

# **News on Solar Neutrinos from Super-Kamiokande**

**VIII International Workshop on "Neutrino Telescopes"**

**23 February 1999**

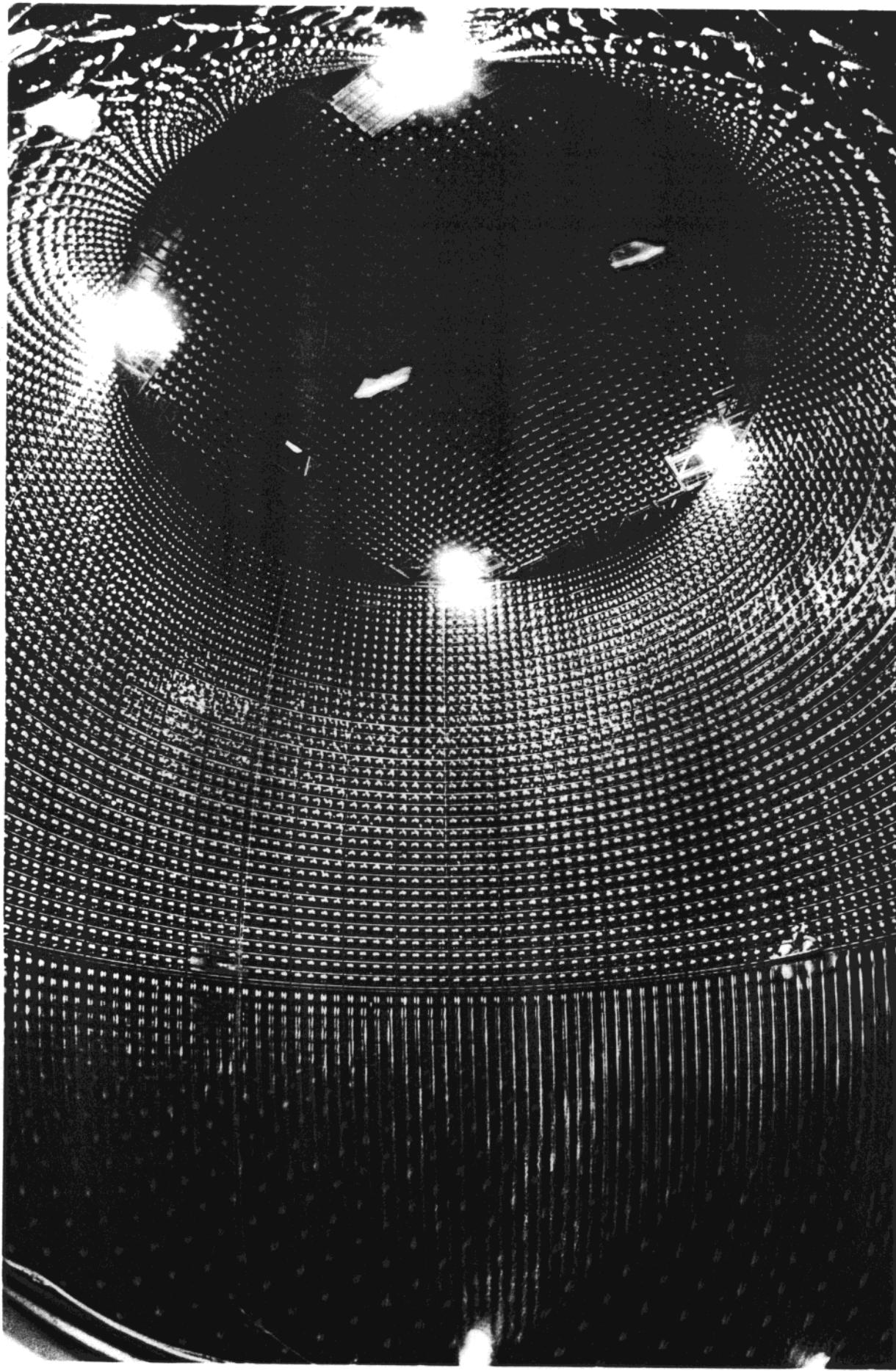
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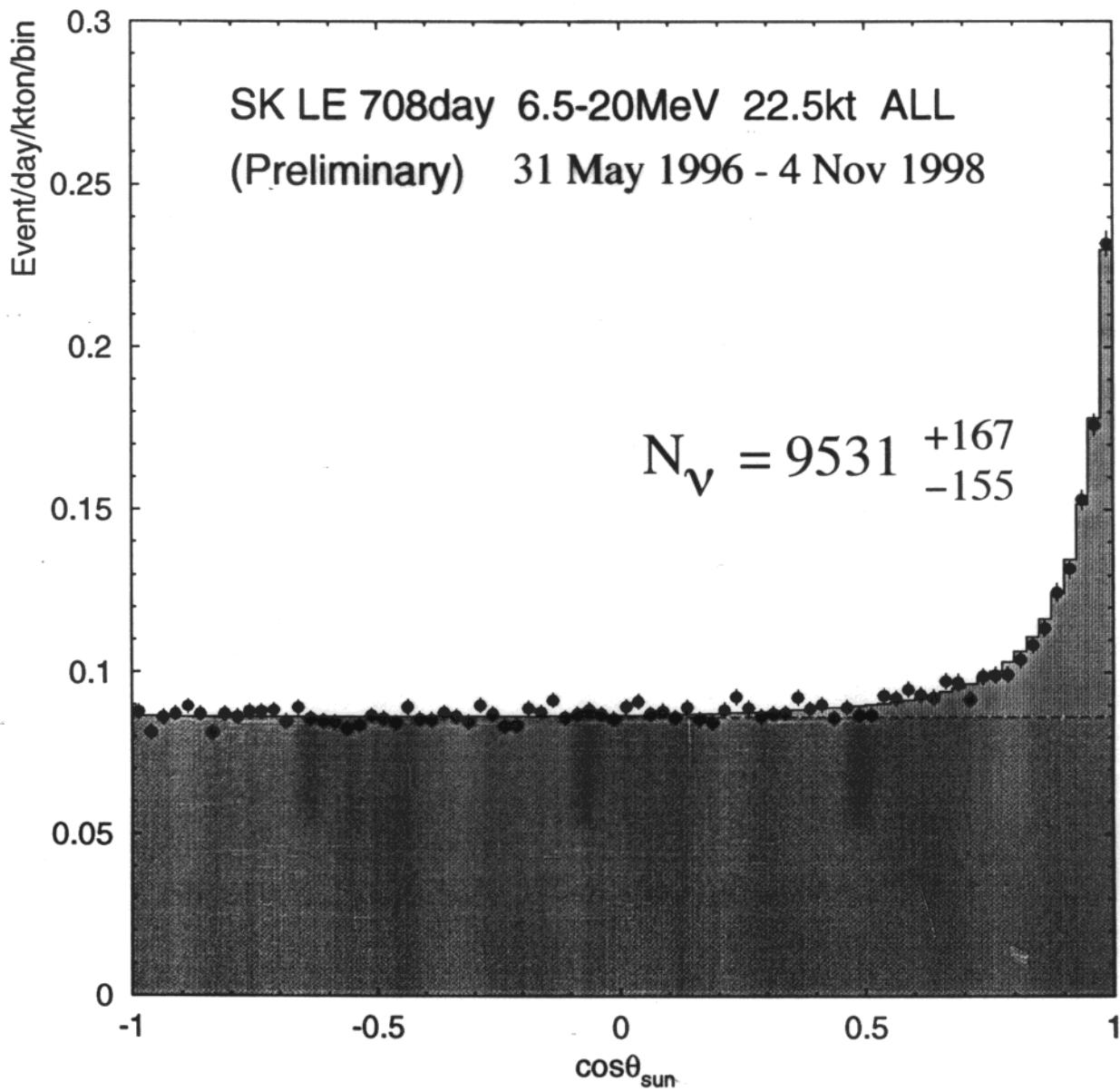
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Jan 1996



