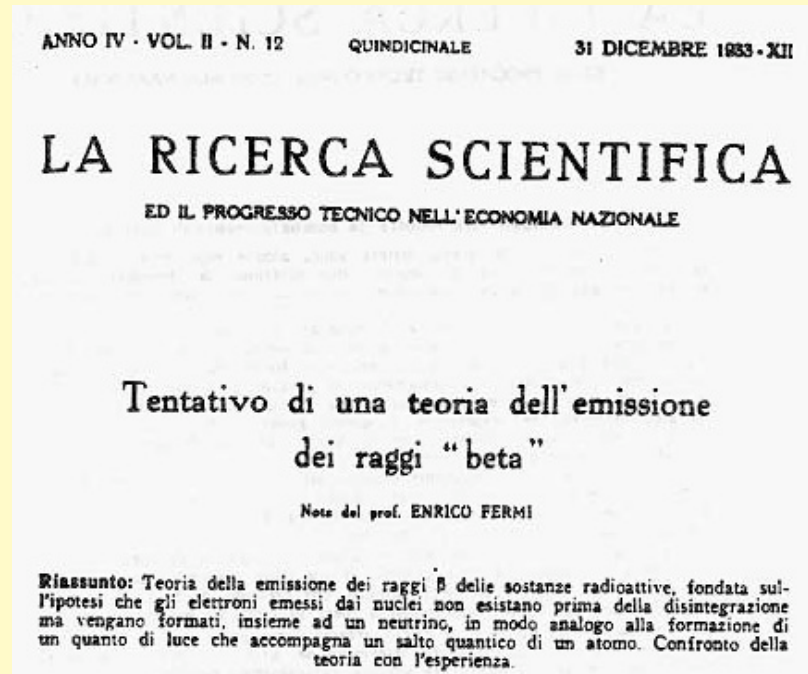
Wie der Ueberbringer dieser Zeilen, den ich baldvollaft  
anzuhören bitte, Ihnen das näherem auseinandersetzen wird, bin ich  
angesichts der "falschen" Statistik der N- und Li-6 Kerne, sowie  
des kontinuierlichen beta-Spektrums auf einen verzweifelten Ausweg  
verfallen: um den "Wechselstich" (1) der Statistik und dem Energiesatz  
zu retten. Nämlich die Möglichkeit, es könnten elektrisch neutrale  
Teilchen, die ich Neutronen nennen will, in den Kernen existieren,  
welche den Spin 1/2 haben und das Ausschliessungsprinzip befolgen und  
sich von Lichtquanten ausserdem noch dadurch unterscheiden, dass sie  
nicht mit Lichtgeschwindigkeit laufen. Die Masse der Neutronen  
müsste von derselben Grössenordnung wie die Elektronenmasse sein und  
jedenfalls nicht grösser als 0.01 Protonenmasse. Das kontinuierliche  
beta-Spektrum wäre dann verständlich unter der Annahme, dass beim  
beta-Zerfall mit dem Elektron jeweils noch ein Neutron emittiert  
wird, derart, dass die Summe der Energien von Neutron und Elektron  
konstant ist.



Kinematics: spin 1/2, tiny mass, zero electric charge



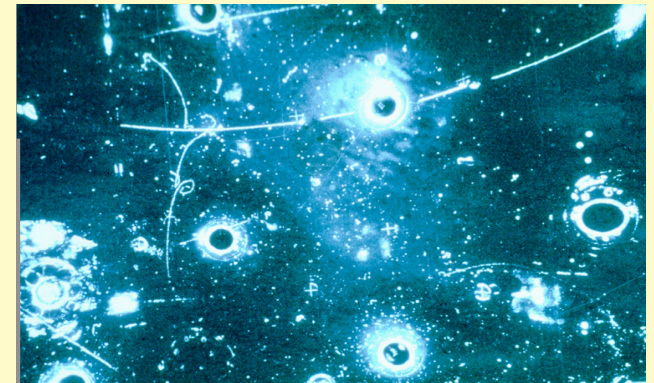
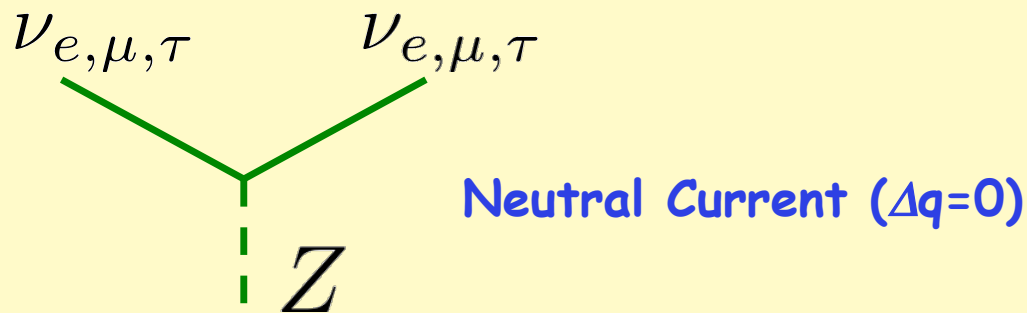
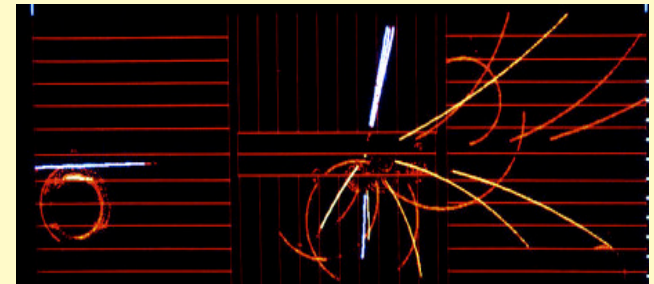
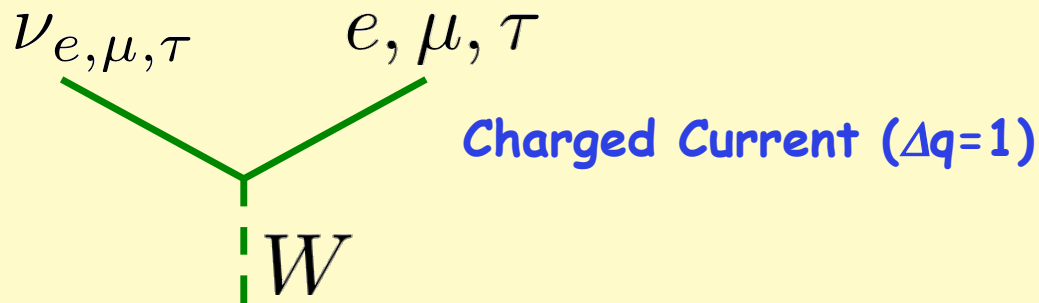
The name "neutrino" (= "little neutral one", in Italian) was actually invented by Enrico Fermi, who first proposed in 1933-34 a theory for its **dynamics** (weak interactions)



Many decades of research have revealed other properties of the **neutrino**. For instance, there are 3 different neutrino "**flavors**"

$$\begin{pmatrix} \nu_e \\ e \end{pmatrix} \begin{pmatrix} \nu_\mu \\ \mu \end{pmatrix} \begin{pmatrix} \nu_\tau \\ \tau \end{pmatrix} \begin{matrix} \leftarrow q = 0 \\ \leftarrow q = -1 \end{matrix} \quad (\Delta q = 1)$$

and their Fermi interactions are mediated by a charged **vector boson W**, with a neutral counterpart, the **Z boson**

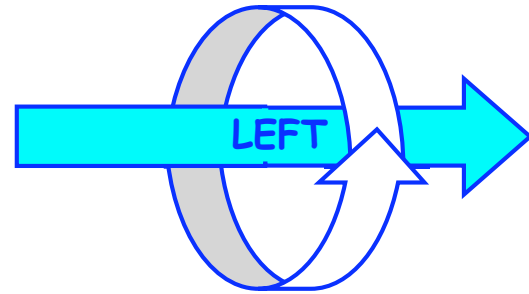


(For the other vertex in  $W$ ,  $Z$  exchange: See lecture by L. Ludovici)

Such interactions are chiral (= not mirror-symmetric):

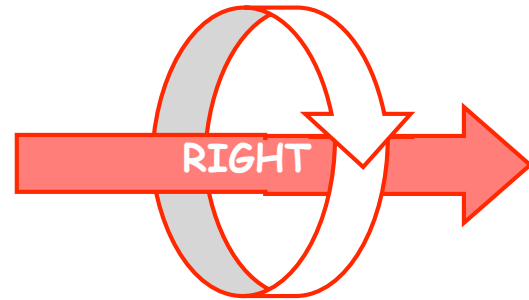
Neutrinos are created in  
a left-handed (LH) state

$\nu$



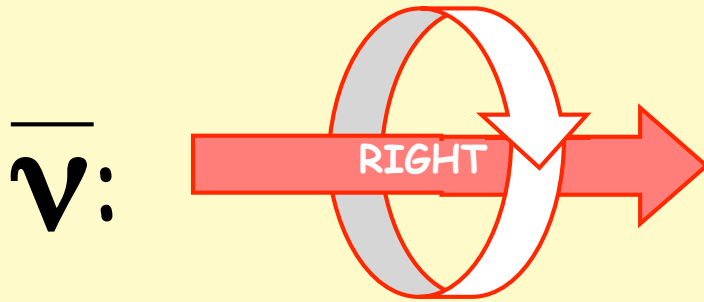
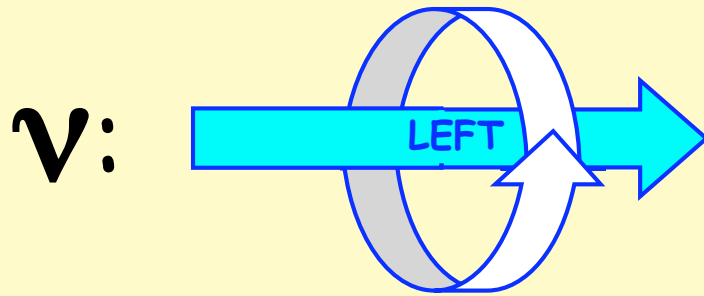
Anti-nus are created in  
a right-handed (RH) state

$\bar{\nu}$



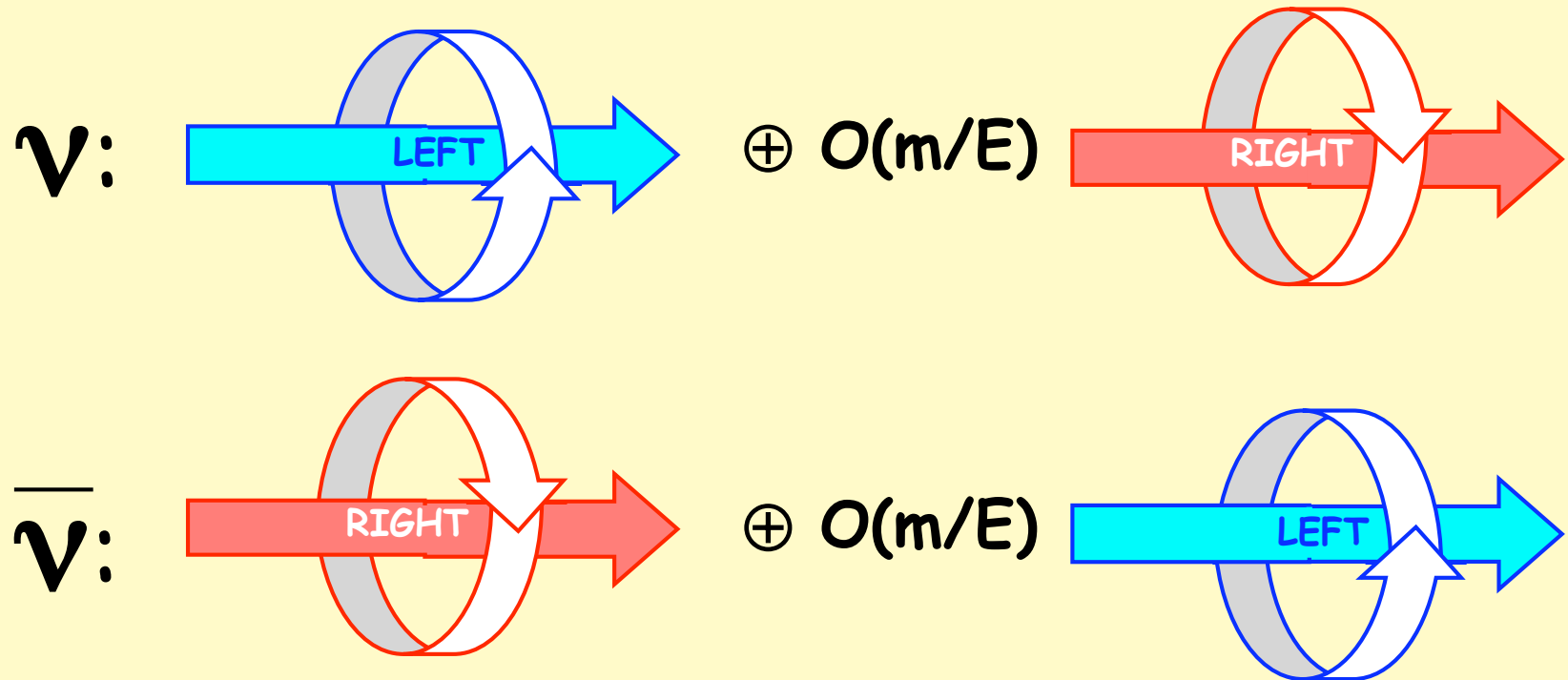
Neutrinos couldn't see themselves in a mirror... like vampires!