# Updated flux result with 708 days of data



$$\phi(^8\text{B}) = 2.419 \begin{array}{l} +0.042 \\ -0.039 \end{array} \text{ (stat)} \begin{array}{l} +0.068 \\ -0.065 \end{array} \text{ (sys)} \times 10^6 \text{ /cm}^2/\text{sec}$$

$$\text{Data/SSM}_{\text{BP98}} = 0.470 \pm 0.008 \pm 0.013$$

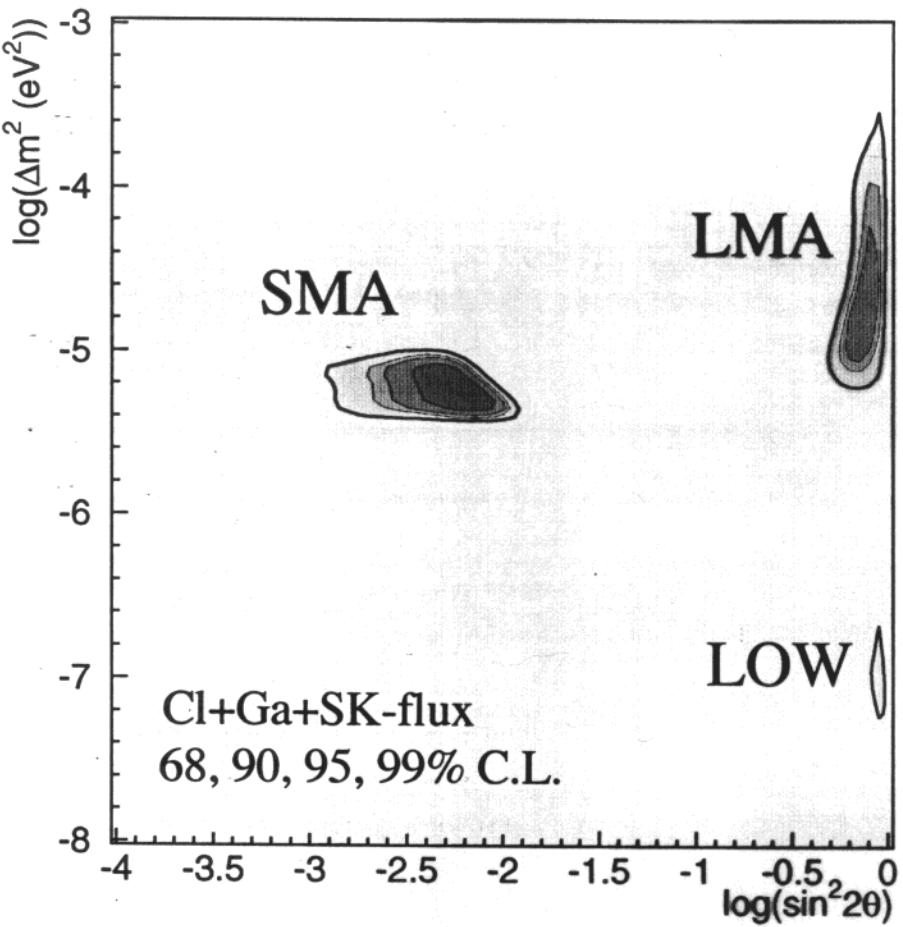
$$\text{Data/SSM}_{\text{BTM98}} = 0.502 \pm 0.008 \pm 0.014$$

# Systematic Errors

for 22.5 kton,  
6.5-20MeV

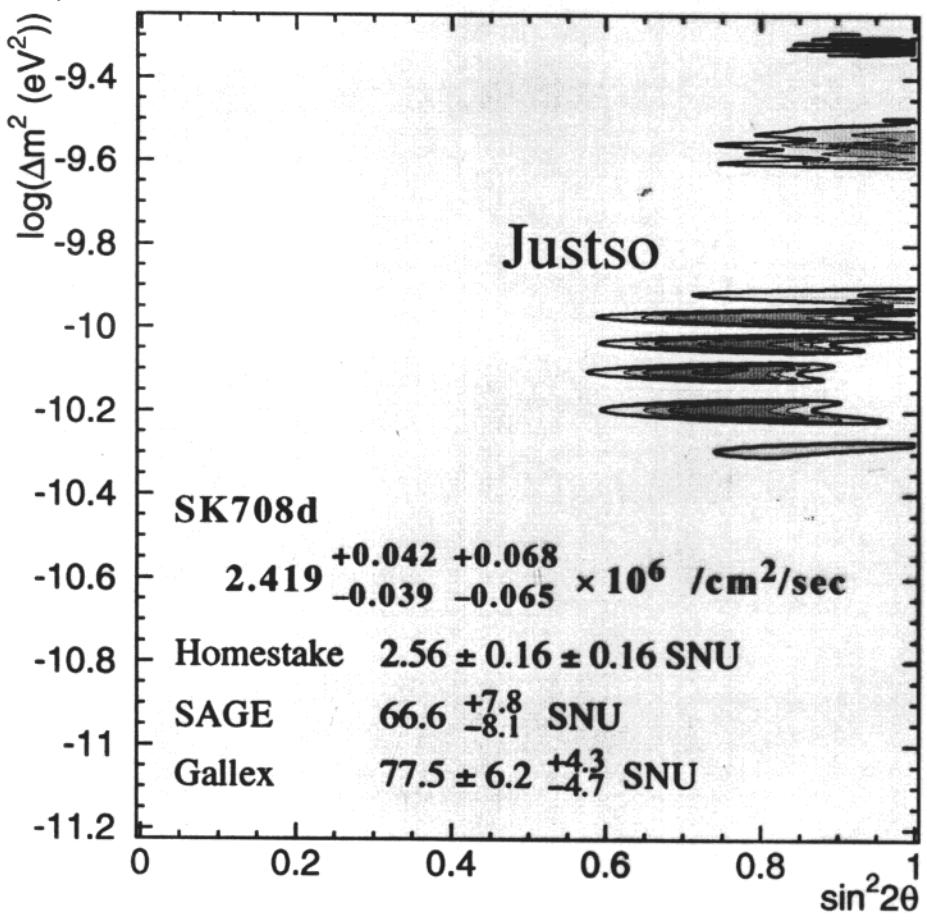
	Flux	Seasonal	Day-Night binned data
<b>E-Scale, Resolution, 8B spectrum</b>	+1.9 -1.8	+1.2 -1.1	+1.2 -1.1
<b>Trigger Efficiency</b>	$\pm 0.2$	$\pm 0.2$	-
<b>Noise event Cut</b>	$\pm 0.7$	-	-
<b>Reduction</b>	$\pm 0.2$	-	$\pm 0.1$
<b>Vertex Shift</b>	$\pm 1.5$	-	-
<b>non-flat B.G.</b>	$\pm 0.1$	$\pm 0.1$	$\pm 0.4$
<b>Direction</b>	$\pm 1.0$	-	-
<b>Cross Section</b>	$\pm 0.5$	-	-
<b>Spallation Dead Time</b>	$\pm 0.1$	$\pm 0.1$	$\pm 0.1$
<b>Live Time</b>	$\pm 0.1$	$\pm 0.1$	$\pm 0.1$
<b>Total</b>	+2.8 -2.7	+1.2 -1.1	+1.3 -1.2