For massless neutrinos: handedness is a constant of motion



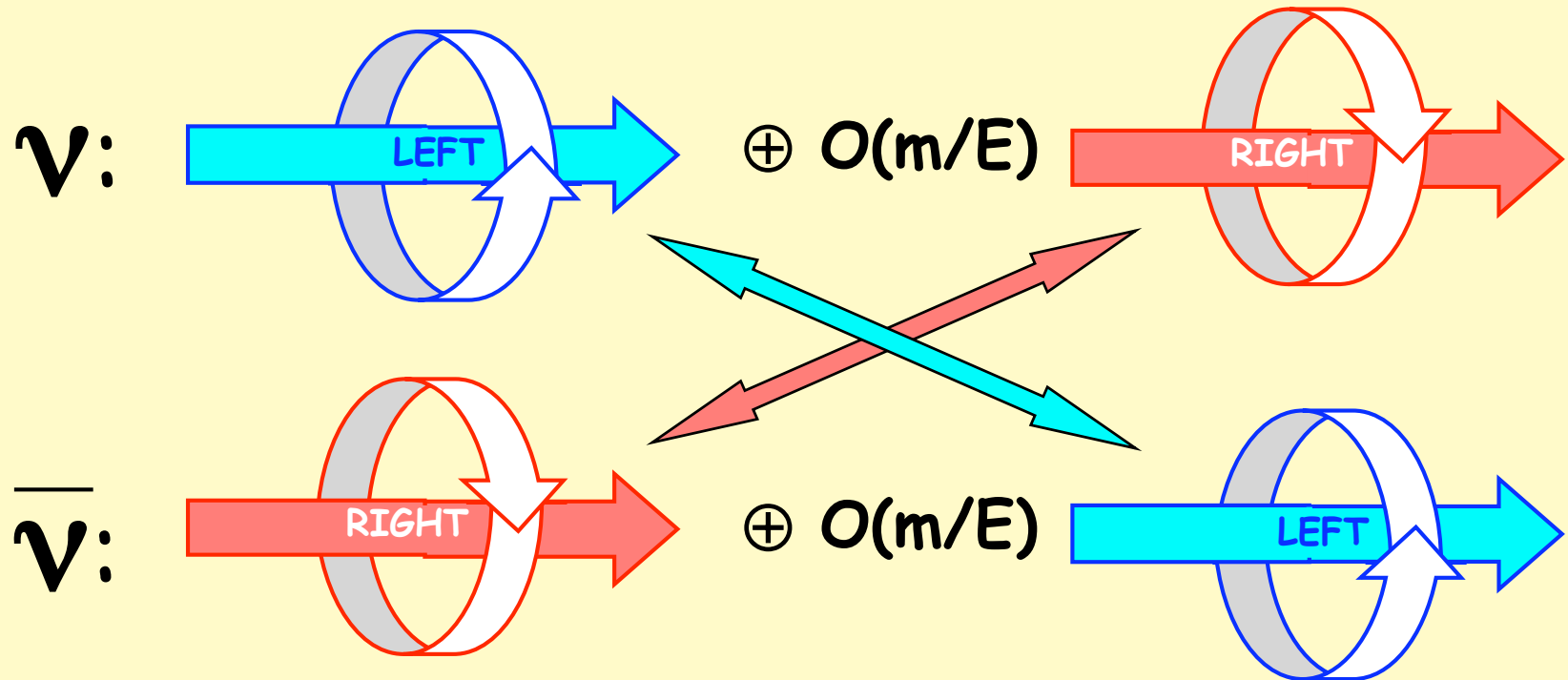
2 independent d.o.f.: massless ("Weyl") 2-spinor

But: massive  $\nu$  can develop the “wrong” handedness at  $O(m/E)$   
 (the Dirac equation mixes RH and LH states for  $m_\nu \neq 0$ ):



If these 4 d.o.f. are independent: massive (“Dirac”) 4-spinor  
 [→ Distinction between neutrinos and antineutrinos, as for electrically charged fermions. Can define a “lepton number”]

But, for neutral fermions, 2 components might be identical !

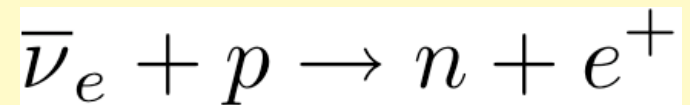
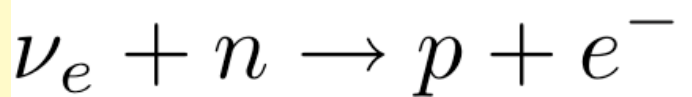


Massive ("Majorana") 4-spinor with 2 independent d.o.f.

[No distinction between neutrinos and antineutrinos, up to a phase:  
 A *very* neutral particle: no electric charge, no leptonic number...]



**Exercise 1.** Define the electron neutrino as the neutral particle emitted in  $\beta^+$  decay, and the electron antineutrino as the neutral particle emitted in  $\beta^-$  decay. Reactions which have been observed:



while the following reactions have not been observed:



If neutrinos and antineutrinos are different (Dirac case), that's easy to understand. Try to understand the same (non)observations in the case of Majorana neutrinos.

# Summary of options for neutrino spinor field:

$m=0$ ,  
Weyl:

$$\psi = \psi_R$$

or

$$\psi = \psi_L$$

massless field  
with 2 d.o.f.

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$m \neq 0$ ,  
Majorana:

$$\psi = \psi_R + \psi_R^c = \psi^c$$

or

$$\psi = \psi_L + \psi_L^c = \psi^c$$

massive field  
with 2 d.o.f.

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$m \neq 0$ ,  
Dirac:

$$\psi = \psi_R + \psi_L \neq \psi^c$$

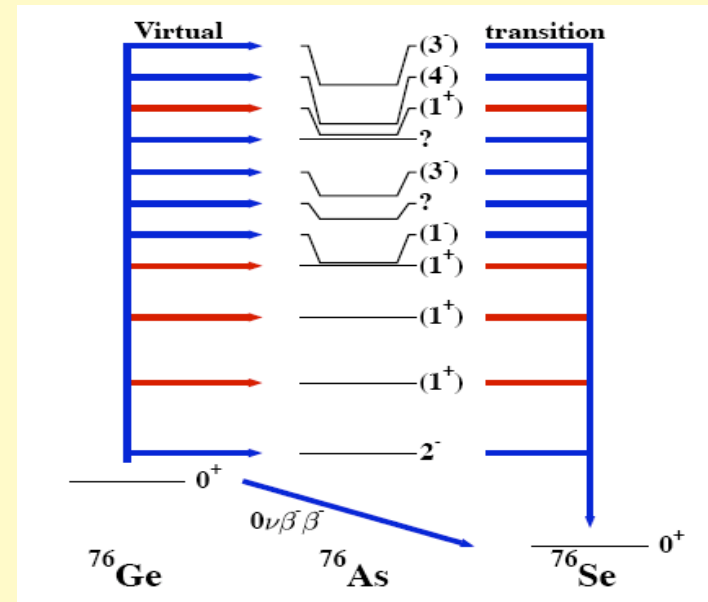
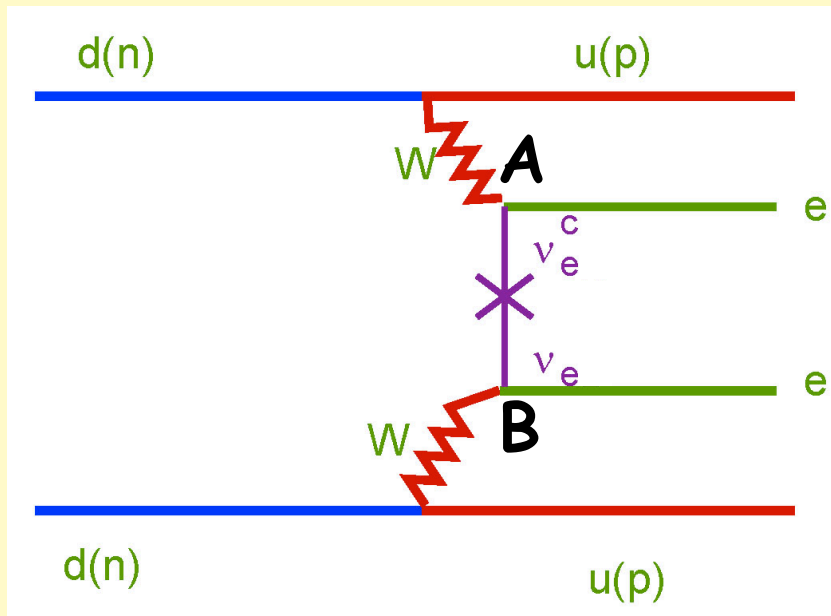
massive field  
with 4 d.o.f.

**Conjugation operator:**  $\psi^c = \mathcal{C}(\psi) = i\gamma^2\psi^*$  .  $\psi_{\text{antiparticle}} = \mathcal{C}(\psi_{\text{particle}})$

Appendix: Majorana masses and "see-saw" mechanism [+ see talk by F. Feruglio]

**Experiments: A unique experimental handle →**

## Neutrinoless double beta decay: $(A,Z) \rightarrow (A,Z+2)+2e$

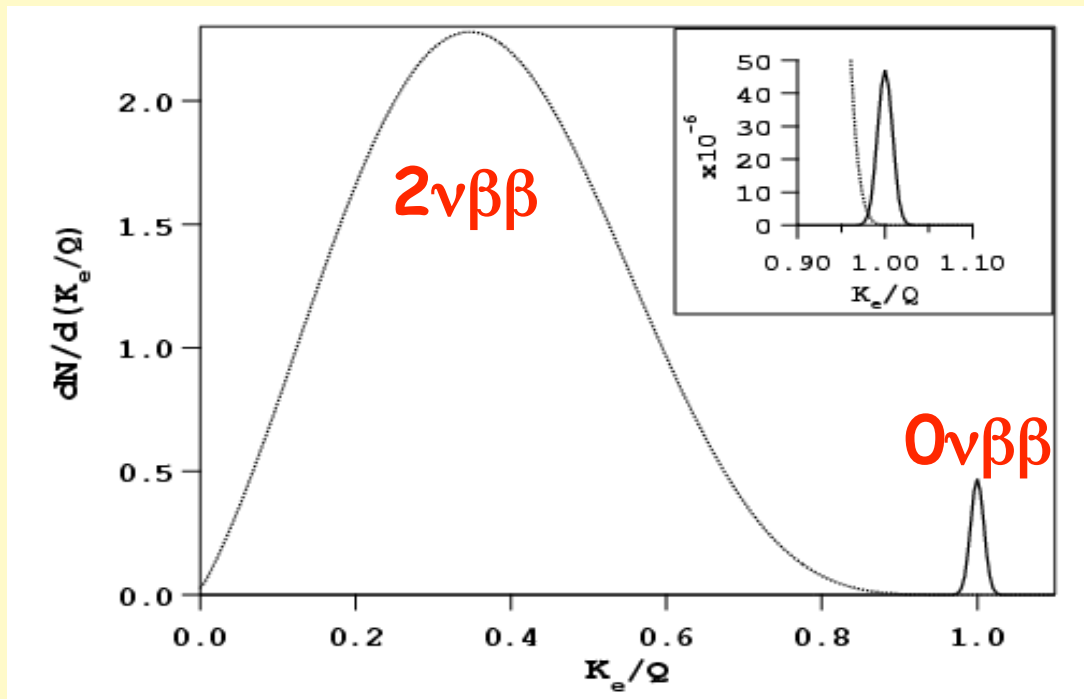


Can occur only for Majorana neutrinos. Intuitive picture:

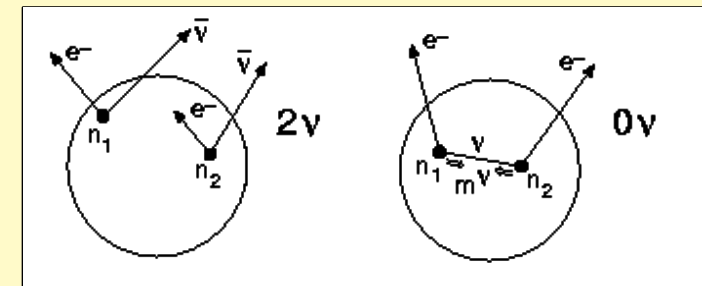
- 1) A RH antineutrino is emitted at point "A" together with an electron
- 2) If it is massive, at  $O(m/E)$  it develops a LH component (not possible if Weyl)
- 3) If neutrino=antineutrino, this component is a LH neutrino (not possible if Dirac)
- 4) The LH (Majorana) neutrino is absorbed at "B" where a 2nd electron is emitted

[EW part is "simple". Nuclear physics part is rather complicated and uncertain.]