# Global fit with SK 708 days flux

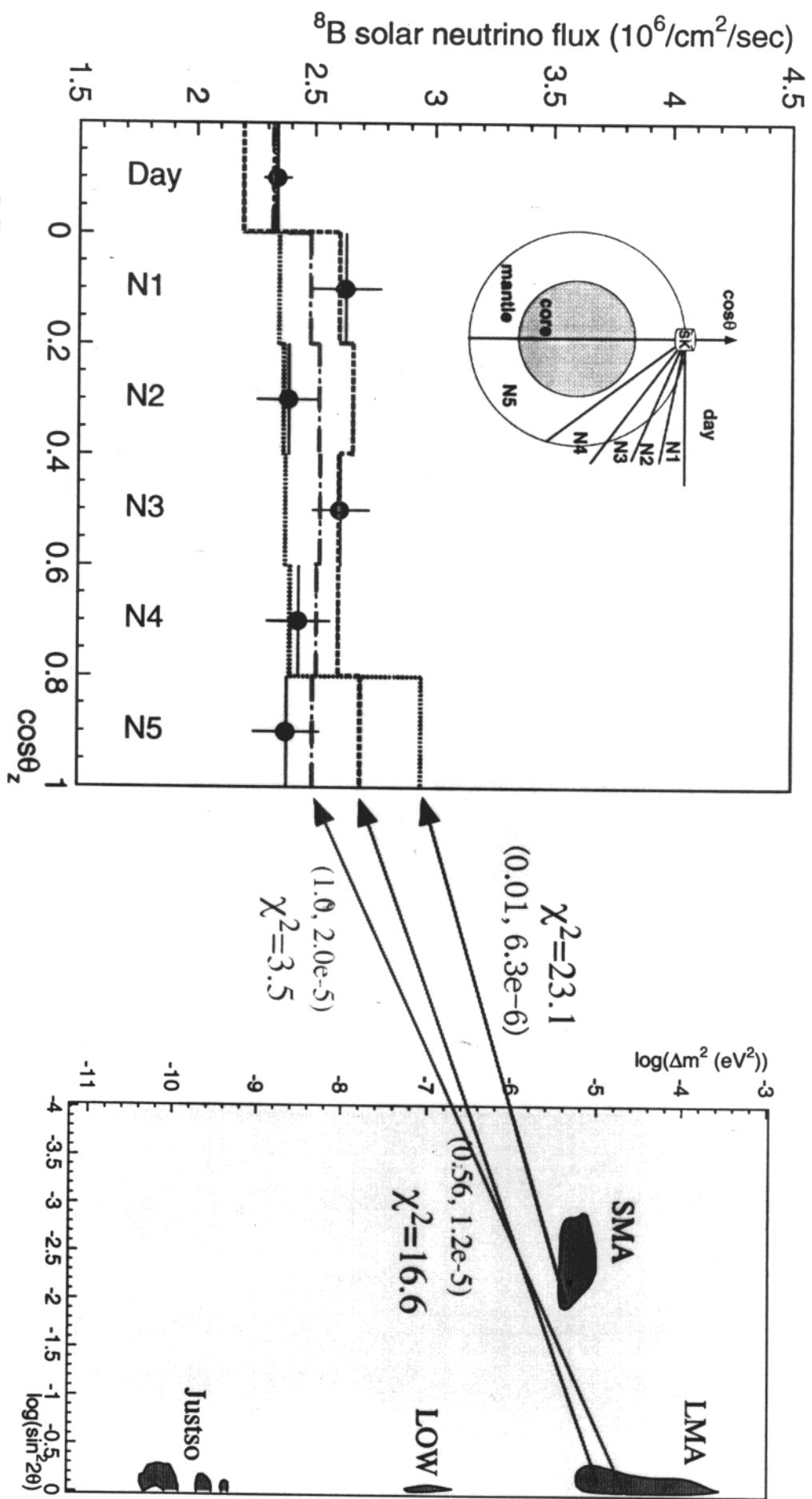


$\chi^2_{\min} = 0.9$   
 $\text{@ } (6.3\text{e-}3, 5.0\text{e-}6)$

no oscillation  
 $\chi^2 = 62.0$



# Day-Night variation



$$\frac{N}{D} - 1 = 0.060 \pm 0.036 \text{ (stat)} \pm 0.008 \text{ (sys)}$$

$$\frac{N_5}{\langle D, N_1-N_4 \rangle} - 1 = -0.013 \quad {}^{+0.060}_{-0.058} \text{ (stat)} \pm 0.013 \text{ (sys)}$$

# $\chi^2$ definitions

Day-Night variation

$$\chi_{DN}^2 = \sum_{i=D,N} \sum_{j=1,18} \left\{ \frac{\phi_i - \alpha \times \phi_i^{osci}(\sin^2 2\theta, \Delta m^2)}{\sqrt{\sigma_i^2 + \sigma_{sys,i}^2}} \right\}^2$$

$\alpha$ : free

DN spectra

$$\chi_{DN-spectra}^2 = \sum_{i=D,N} \sum_{j=1,18} \left\{ \frac{\phi_{ij} - \phi_{ij}^{osci} \times \alpha \times f_j(\varepsilon_{exp}, \varepsilon_{theo})}{\sigma_{ij}} \right\}^2 + \left( \frac{\varepsilon_{exp}}{\sigma_{exp}} \right)^2 + \left( \frac{\varepsilon_{theo}}{\sigma_{theo}} \right)^2$$

f: response function

$$\alpha: \text{free} \quad \sigma_{ij} = \sqrt{\sigma_{stat,ij}^2 + \sigma_{uncorr.sys,ij}^2}$$

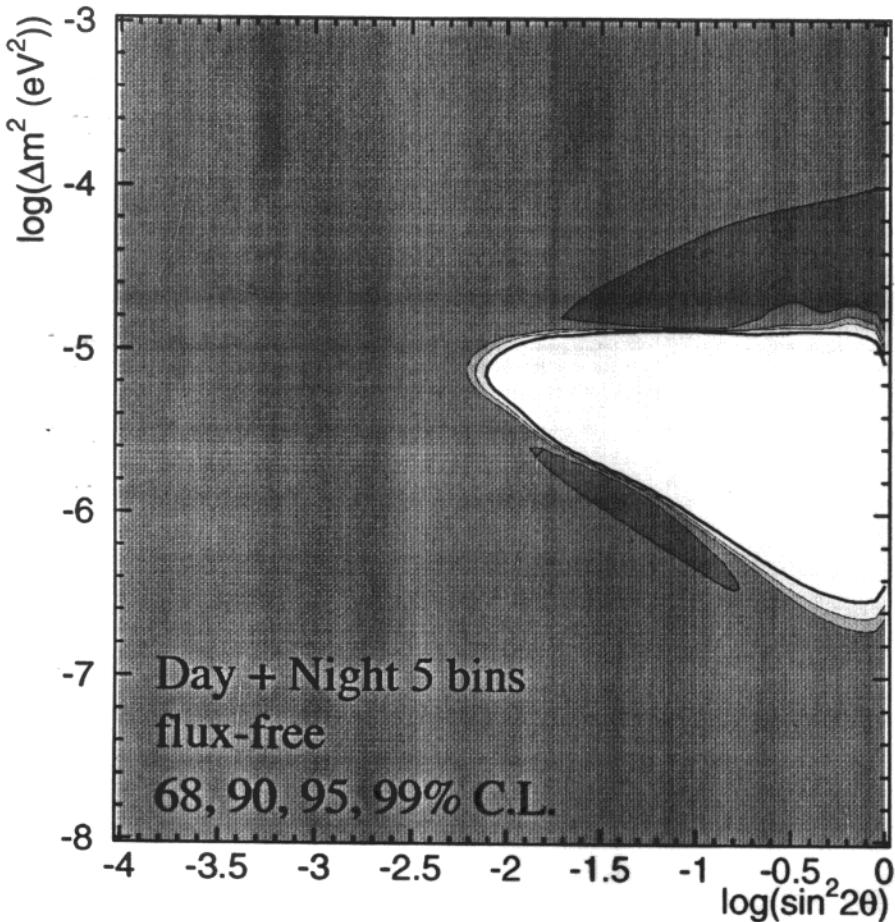
$\varepsilon$ : correlated errors

DN spectra, hep-free

$$\chi_{DN-spectra, hep-free}^2 = \sum_{i=D,N} \sum_{j=1,18} \left\{ \frac{\phi_{ij} - (\phi_{8B,ij}^{osci} \times \alpha_{8B} + \phi_{hep,ij}^{osci} \times \alpha_{hep}) \times f_j(\varepsilon_{exp}, \varepsilon_{theo})}{\sigma_{ij}} \right\}^2 + \left( \frac{\varepsilon_{exp}}{\sigma_{exp}} \right)^2 + \left( \frac{\varepsilon_{theo}}{\sigma_{theo}} \right)^2$$

$\alpha_{8B}, \alpha_{hep}$  : free

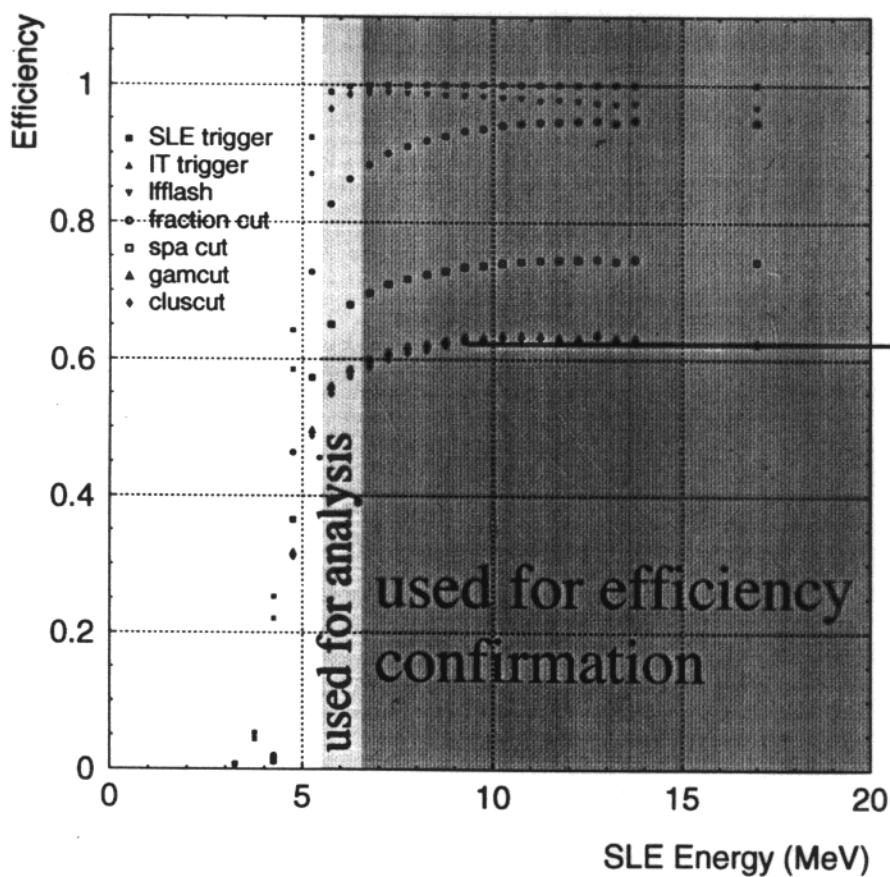
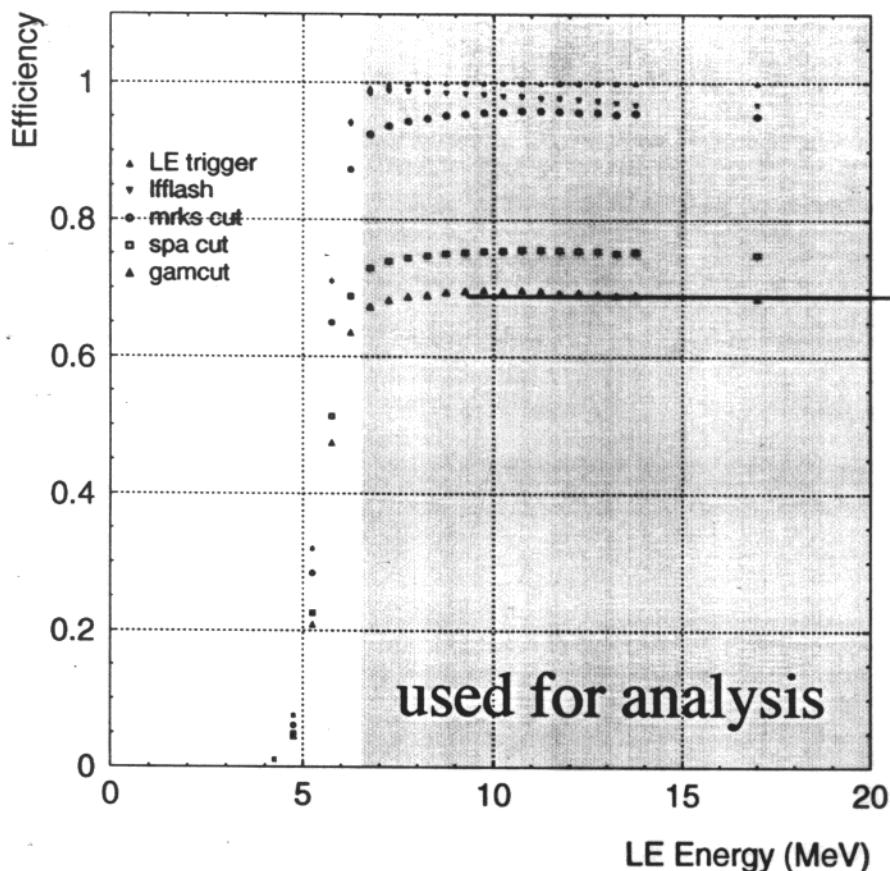
# Day–Night variation