Experimentally: Look at sum energy of both electrons



Need to see the  $0\nu\beta\beta$  line emerge above background, at the endpoint of spectrum from "conventional" (and observed)  $2\nu\beta\beta$  decay.



Very rare to detect (if it occurs): doubly-weak and suppressed by  $m/E$ .  
Need to be tenacious... like E. Fiorini (see next lecture)

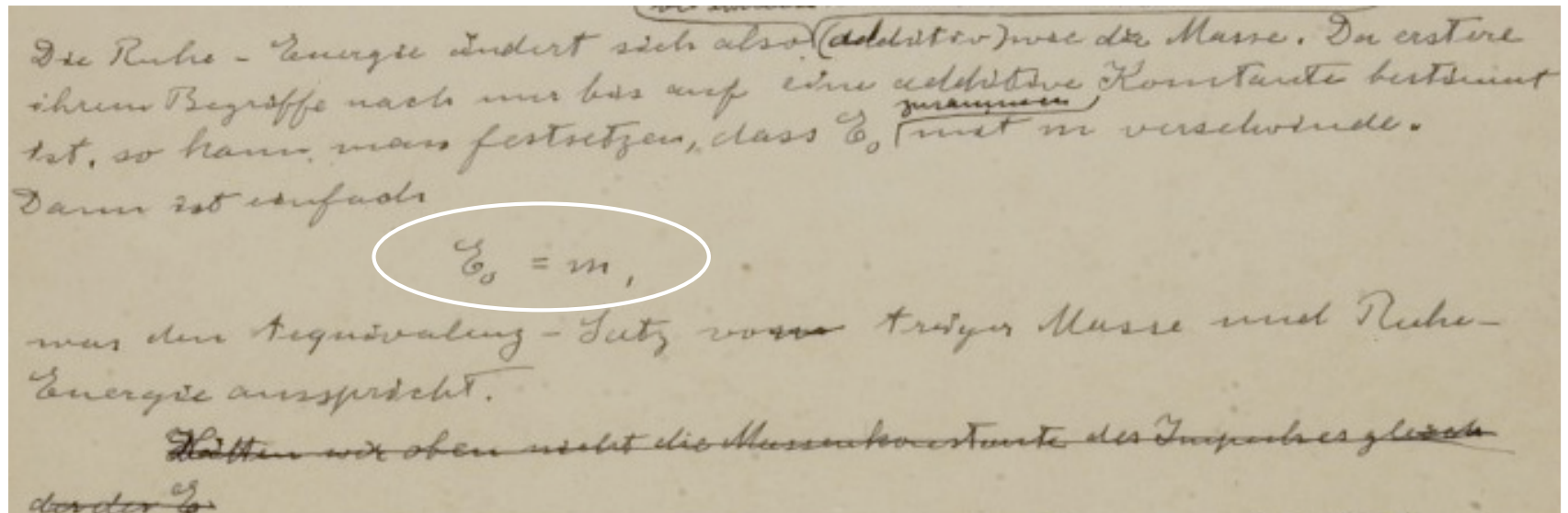
Recap: if neutrinos have mass, they can develop the “**wrong handedness**” with **amplitude of  $O(m_{\text{mass}}/E_{\text{energy}})$** . The only known chance to observe this tiny effect is  **$0\nu\beta\beta$  decay**.

But, if neutrinos are not only massive but mixed, they can also develop in the “**wrong flavor**” as a major consequence (“**neutrino flavor oscillations**”). This effect, despite being only of  **$O(m^2/E)$  in the phase**, can become observable over macroscopic distances (similar to optical interferometry).

**We shall now discuss the phenomenon of flavor oscillations, going from simplified approximation to more realistic scenarios.**

# Neutrino flavor oscillations in vacuum (2v)

The starting point is a century-old equation ...



... namely, for  $p \neq 0$ :

$$E = \sqrt{m^2 + p^2}$$

(in natural units)



Our ordinary experience takes place in the limit:  $p \ll m$

$$E \simeq m + \frac{p^2}{2m}$$

... while for neutrinos the proper limit is:  
 $p \gg m$

$$E \simeq p + \frac{m^2}{2p}$$

Energy difference between two neutrinos  $\nu_i$  e  $\nu_j$  with mass  $m_i$  e  $m_j$  in the same beam ( $p_i = p_j \simeq E$ ):

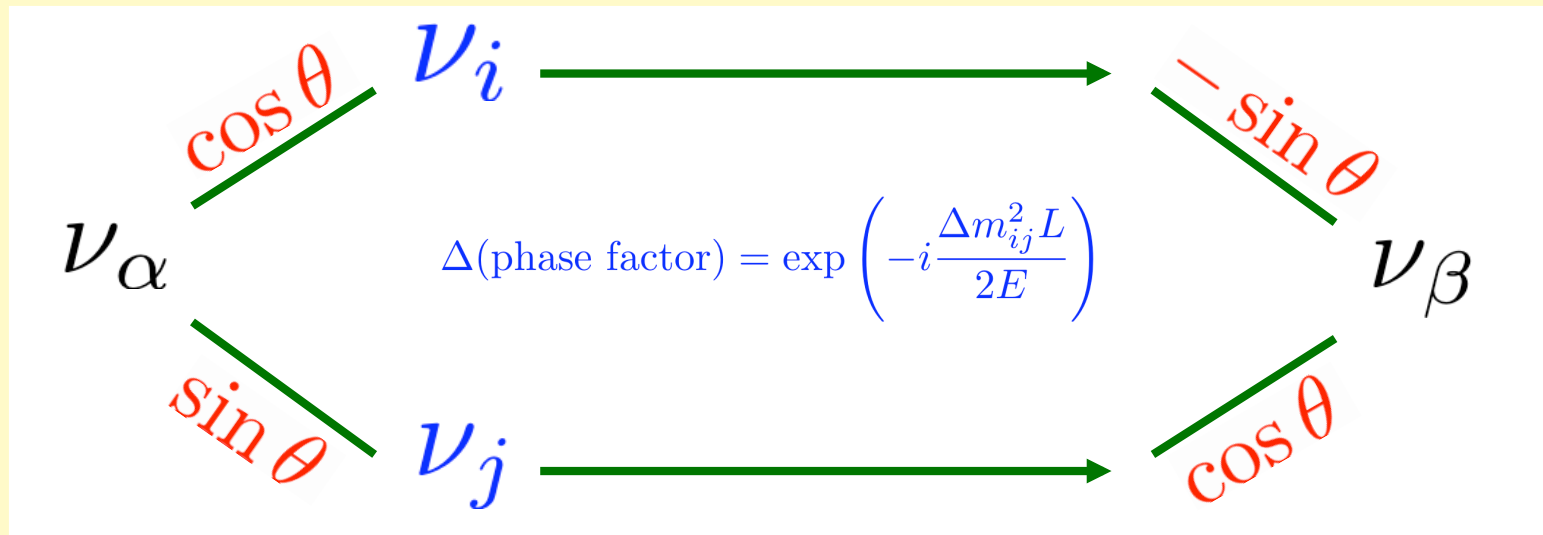
$$\Delta E \simeq \frac{\Delta m_{ij}^2}{2E}$$

**PMNS\***: neutrinos with definite mass ( $\nu_i$  and  $\nu_j$ ) might have NO definite flavor ( $\nu_\alpha$  e  $\nu_\beta$ ), e.g.,

$$\begin{pmatrix} \nu_\alpha \\ \nu_\beta \end{pmatrix} = \begin{pmatrix} \cos \theta & \sin \theta \\ -\sin \theta & \cos \theta \end{pmatrix} \begin{pmatrix} \nu_i \\ \nu_j \end{pmatrix}$$

\*Pontecorvo; Maki, Nakagawa & Sakata

Analogy with a two-slit interference experiment in vacuum:



This is the simplest case (only 2 neutrinos involved, no interactions with matter). It shows that, if neutrinos are massive and mixed (like quarks), then flavor is not a good quantum number during propagation. Indeed, it changes ("oscillates") significantly over a distance  $L$  ( $\approx \Delta t$ ) dictated by the uncertainty relation:

$$1 \sim \Delta E \Delta t \simeq \frac{m_i^2 - m_j^2}{2E} L$$

**Exercise 2.** Prove that a neutrino created with flavor  $\alpha$  can develop a different flavor  $\beta$  with a periodical oscillation probability in  $L/E$ :

$$P(\nu_\alpha \rightarrow \nu_\beta) = 4 \sin^2 \theta \cos^2 \theta \sin^2 \left( \frac{\Delta m_{ij}^2 L}{4E} \right) \quad (\text{B. Pontecorvo})$$

Amplitude  
(vanishes for  $\theta=0$  or  $\pi/2$ )

Phase difference  
(vanishes for degenerate masses)

**Note :** This is the flavor "appearance" probability.

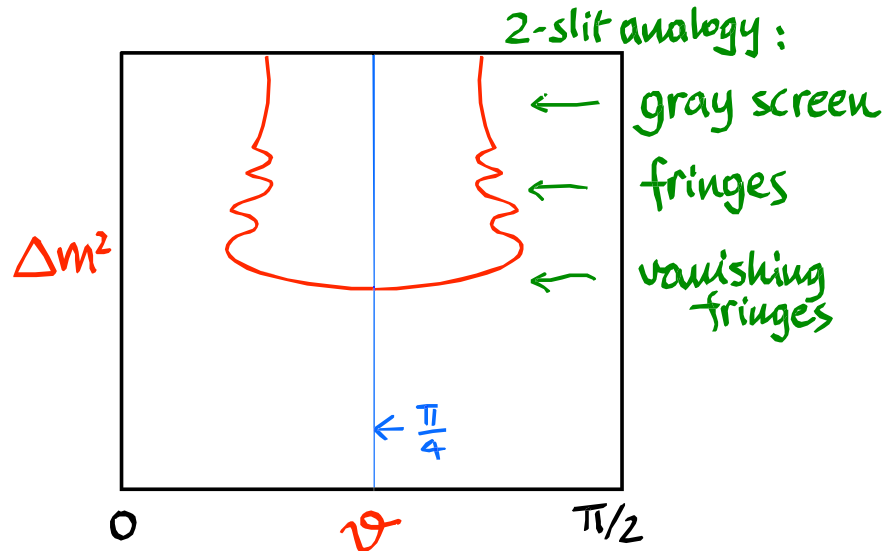
The flavor "disappearance" probability is the complement to 1.

**Exercise 3.** The oscillation effect depends on the **difference** of (squared) masses, not on the **absolute masses**. Why?

**Exercise 4.** Show that:

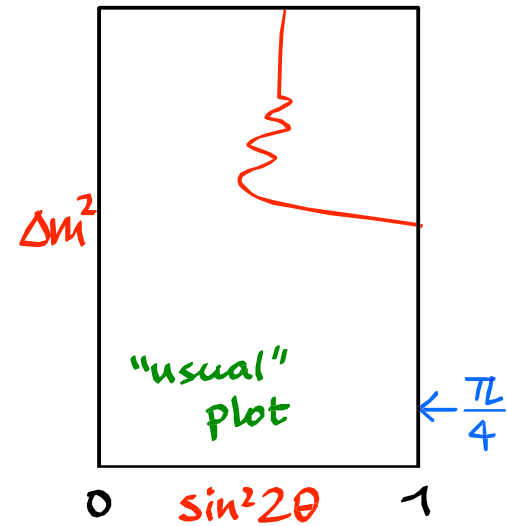
$$\frac{\Delta m^2 L}{4E} = 1.267 \left( \frac{\Delta m^2}{\text{eV}^2} \right) \left( \frac{L}{\text{km}} \right) \left( \frac{\text{GeV}}{E} \right)$$

## Typical iso- $\langle P_{\alpha\beta} \rangle$ contours



Octant symmetry:  $\theta \rightarrow \frac{\pi}{2} - \theta$  in  $P_{\mu\nu}$

If 2nd octant folded onto the 1st one:



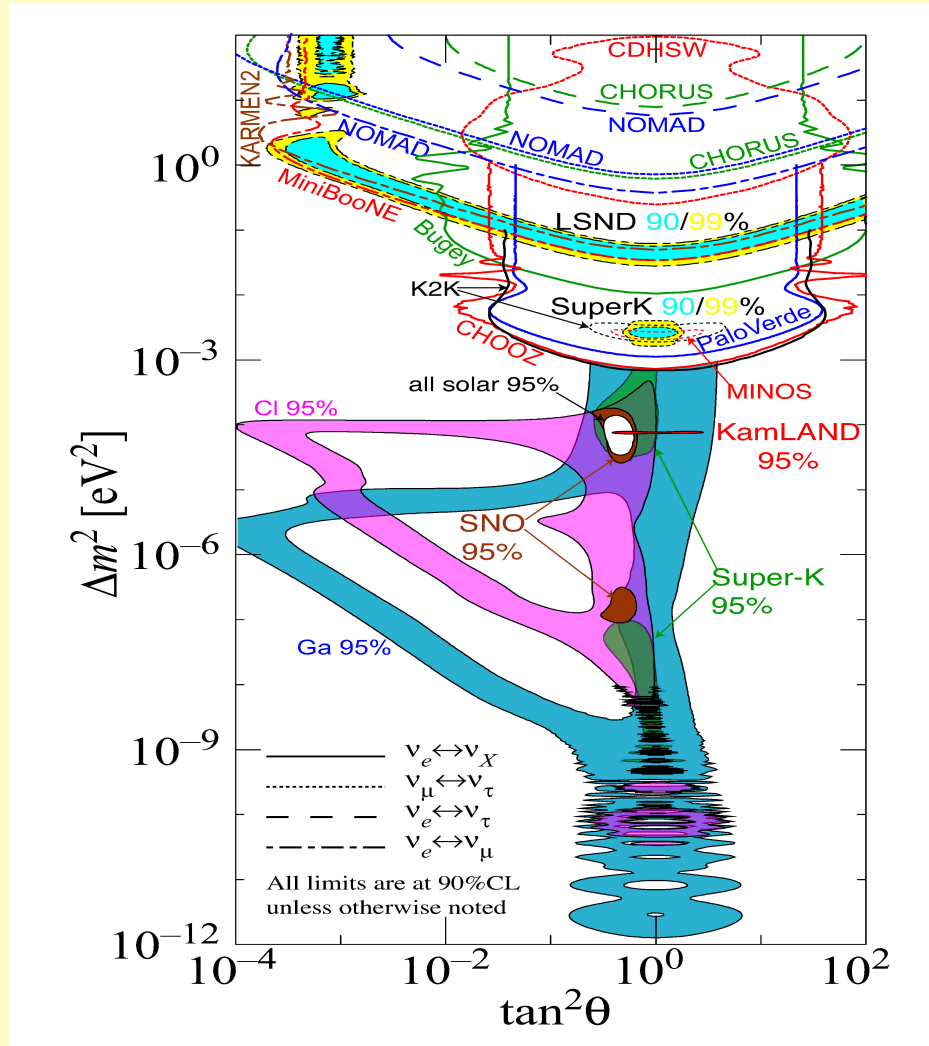
Basically obsolete

In general, better to use  
(preserve octant-symmetry)

$\log \tan^2 \theta$   
or  $\sin^2 \theta$

(Note: Octant symmetry broken by  $3\nu$  and/or matter effects)

# Octant (a)symmetric contours:

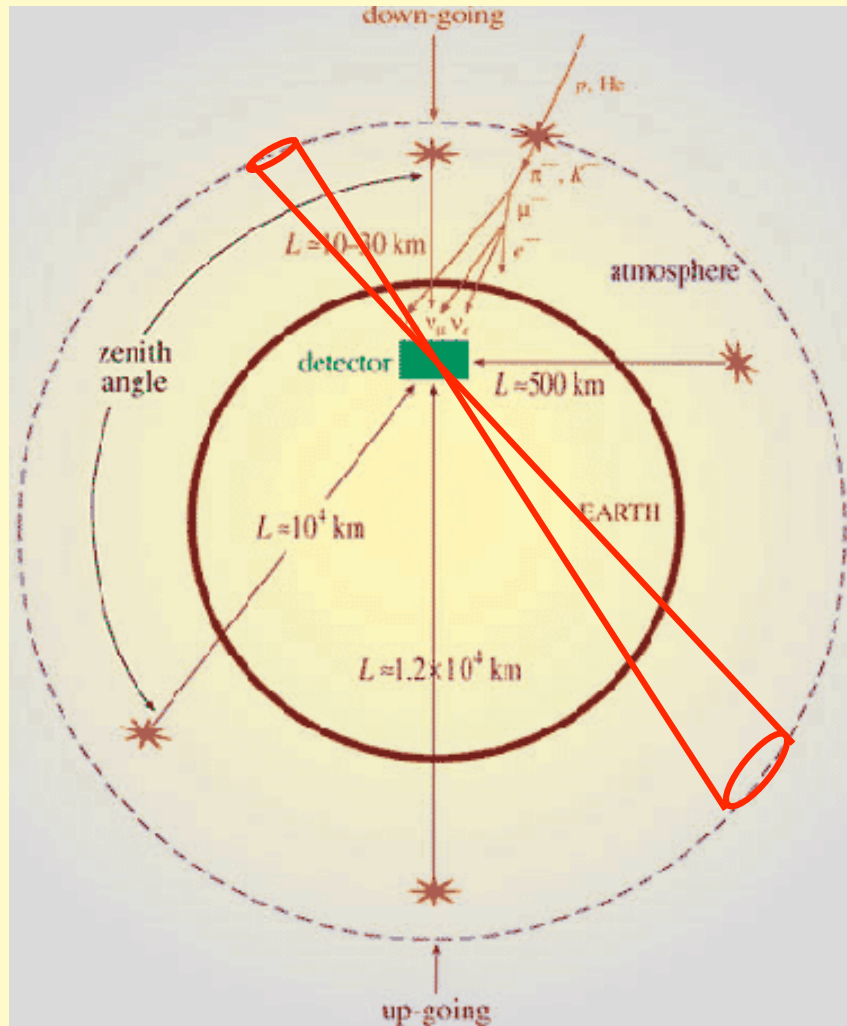


[Particle Data Group 2008]

## Observation of "effective $2\nu$ " oscillations of atmospheric $\nu$ 's

Cosmic rays hitting the atmosphere can generate secondary (anti)neutrinos with electron and muon flavor via meson decays.

Energies:  $E \sim 0.1 - 100 \text{ GeV}$ . Pathlengths:  $L \sim 10 - 10000 \text{ km}$



Same  $\nu$  flux expected from opposite solid angles (up-down symmetry)

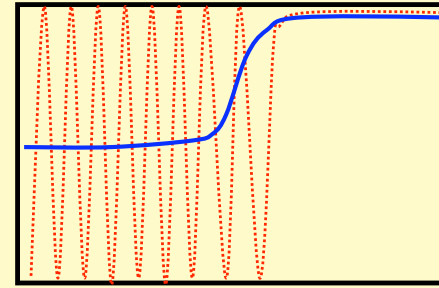
[Flux dilution ( $\sim 1/r^2$ ) is compensated by larger production surface ( $\sim r^2$ )]

Should be reflected in symmetry of event zenith spectra, if energy & angle can be reconstructed well enough



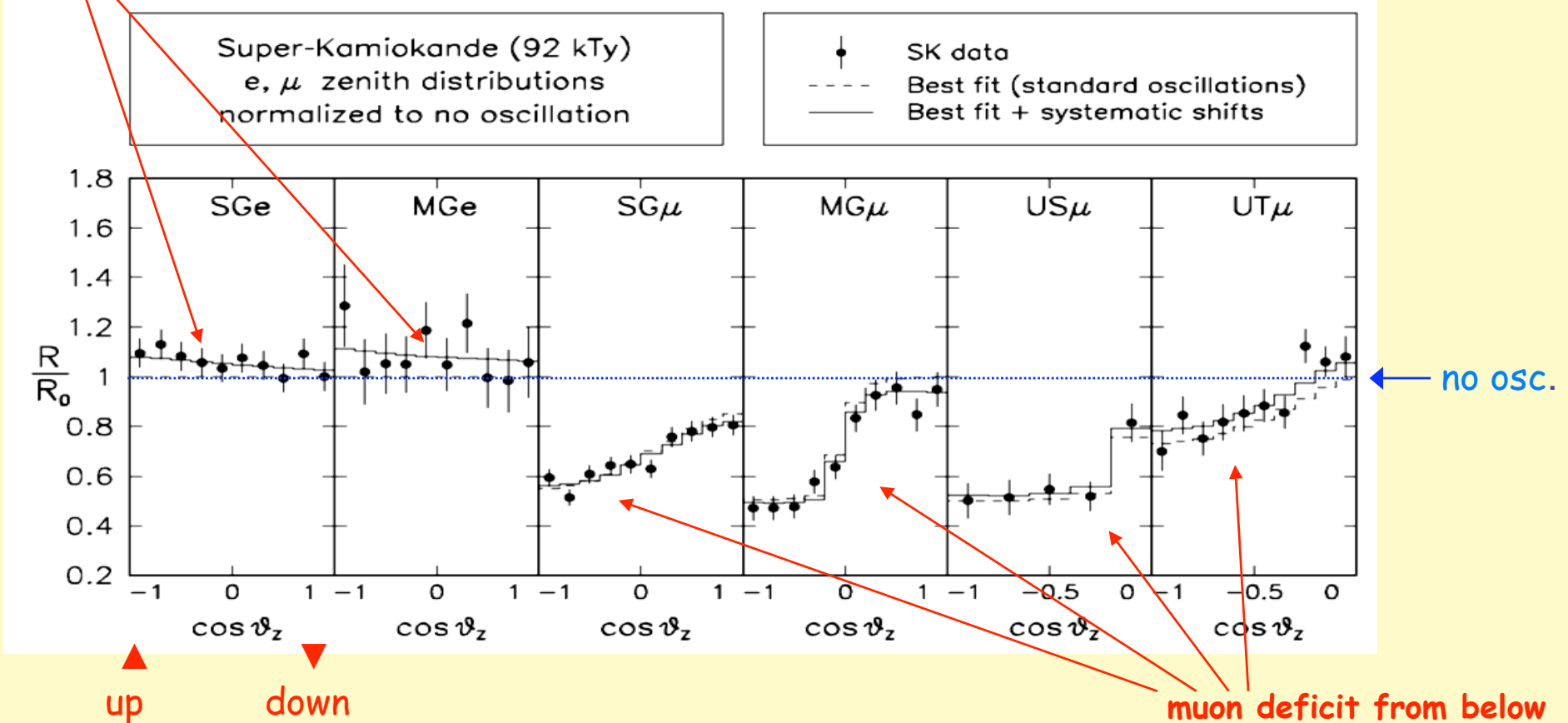
# Atmospheric neutrinos: Super-Kamiokande

<b>S<sub>Ge</sub></b>	Sub-GeV electrons
<b>M<sub>Ge</sub></b>	Multi-GeV electrons
<b>S<sub>Gμ</sub></b>	Sub-GeV muons
<b>M<sub>Gμ</sub></b>	Multi-GeV muons
<b>U<sub>Sμ</sub></b>	Upward Stopping muons
<b>U<sub>Tμ</sub></b>	Upward Through-going muons



$\cos\theta_z$

electrons ~OK



Observations over several decades in L/E:

$\nu_e$  induced events:  $\sim$  as expected

$\nu_\mu$  induced events: disappearance from below

Interpretation in terms of oscillations:

Channel  $\nu_\mu \rightarrow \nu_e$ ? No (or subdominant)

Channel  $\nu_\mu \rightarrow \nu_\tau$ ? Yes (dominant)

2 $\nu$ -like approximation works well...