$\chi^2_{\text{min}} = 3.5$   
@ (1.0, 2.0e-5)

$\chi^2_{\text{flat}} = 6.7$

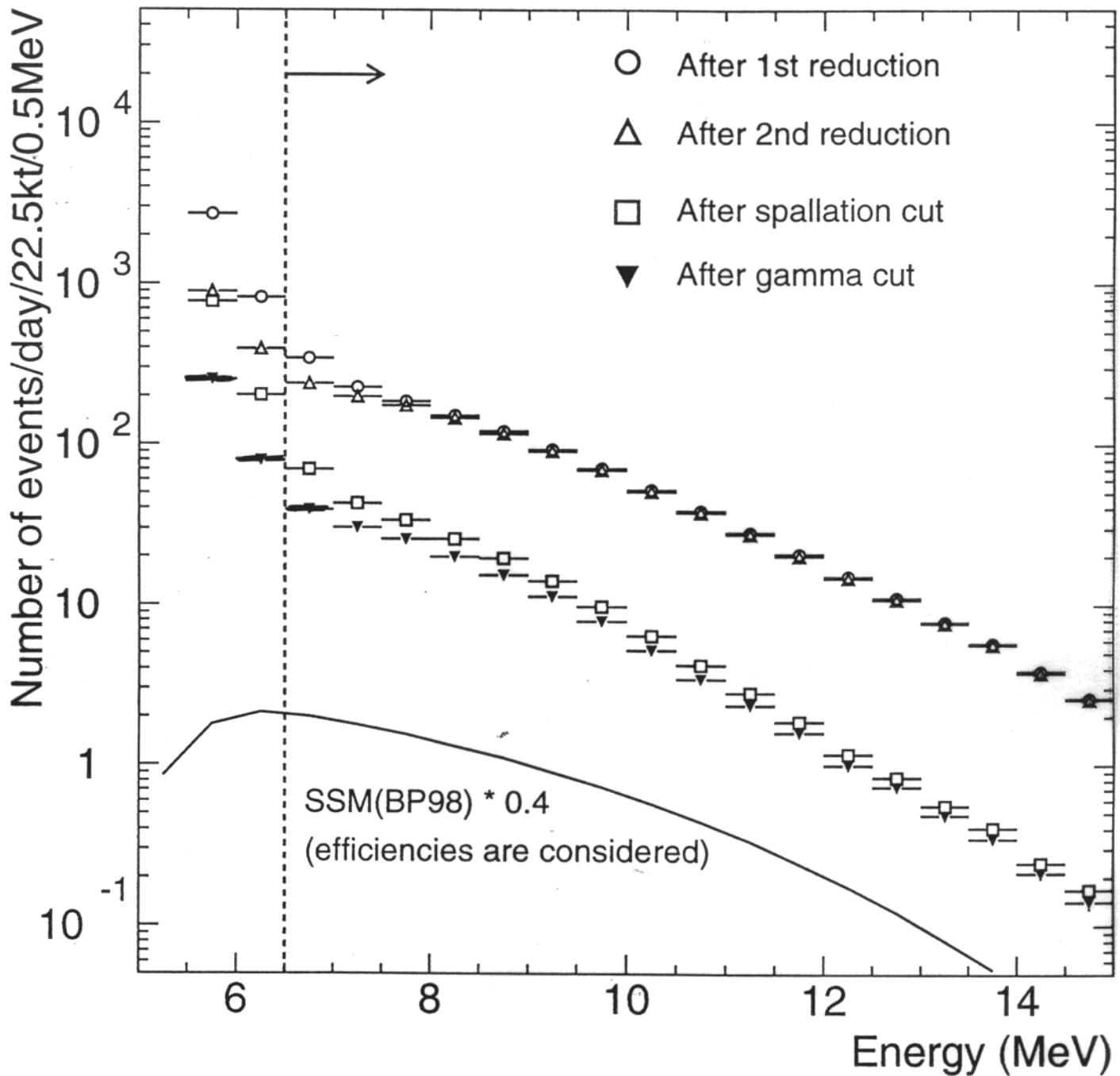
# Super Low Energy (SLE) Data



(Preliminary)