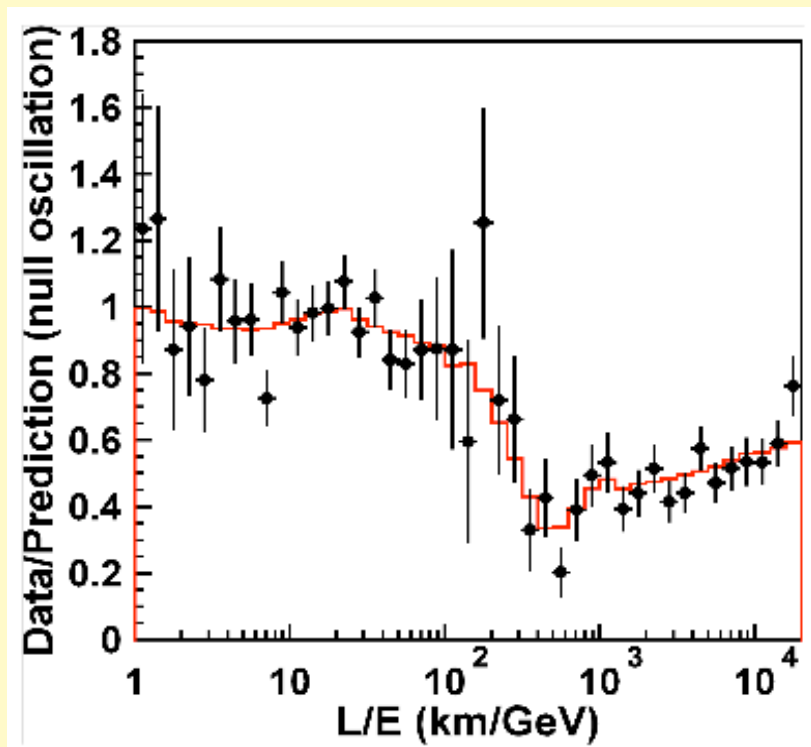
$$P_{\mu\tau} = \sin^2(2\theta) \sin^2(\Delta m^2 L / 4E_\nu)$$

[In this channel, oscillations are  $\sim$ vacuum-like,  
despite the presence of Earth matter]

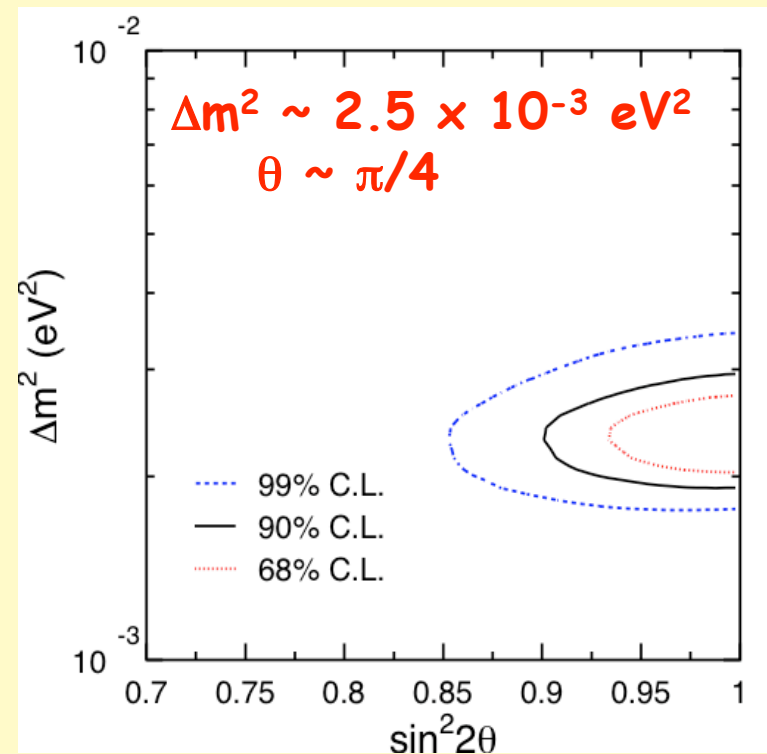
... but where are the "oscillations" ?

## Dedicated L/E analysis to “see” half-period of oscillations

1st oscillation dip still visible despite large L & E smearing

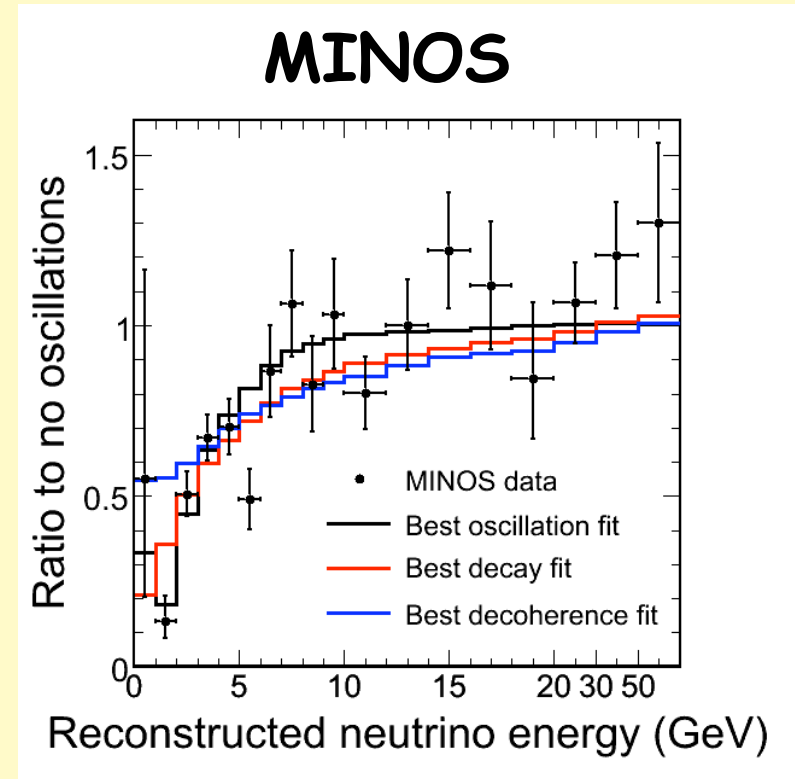
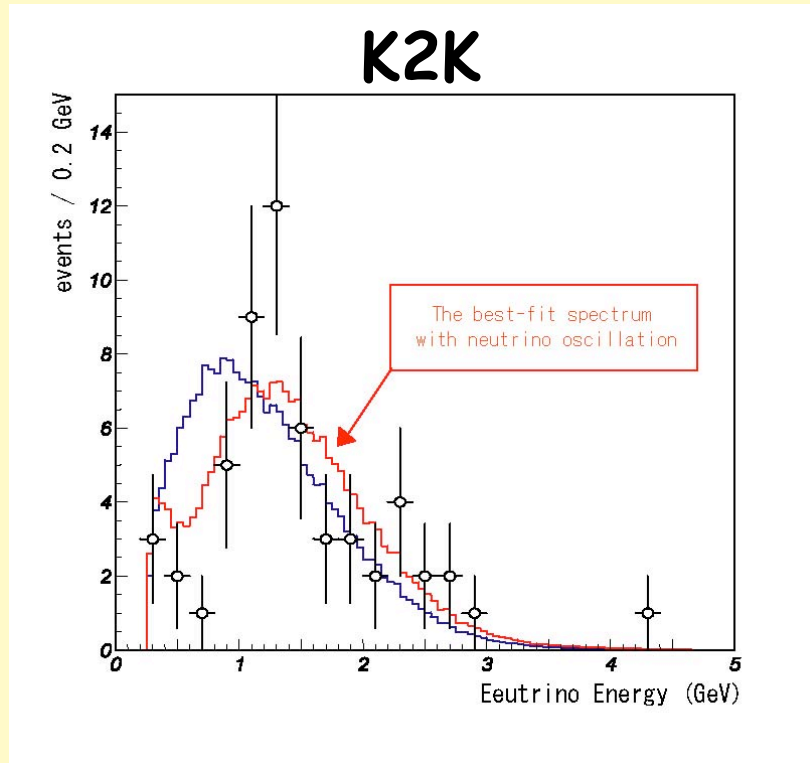


Strong constraints on the parameters ( $\Delta m^2$ ,  $\theta$ )



Same mass/mixing parameters confirmed in disappearance mode ( $\nu_\mu \rightarrow \nu_\mu$ ) by other atmospheric expts (MACRO, Soudan2) and by expts with accelerator beams (K2K, MINOS)

# Accelerator Results (muon disappearance mode)



**1<sup>st</sup> oscillation dip also observed.**

**[Exotic explanations without dip (decay, decoherence) disfavored]**

## Open questions for $\Delta m^2$ -driven $\nu_\mu$ oscillations:

**The quest for hierarchy and octant:** Is the sign of  $\Delta m^2$  positive ("normal hierarchy") or negative ("inverted hierarchy")? Is  $\theta >$  or  $< \pi/4$ ?

**The quest for  $\nu_\tau$  appearance:** We expect dominant  $\nu_\mu \rightarrow \nu_\tau$  transitions, but haven't seen the  $\tau$  flavor directly - the hunt is going on with the CNGS beam. See talks by L. Stanco, A. Guglielmi

**The quest for  $\nu_e$  appearance:** We haven't seen  $\nu_\mu \rightarrow \nu_e$  transitions; are they absent or just suppressed? This is a crucial problem for its implications on leptonic CP violation. See later, & talk by M. Mezzetto

**The quest for sterile neutrinos:** Besides the known neutrinos  $\nu_{e\mu\tau,L}$  (LH, gauge doublets) there might be new "sterile" states  $\nu_{s,R}$  (RH, gauge singlets) leading to further disappearance  $\nu_{\mu L} \rightarrow (\nu_{s,R})^c$   
See talk by C. Giunti

Useful to rephrase some of these questions in  $3\nu$  language  $\rightarrow$

## 3ν, 1st step: one dominant mass splitting

- 3 flavor and mass states:

$$(\nu_e, \nu_\mu, \nu_\tau)^T = U (\nu_1, \nu_2, \nu_3)^T$$

Unitary matrix  $U$  depends on: 3 rotation angles  $\theta_{ij}$  + 1 complex  $CP$  phase.  
Conventionally, same ordering of the CKM quark matrix used for neutrinos:

$$U = \begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix} \begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix} \begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}$$

where  $c_{ij} = \cos(\theta_{ij})$  etc.

[For antineutrinos:  $U \rightarrow U^*$ ]



- For the 3 masses, let's assume for the moment a single dominant splitting:

$$m_1 \simeq m_2 \quad \text{and} \quad \Delta m^2 = |m_3^2 - m_{1,2}^2|$$

which is a reasonable approx. for all experiments where  $\Delta m^2 L / 4E \sim O(1)$  namely, atmospheric, long-baseline accelerator, short-baseline reactor expts.

Then, the vacuum oscillation probabilities are generalized as (2v  $\rightarrow$  3v):

$$P_{\alpha\beta} \simeq \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E} \right) \quad \longrightarrow \quad P_{\alpha\beta} \simeq 4|U_{\alpha 3}|^2 |U_{\beta 3}|^2 \sin^2 \left( \frac{\Delta m^2 L}{4E} \right)$$

$$P_{\alpha\alpha} \simeq 1 - \sin^2 2\theta \sin^2 \left( \frac{\Delta m^2 L}{4E} \right) \quad \longrightarrow \quad P_{\alpha\alpha} \simeq 1 - 4|U_{\alpha 3}|^2 (1 - |U_{\alpha 3}|^2) \sin^2 \left( \frac{\Delta m^2 L}{4E} \right)$$

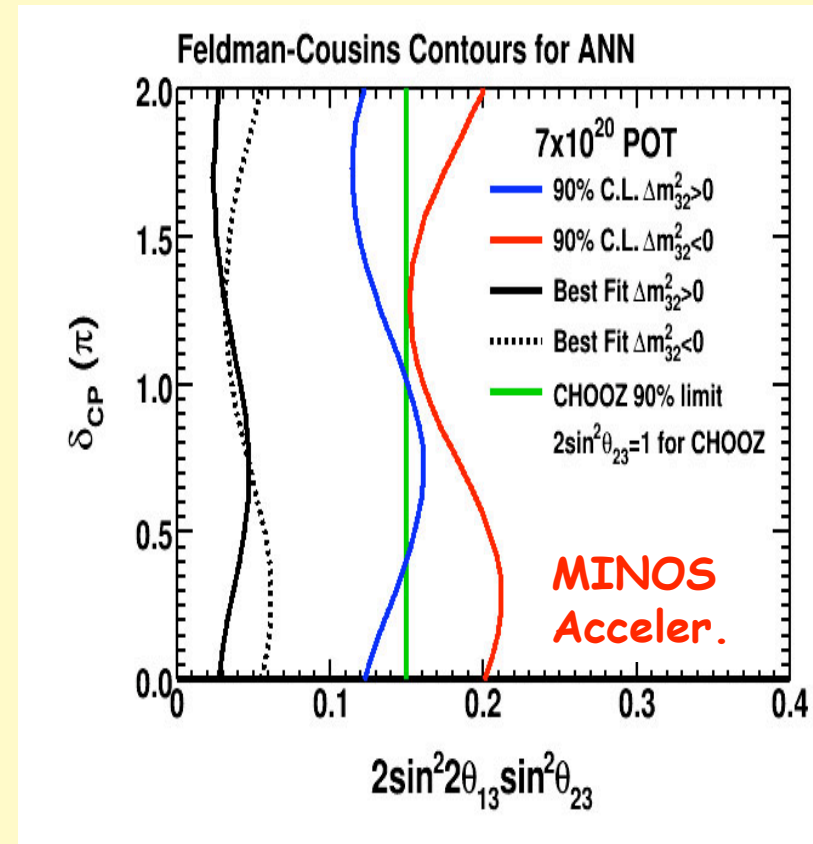
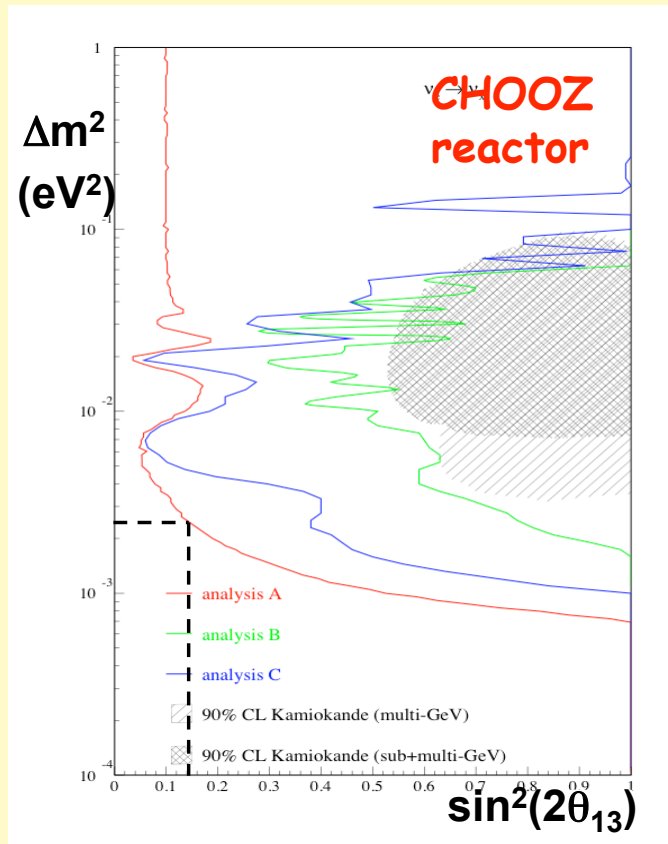
The amplitudes now differ in different oscillation channels, yet they do not depend on the hierarchy or the CP phase.