## Super-Kamiokande LE 708day

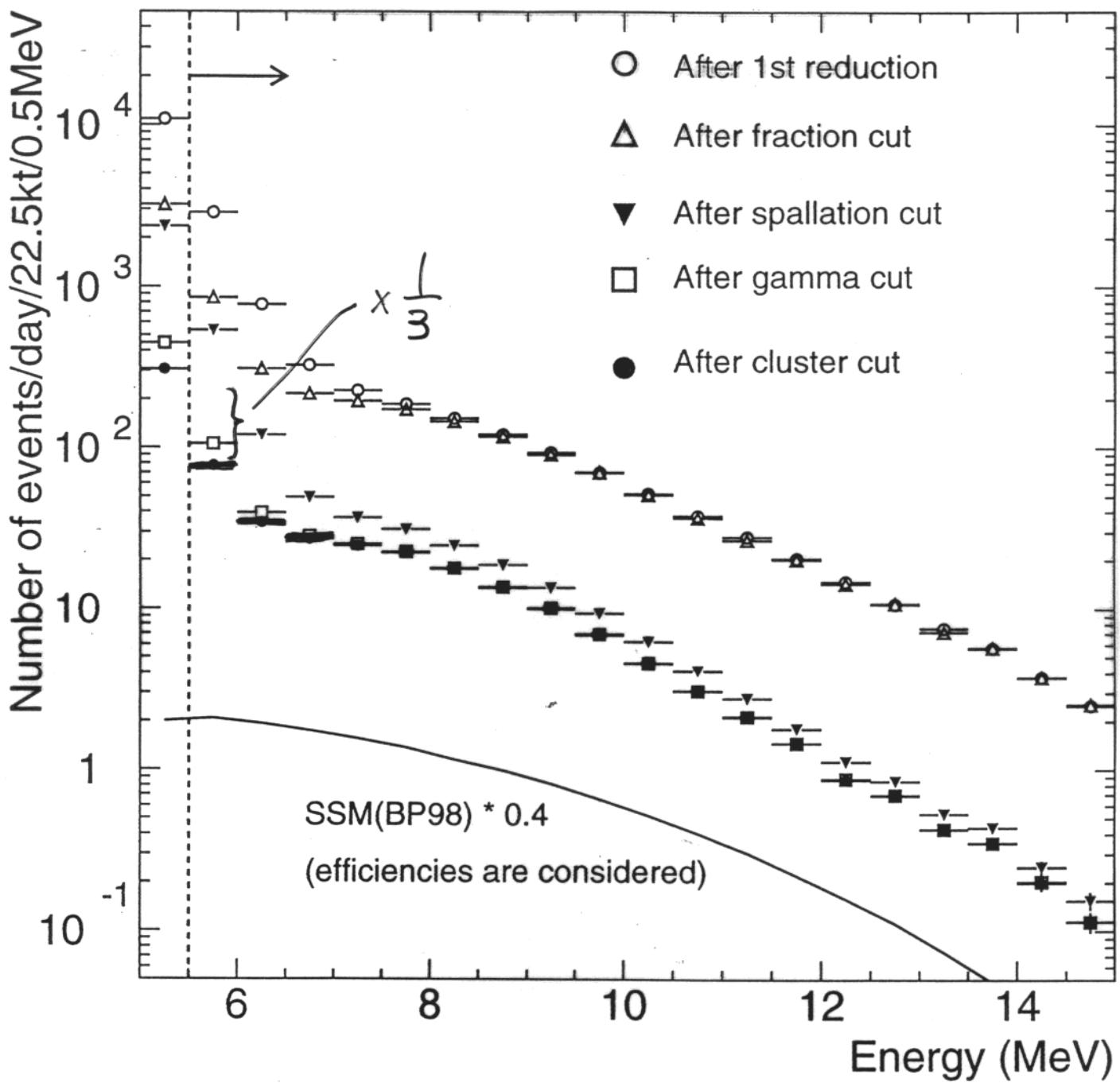
Fid. vol. 22.5kton, ALL



(Preliminary)