Also, they do not depend on  $\theta_{12}$ , due to the assumed degeneracy  $m_1 \simeq m_2$

In such notation, the previous " $\nu_\mu \rightarrow \nu_\tau$ " mixing angle is  $\theta_{23} \sim \pi/4$ , while  $\theta_{13}$  modulates the oscillation amplitude in the  $\nu_e \rightarrow \nu_e$  and  $\nu_\mu \rightarrow \nu_e$  channels where, unfortunately, no signal has been found so far...

$$P_{ee} = 1 - \sin^2(2\theta_{13}) \sin^2(\Delta m^2 L / 4E_\nu)$$

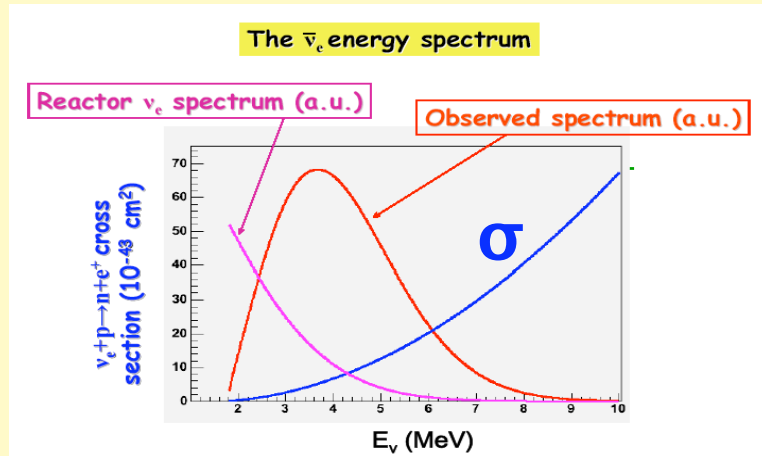
$$P_{\mu e} = \sin^2\theta_{23} \sin^2(2\theta_{13}) \sin^2(\Delta m^2 L / 4E_\nu)$$



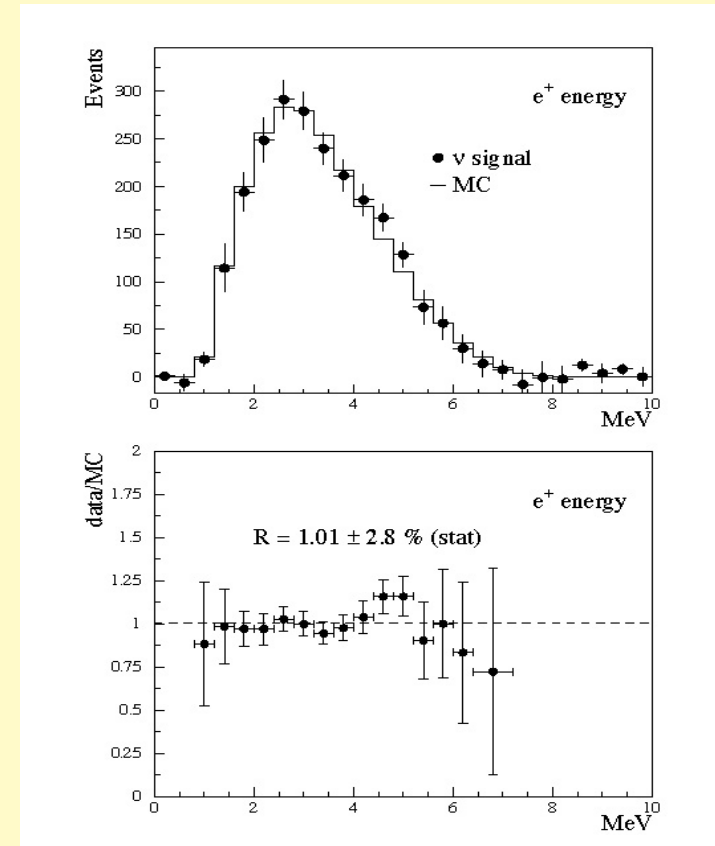
World data consistent with  $\sin^2\theta_{13} < \text{few } \%$ .

# - More about CHOOZ results ( $L \sim 1\text{km}$ ) -

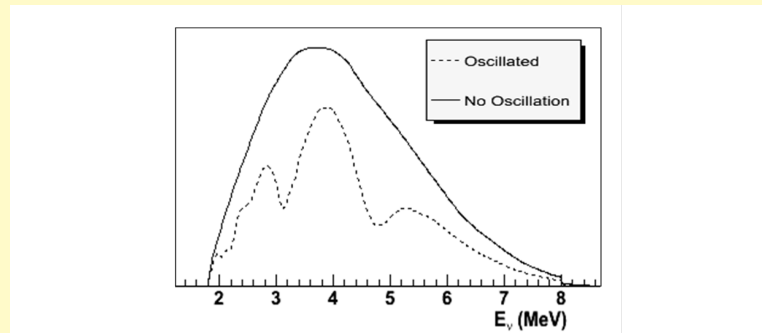
Expected spectrum (no oscill.):



Data: no oscillations within few % error



With oscillations (qualitative):



Future reactor expts: Reduce syst's with near/far detectors.

But: Why hope for  $\theta_{13} > 0$  after all?

## 3ν, 2nd step: two mass splittings

We have seen that atmospheric (and long-baseline accelerator) experiments have established the mass splitting of  $\nu_3$  with respect to  $\nu_{1,2}$ , with oscillation parameters:

$$\Delta m^2 = |m_3^2 - m_{1,2}^2| \simeq 2.5 \times 10^{-3} \text{ eV}^2$$

$$\sin^2 \theta_{23} \simeq 0.5$$

We shall see tomorrow that solar and long-baseline reactors, sensitive to much larger  $L/E$ , have established the splitting between  $\nu_1$  and  $\nu_2$  with oscillation parameters:

$$\delta m^2 = m_2^2 - m_1^2 \simeq 8 \times 10^{-5} \text{ eV}^2$$

$$\sin^2 \theta_{12} \simeq 0.3$$

**This opens the door to leptonic CP violation, iff  $\theta_{13} > 0$  !**

In a full 3ν scenario, a CP violating difference may arise between neutrino and antineutrino oscillation probabilities,

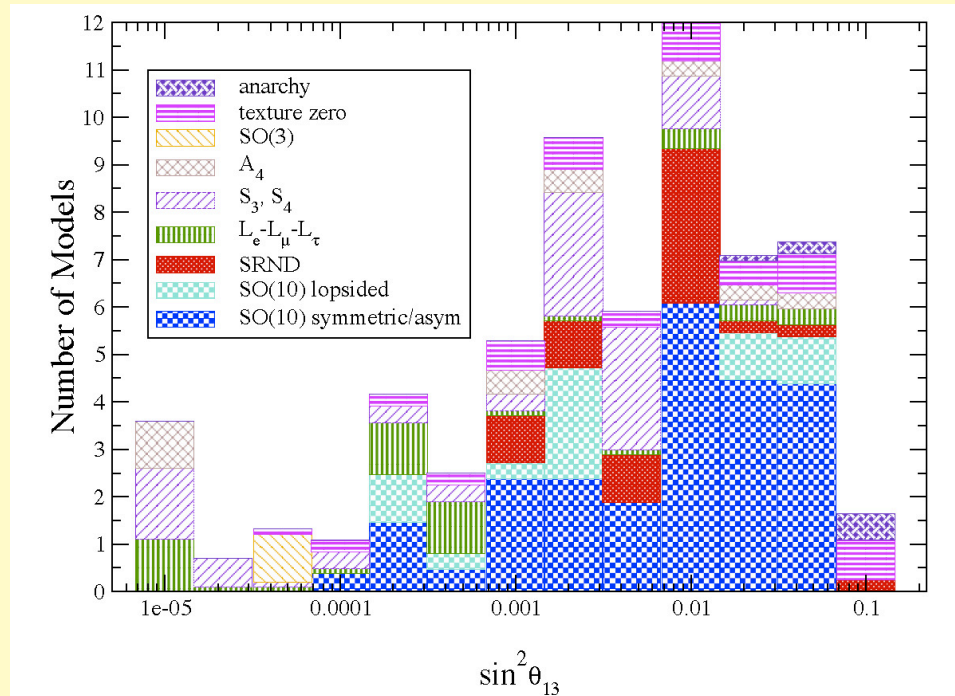
$$P_{\alpha\beta}(\nu) - P_{\alpha\beta}(\bar{\nu}) = 2 \sin 2\theta_{12} \sin 2\theta_{23} \sin 2\theta_{13} \cos \theta_{13} \sin \delta \\ \times \sin \left( \frac{\Delta m^2 - \frac{\delta m^2}{2}}{4E} L \right) \sin \left( \frac{\Delta m^2 + \frac{\delta m^2}{2}}{4E} L \right) \sin \left( \frac{\delta m^2}{4E} L \right)$$

provided that:

- $\sin 2\theta_{13}$  is nonzero
- $\sin \delta$  is nonzero
- the oscillation phases are neither too small nor too large

Hunt for  $\theta_{13}$  crucial in current neutrino research

Also:  $\theta_{13}$  very important to restrict theoretical models for  $\nu$  masses