## Super-Kamiokande SLE 419day

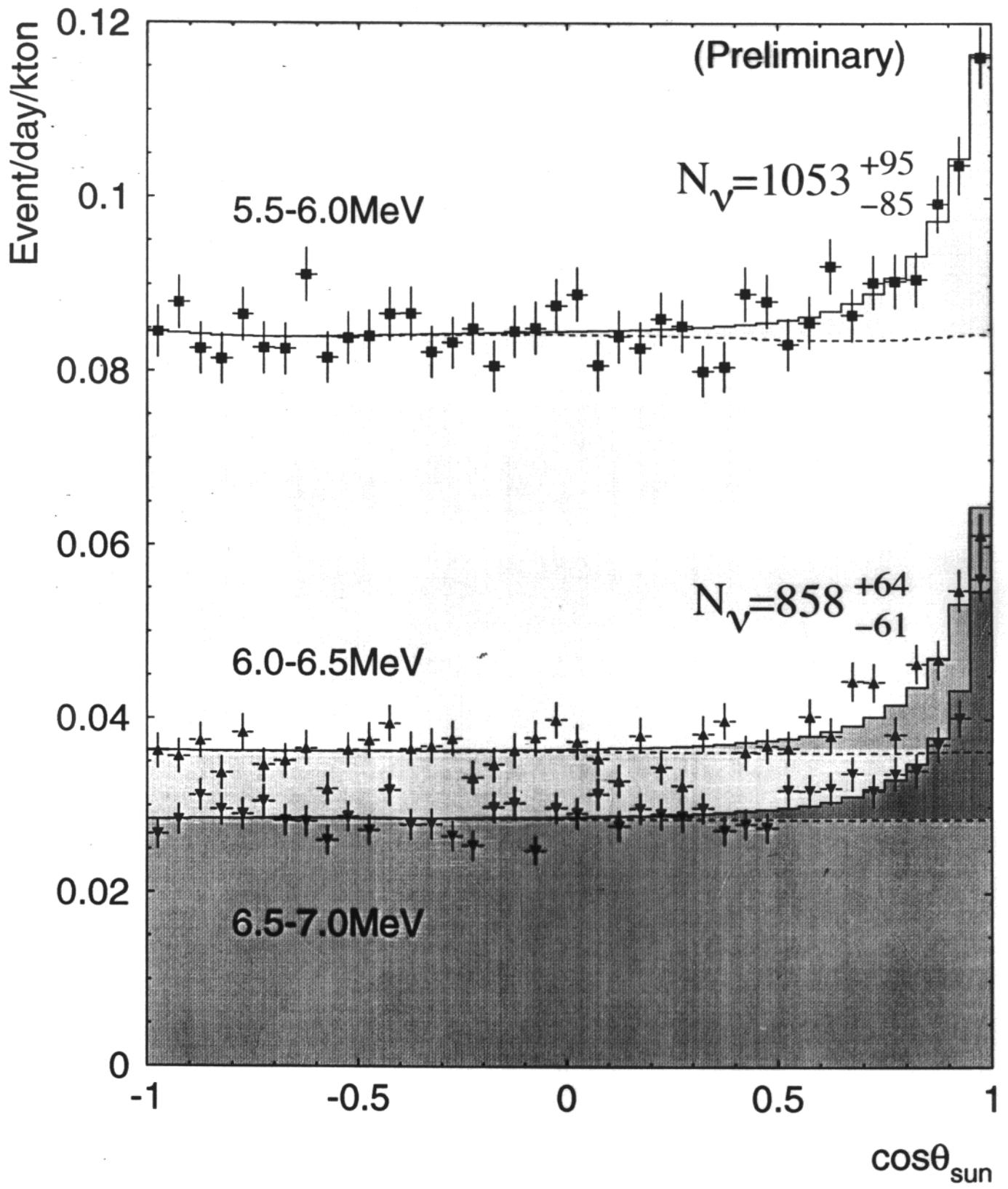
Fid. vol. 22.5kton, ALL



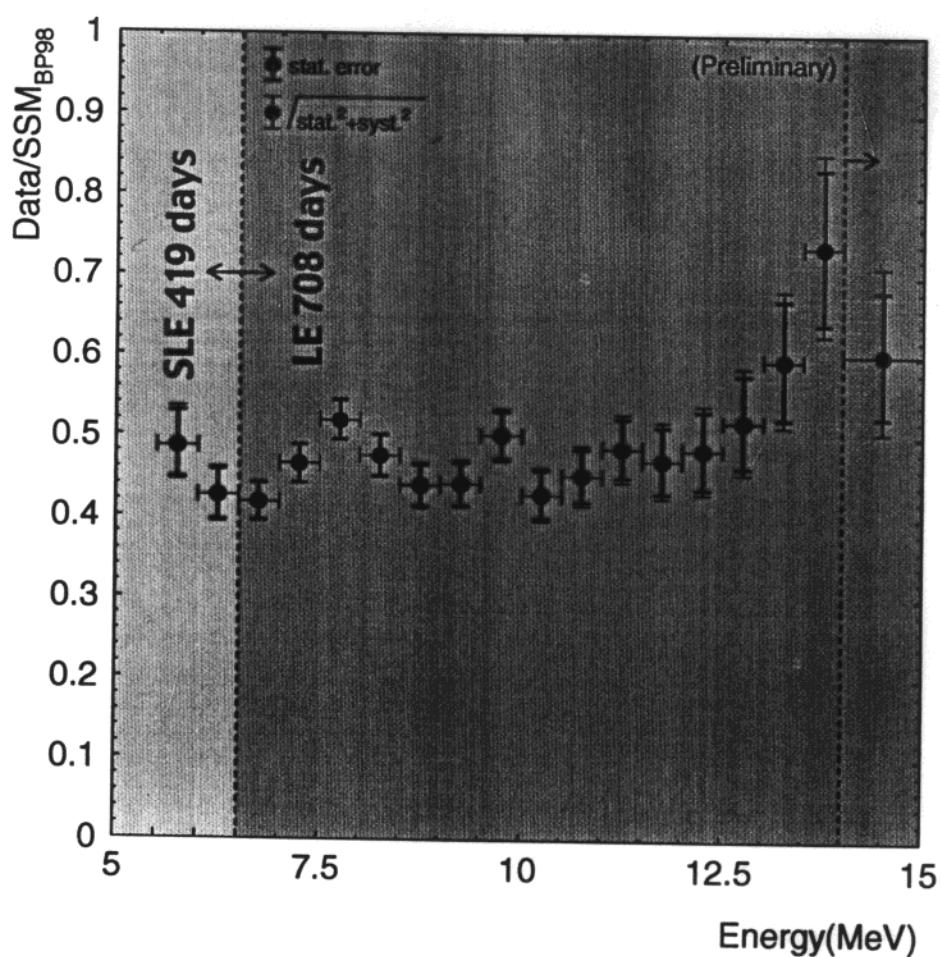
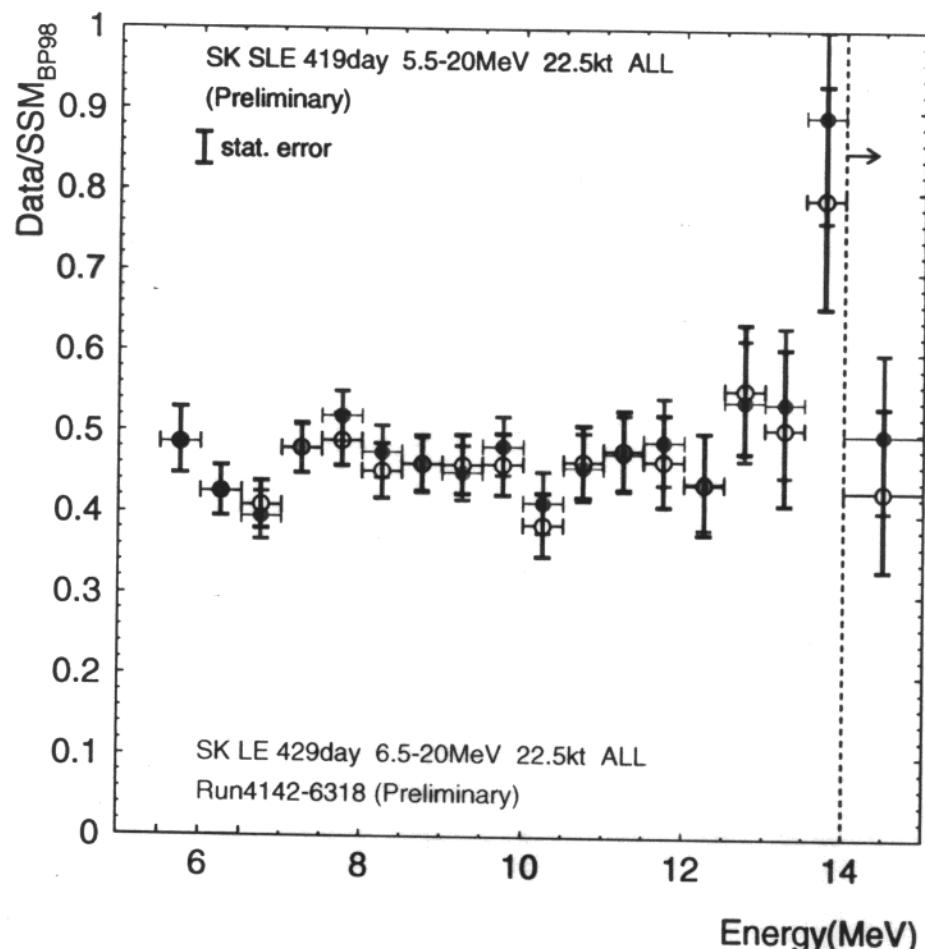
# Solar neutrino signal in the SLE data