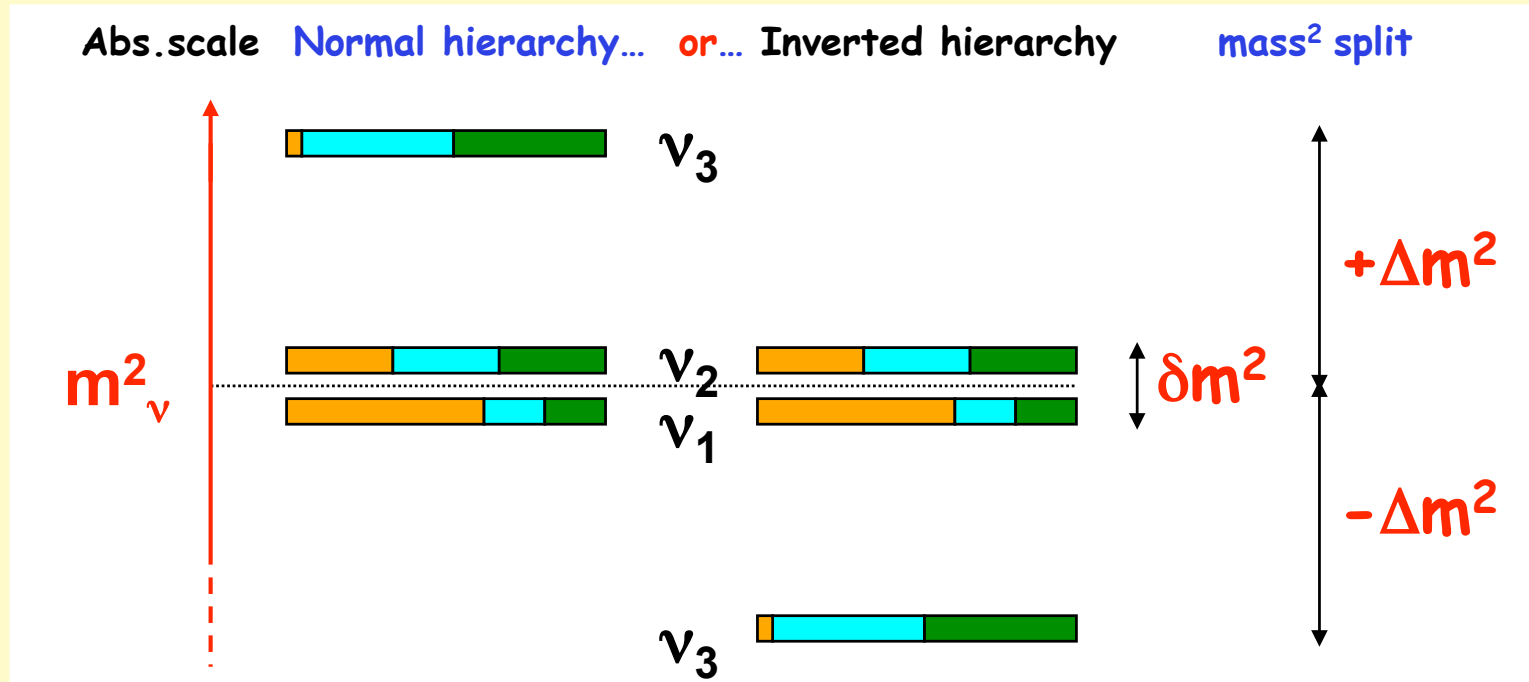


E.g.: CH Albright, 2008, “distribution” of published predictions

See talk by F. Feruglio

# 3ν mass-mixing overview in just one slide

(here, with 1 digit accuracy). Flavors =  $e \mu \tau$



$$\delta m^2 \sim 8 \times 10^{-5} \text{ eV}^2$$

$$\Delta m^2 \sim 3 \times 10^{-3} \text{ eV}^2$$

$$\sin^2 \theta_{12} \sim 0.3$$

$$\sin^2 \theta_{23} \sim 0.5$$

$$m_\nu < O(1) \text{ eV}$$

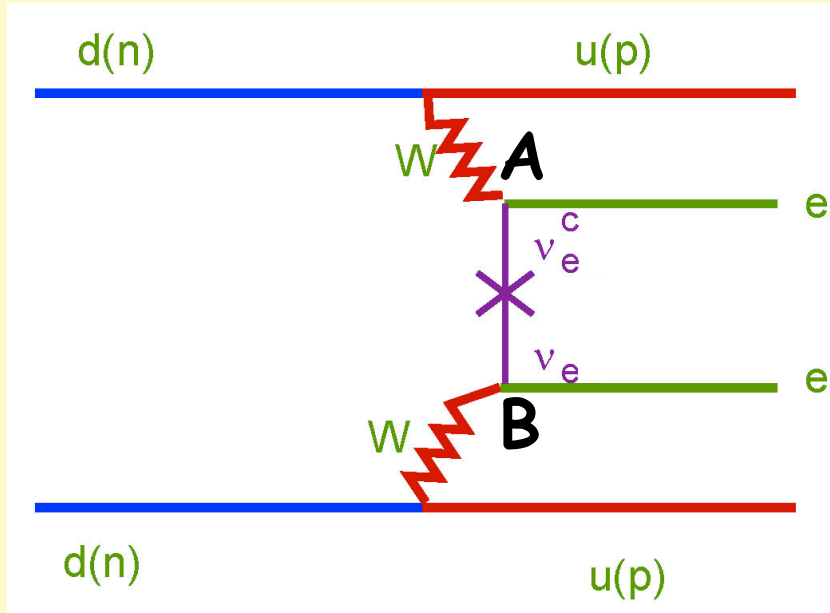
$$\sin^2 \theta_{13} < \text{few}\%$$

sign( $\pm \Delta m^2$ ) unknown

$\delta$  (CP) unknown

## Implications of $3\nu$ mixing for $0\nu\beta\beta$ decay

For each mass state  $\nu_i$ ,  $0\nu\beta\beta$  amplitude proportional to:



- ... mixing of  $\nu_e$  with  $\nu_i$
- ... mass of  $\nu_i$
- ... mixing of  $\nu_i$  with  $\nu_e$
- (times an unknown  $\nu_i$  phase)

Summing up for three massive neutrinos:

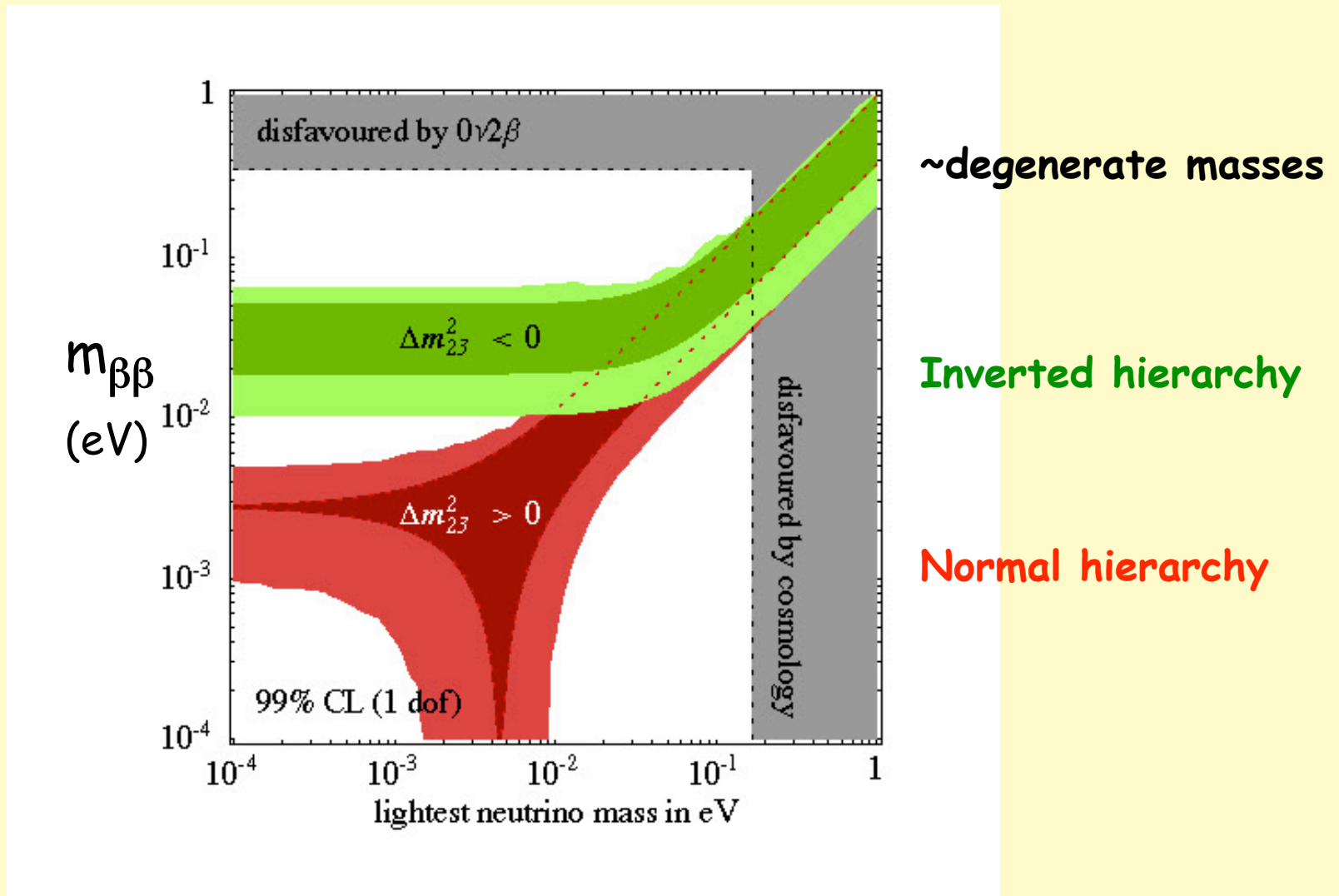
**Amplitude ~ "effective Majorana mass"**

$$m_{\beta\beta} = \left| c_{13}^2 c_{12}^2 m_1 + c_{13}^2 s_{12}^2 m_2 e^{i\phi_2} + s_{13}^2 m_3 e^{i\phi_3} \right|$$

[complex linear combination of masses;  $c_{ij} = \cos \theta_{ij}$  etc.]



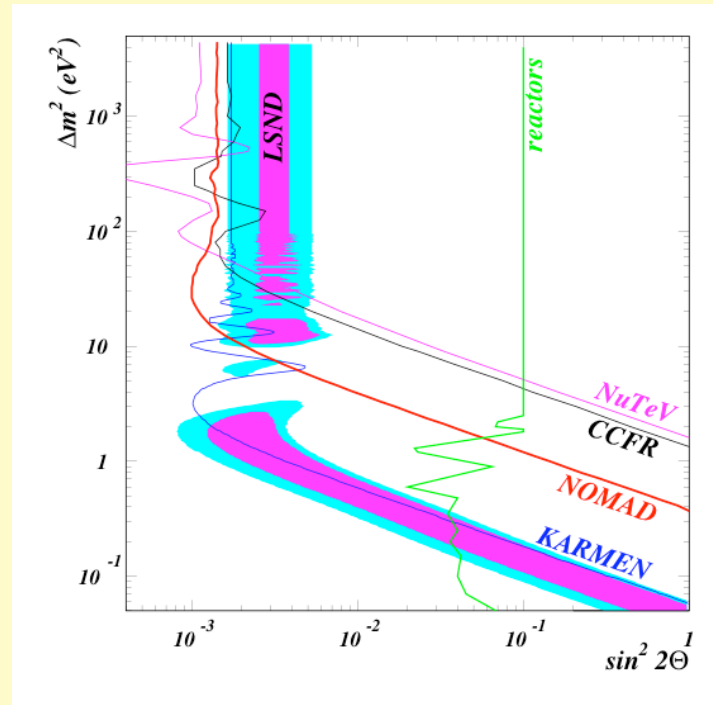
# Typical plot of $m_{\beta\beta}$ versus lightest neutrino mass:



$$m_{\beta\beta} = \left| c_{13}^2 c_{12}^2 m_1 + c_{13}^2 s_{12}^2 m_2 e^{i\phi_2} + s_{13}^2 m_3 e^{i\phi_3} \right|$$

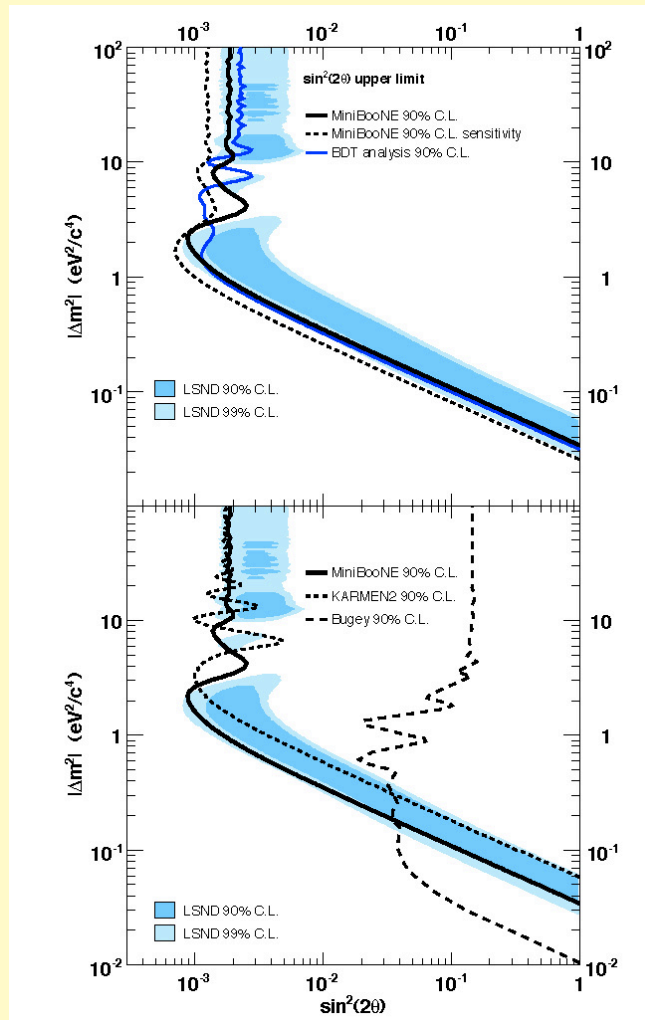
# Short baseline accelerator expts: Beyond 3 neutrinos?

The LSND experiment found a signal of possible  $\nu_\mu \rightarrow \nu_e$  oscillations at a relatively high  $\Delta M^2$  scale of  $O(0.1-1) \text{ eV}^2$



Large literature on attempts to reconcile LSND with other data, by using new (sterile) neutrino states and/or new neutrino interactions. No satisfactory model emerged so far. Moreover...

... simplest LSND "oscillations" excluded by a dedicated test experiment, MiniBoone:



But, the MiniBoone data have some new, unexplained anomalies at low energy!

So the LSND/MiniBoone saga may have not yet ended ...

Sterile neutrinos and new physics at work? See talk by C. Giunti



# RECAP and end of LECTURE I



Poster of the Neutrino Oscillation Workshop 2004 (NOW 2004, Otranto, Italy)

# **HOMEWORK**



# Solution 1

- If  $\nu$ 's are Dirac, then  $\nu_e \neq \bar{\nu}_e$ , and one can attach a leptonic number to the doublets  $(\nu_e, e^-)$  and  $(\bar{\nu}_e, e^+)$ , which is conserved in the observed reactions ( $\Delta L = 0$ ) and would be violated in the other two ( $\Delta L = 2$ ).
- If  $\nu$ 's are Majorana, then  $\nu_e = \bar{\nu}_e$ , and we are just naming:
  - " $\nu_e$ " = LH component of  $\nu$  state
  - " $\bar{\nu}_e$ " = RH component of  $\nu$  state

The initial " $\nu_e$ " is LH, being produced in a weak ( $\beta^+$ ) decay. While propagating, it remains dominantly LH, but can develop a small RH component (" $\bar{\nu}_e$ ") at  $\mathcal{O}(m/E)$ . Then also the reaction  $\bar{\nu}_e + n \rightarrow p + e^-$  can take place in principle, but is so suppressed to be practically unobservable - lepton number violation ( $\Delta L = 2$ ) is allowed in principle, but suppressed at  $\mathcal{O}(m/E)$  in practice.

## Solution 2

- Mass basis  $\begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$  and flavor basis  $\begin{pmatrix} \nu_\alpha \\ \nu_\beta \end{pmatrix}$  are related by:

$$\begin{pmatrix} \nu_\alpha \\ \nu_\beta \end{pmatrix} = U \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix} = \begin{pmatrix} \cos\theta & \sin\theta \\ -\sin\theta & \cos\theta \end{pmatrix} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix} \quad \text{with } \Delta m^2 = m_2^2 - m_1^2$$

- Evolution equation in mass basis (mb):

$$i \frac{d}{dt} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix} = \mathcal{H}_{mb} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}$$

← Schrödinger eq. in natural units ( $\hbar = c = 1$ )

where the Hamiltonian is simply

$$\mathcal{H}_{mb} = \begin{pmatrix} E_1 & 0 \\ 0 & E_2 \end{pmatrix} \approx \begin{pmatrix} p + \frac{m_1^2}{2E} & 0 \\ 0 & p + \frac{m_2^2}{2E} \end{pmatrix} = \underbrace{\left( p + \frac{m_1^2 + m_2^2}{4E} \right)}_{\propto \mathbb{1}} + \underbrace{\begin{pmatrix} -\frac{\Delta m^2}{4E} & 0 \\ 0 & +\frac{\Delta m^2}{4E} \end{pmatrix}}_{\text{Traces}}$$

Final results do not depend on the part proportional to  $\mathbb{1}$  - check it.

(Reason: it gives an overall phase which disappears in observable real quantities). So we take:

$$\mathcal{H}_{mb} = \frac{\Delta m^2}{4E} \begin{pmatrix} -1 & 0 \\ 0 & +1 \end{pmatrix}$$

## Solution 2 (ctd)

- Evolution operator in mass basis:

$$\begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}_t = S_{\text{mb}} \begin{pmatrix} \nu_1 \\ \nu_2 \end{pmatrix}_0 \quad \text{where}$$

$$S_{\text{mb}} = e^{-iH_{\text{mb}}t} \simeq e^{-iH_{\text{mb}}x} = \begin{pmatrix} e^{i\frac{\Delta m^2}{4E}x} & 0 \\ 0 & e^{-i\frac{\Delta m^2}{4E}x} \end{pmatrix}$$

←  $x \simeq t$  for ultrarelativistic neutrinos

- Evolution operator in flavor basis (fb):

$$\begin{aligned} S_{\text{fb}} &= U S_{\text{mb}} U^T \\ &= \cos\left(\frac{\Delta m^2 x}{4E}\right) \begin{pmatrix} 1 & 0 \\ 0 & 1 \end{pmatrix} - i \sin\left(\frac{\Delta m^2 x}{4E}\right) \begin{pmatrix} -\cos 2\theta & \sin 2\theta \\ \sin 2\theta & \cos 2\theta \end{pmatrix} \end{aligned}$$

- Amplitudes for flavor transitions:

$$\begin{pmatrix} \nu_\alpha \\ \nu_\beta \end{pmatrix}_t = S_{\text{fb}} \begin{pmatrix} \nu_\alpha \\ \nu_\beta \end{pmatrix}_0 \quad \rightarrow \quad \text{Off-diagonal elements of } S_{\text{fb}} \text{ give} \\ \text{amplitudes for } \nu_\alpha \rightarrow \nu_\beta \text{ and } \nu_\beta \rightarrow \nu_\alpha$$



## Solution 2 (ctd)

- Probability of flavor transition is the square modulus of the amplitude:

$$\begin{aligned} P(\nu_\alpha \rightarrow \nu_\beta) &= P(\nu_\beta \rightarrow \nu_\alpha) = \left| -i \sin 2\theta \sin\left(\frac{\Delta m^2 x}{4E}\right) \right|^2 \\ &= \sin^2 2\theta \sin^2\left(\frac{\Delta m^2 x}{4E}\right) \quad (x=L \text{ in the lecture}). \end{aligned}$$

- The diagonal elements of the evolution operator would give the "flavor survival" amplitude. Check that

$$1 - P(\nu_\alpha \rightarrow \nu_\beta) = P(\nu_\alpha \rightarrow \nu_\alpha) = P(\nu_\beta \rightarrow \nu_\beta).$$

$$\rightarrow P(\nu_\alpha \rightarrow \nu_\alpha) = 1 - \sin^2 2\theta \sin^2\left(\frac{\Delta m^2 x}{4E}\right)$$

## Solution 3

Oscillations depend only on the difference of phases, and thus of neutrino energies. Indeed, the results do not change by an overall shift of the Hamiltonian:

$$H \rightarrow H + \text{const.} \mathbb{1}$$

Since the zero-point energy is irrelevant in this context, the absolute neutrino mass scale ~~is~~ is unobservable (in oscillation searches).

## Solution 4

$$\frac{\hbar c}{197.327 \text{ MeV} \cdot \text{fm}} = 1 \text{ in natural units.}$$

$$\text{Therefore: } 1 \text{ MeV} \cdot 1 \text{ m} = 5.0677 \times 10^{12}$$

$$\begin{aligned} \text{Then: } \frac{\Delta m^2 L}{4E} &= \frac{1}{4} \left( \frac{\Delta m^2}{\text{eV}^2} \text{ eV}^2 \right) \left( \frac{L}{\text{m}} \cdot \text{m} \right) \left( \frac{\text{MeV}}{E} \cdot \frac{1}{\text{MeV}} \right) \\ &= \frac{1}{4} \left( \frac{1 \text{ eV}^2 \cdot 1 \text{ m}}{1 \text{ MeV}} \right) \left( \frac{\Delta m^2}{\text{eV}^2} \right) \left( \frac{L}{\text{m}} \right) \left( \frac{\text{MeV}}{E} \right) \end{aligned}$$

$$\frac{1}{4} \frac{\text{eV}^2 \text{ m}}{\text{MeV}} = \frac{1}{4} \times 10^{-12} \frac{\text{MeV}^2 \cdot 1 \text{ m}}{1 \text{ MeV}} = \frac{10^{-12}}{4} (\text{MeV} \cdot \text{m}) = 0.25 \times 10^{-12} \times 5.0677 \times 10^{12} = 1.267$$

$$\frac{\Delta m^2 L}{4E} = 1.267 \left( \frac{\Delta m^2}{\text{eV}^2} \right) \left( \frac{L}{\text{m}} \right) \left( \frac{\text{MeV}}{E} \right) = 1.267 \left( \frac{\Delta m^2}{\text{eV}^2} \right) \left( \frac{L}{\text{km}} \right) \left( \frac{\text{GeV}}{E} \right)$$

# Appendix on Majorana mass terms

## Dirac and Majorana mass terms (1 family)

- Dirac mass terms are of the form  $m\bar{\psi}\psi$  (4 dof  $\psi$ )
- Majorana " " " " "  $\frac{1}{2}m\bar{\psi}\psi$  (2 dof  $\psi$ )

Three possibilities:

$$\text{Dirac} : \psi = \psi_L + \psi_R \rightarrow \bar{\psi}\psi = \bar{\psi}_L\psi_R + \bar{\psi}_R\psi_L$$

$$\text{Majorana (L)} : \psi = \psi_L + \psi_L^c \rightarrow \bar{\psi}\psi = \bar{\psi}_L\psi_L^c + \bar{\psi}_L^c\psi_L$$

$$\text{Majorana (R)} : \psi = \psi_R + \psi_R^c \rightarrow \bar{\psi}\psi = \bar{\psi}_R\psi_R^c + \bar{\psi}_R^c\psi_R$$

Most general mass term for one neutrino family:

$$m_D (\bar{\psi}_L\psi_R + \bar{\psi}_R\psi_L) + \frac{1}{2}m_L (\bar{\psi}_L\psi_L^c + \bar{\psi}_L^c\psi_L) + \frac{1}{2}m_R (\bar{\psi}_R\psi_R^c + \bar{\psi}_R^c\psi_R)$$

[Last two terms absent for charged fermions.]



- Previous mass term can be rewritten as:

$$\frac{1}{2} [\bar{\Psi}_L + \bar{\Psi}_L^c, \bar{\Psi}_R + \bar{\Psi}_R^c] \begin{bmatrix} m_L & m_D \\ m_D & m_R \end{bmatrix} \begin{bmatrix} \Psi_L + \Psi_L^c \\ \Psi_R + \Psi_R^c \end{bmatrix}$$

- Diagonalization provides fields with definite masses.  
(If mass  $< 0$ , redefine field  $\psi \rightarrow \gamma_5 \psi$  so that  $m \rightarrow -m$ )
- Since the basis fields  $(\bar{\Psi}_L + \bar{\Psi}_L^c)$  and  $(\bar{\Psi}_R + \bar{\Psi}_R^c)$  are Majorana, diagonalization will generally produce mass eigenvectors which are also Majorana

Diagonalize  $M = \begin{bmatrix} m_L & m_D \\ m_D & m_R \end{bmatrix}$

$$M = \begin{bmatrix} m_L & m_D \\ m_D & m_R \end{bmatrix}$$

$$T = \text{Tr } M = m_L + m_R$$

$$D = \det M = m_L m_R - m_D^2$$

$$\text{Eigenvalues: } m_{\pm} = \frac{1}{2} (T \pm \sqrt{T^2 - 4D})$$

$$\text{Diagonalization angle: } \sin 2\theta = \frac{m_D}{\sqrt{T^2 - 4D}}$$

(not a mixing angle!)

$$\cos 2\theta = \frac{m_L - m_R}{\sqrt{T^2 - 4D}}$$

$$\begin{bmatrix} m_+ & 0 \\ 0 & m_- \end{bmatrix} = \begin{bmatrix} c_\theta & s_\theta \\ -s_\theta & c_\theta \end{bmatrix} \begin{bmatrix} m_L & m_D \\ m_D & m_R \end{bmatrix} \begin{bmatrix} c_\theta & -s_\theta \\ s_\theta & c_\theta \end{bmatrix}$$

Eigenvectors

$$\begin{bmatrix} v_1 & v_2 \end{bmatrix} \begin{bmatrix} m_L & m_D \\ m_D & m_R \end{bmatrix} \begin{bmatrix} v_1 \\ v_2 \end{bmatrix} = \begin{bmatrix} v'_1 & v'_2 \end{bmatrix} \begin{bmatrix} m_+ & 0 \\ 0 & m_- \end{bmatrix} \begin{bmatrix} v'_1 \\ v'_2 \end{bmatrix}$$

$$\begin{bmatrix} v'_1 \\ v'_2 \end{bmatrix} = \begin{bmatrix} c_\theta & s_\theta \\ -s_\theta & c_\theta \end{bmatrix} \begin{bmatrix} v_1 \\ v_2 \end{bmatrix}$$

## The see-saw mechanism

Many extensions of the Standard Model predict the existence of singlet neutrinos ( $\nu_R$ ).

E.g., in the 16 representation of  $SO(10)$ :

$$\begin{pmatrix} u_L & u_L & u_L & \nu_L \\ d_L & d_L & d_L & e_L \\ u_R & u_R & u_R & \nu_R \\ d_R & d_R & d_R & e_R \end{pmatrix}$$

→ can get a Majorana mass term  $\sim m_R (\bar{\nu}_R^c \nu_R + \bar{\nu}_R \nu_R^c)$ , where  $m_R$  is presumably a large mass scale characterizing the SM extension.

For  $m \ll M$ , diagonalization of  $\begin{bmatrix} 0 & m \\ m & M \end{bmatrix}$  gives:

Eigenvectors ( $\nu$ fields)	Eigenvalues (masses)
$\nu_{\text{heavy}} \cong (\nu_R + \nu_R^c) + \frac{m}{M} (\nu_L + \nu_L^c)$	$M$
$\nu_{\text{light}} \cong -(\nu_L + \nu_L^c) + \frac{m}{M} (\nu_R + \nu_R^c)$	$(-)\frac{m^2}{M} \ll m$

← see-saw

The light state is active (contains  $\nu_L$ ) and has a very small mass  $\sim m^2/M$

Presumably:  $m \sim \mathcal{O}(m_{\text{quarks}}, m_{\text{leptons}})$   
 $M \sim \mathcal{O}(\Lambda_{\text{beyond SM}})$

**The see-saw mechanism might explain the smallness of neutrino masses**