SK SLE 419day 22.5kton

29 May 1997 - 4 Nov 1998

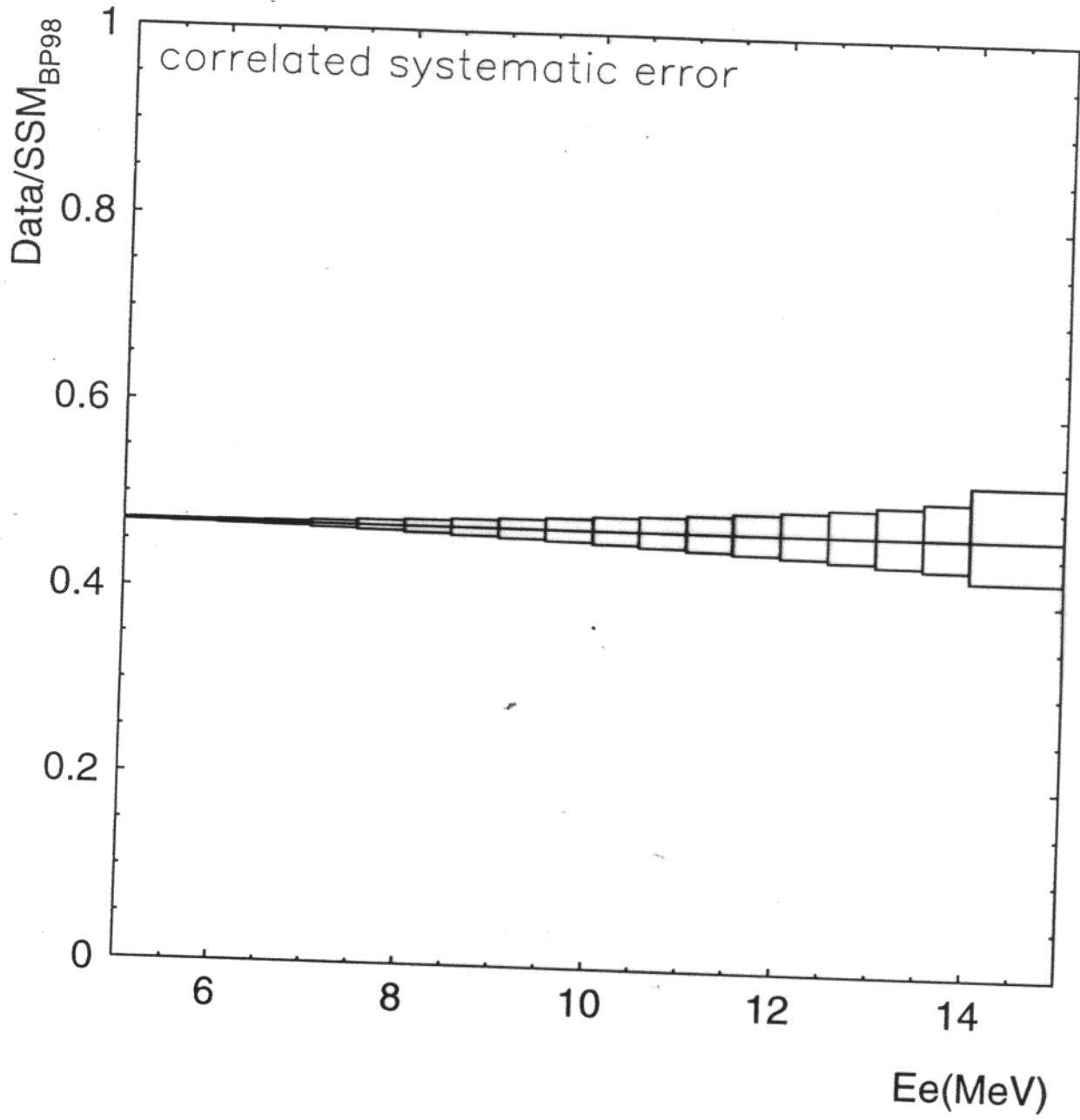


# Spectrum with SLE data

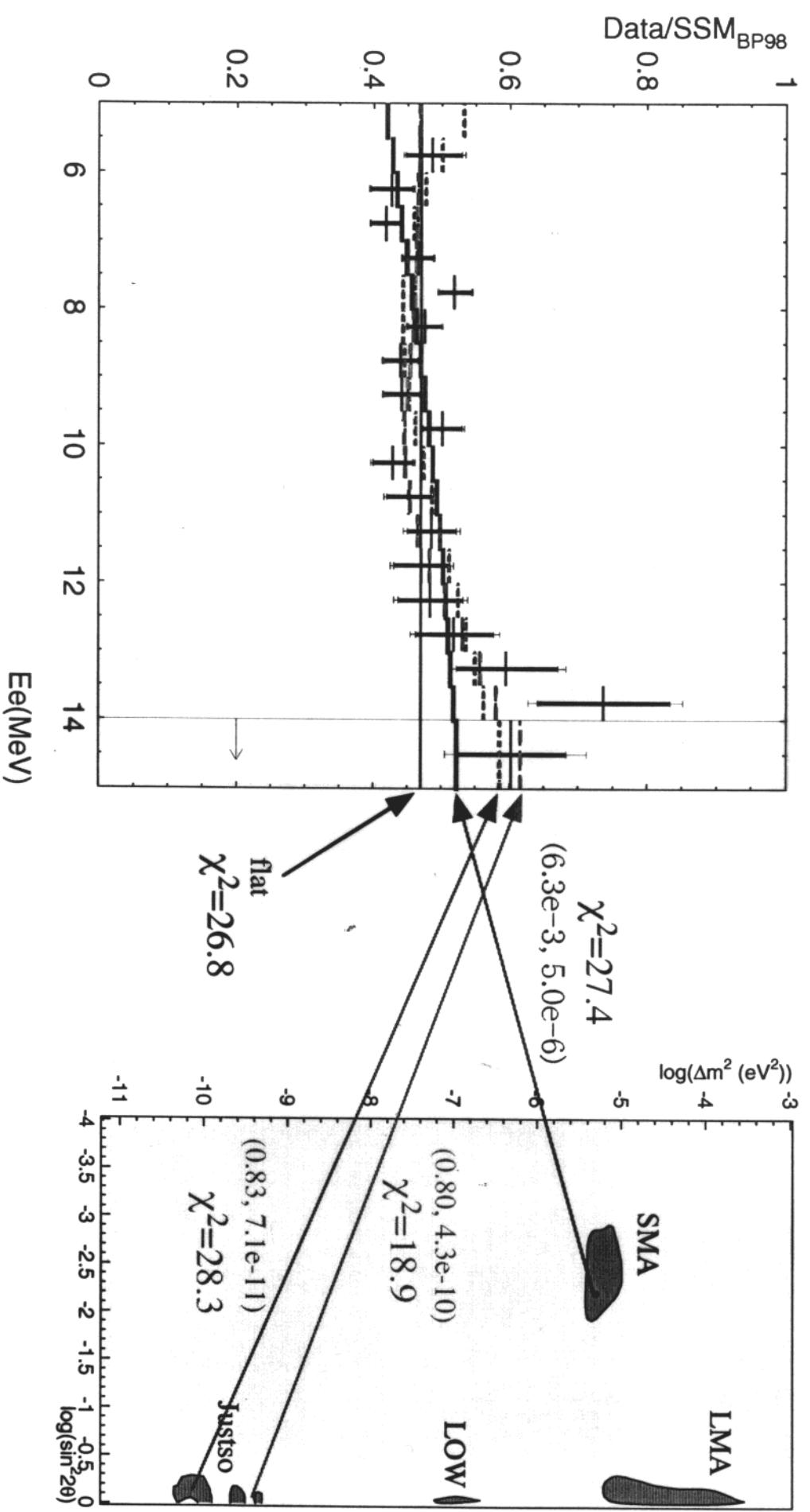


# Systematic Errors for energy spectrum

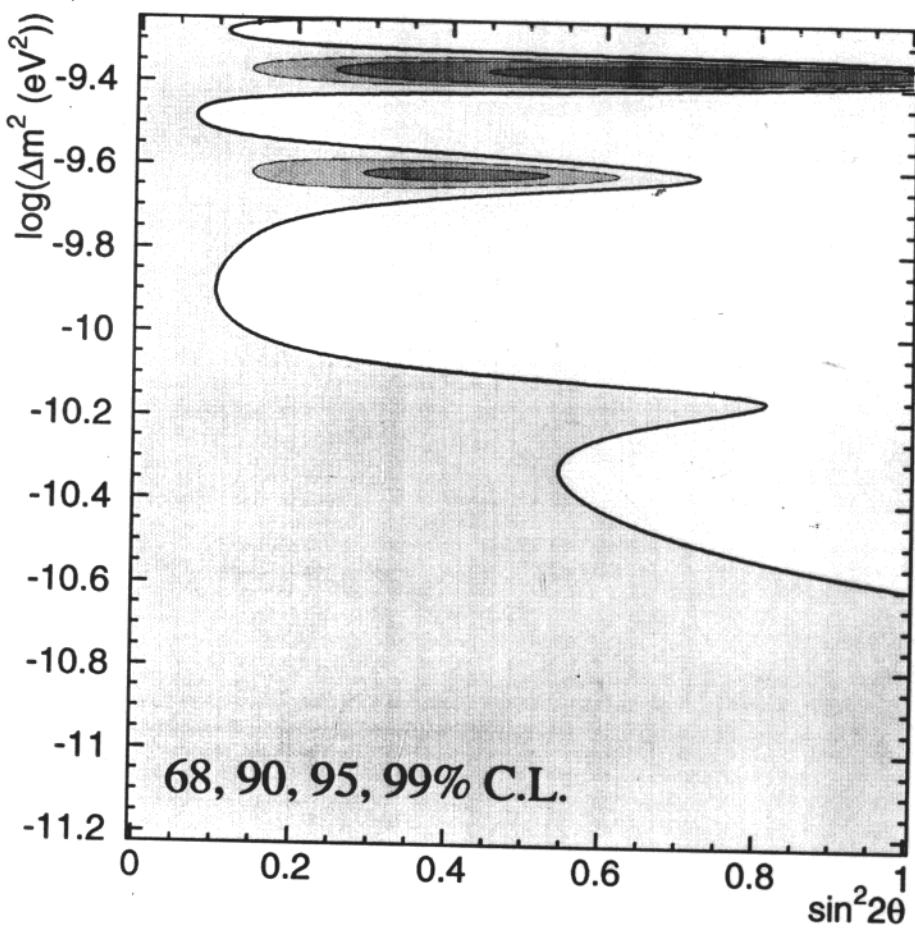
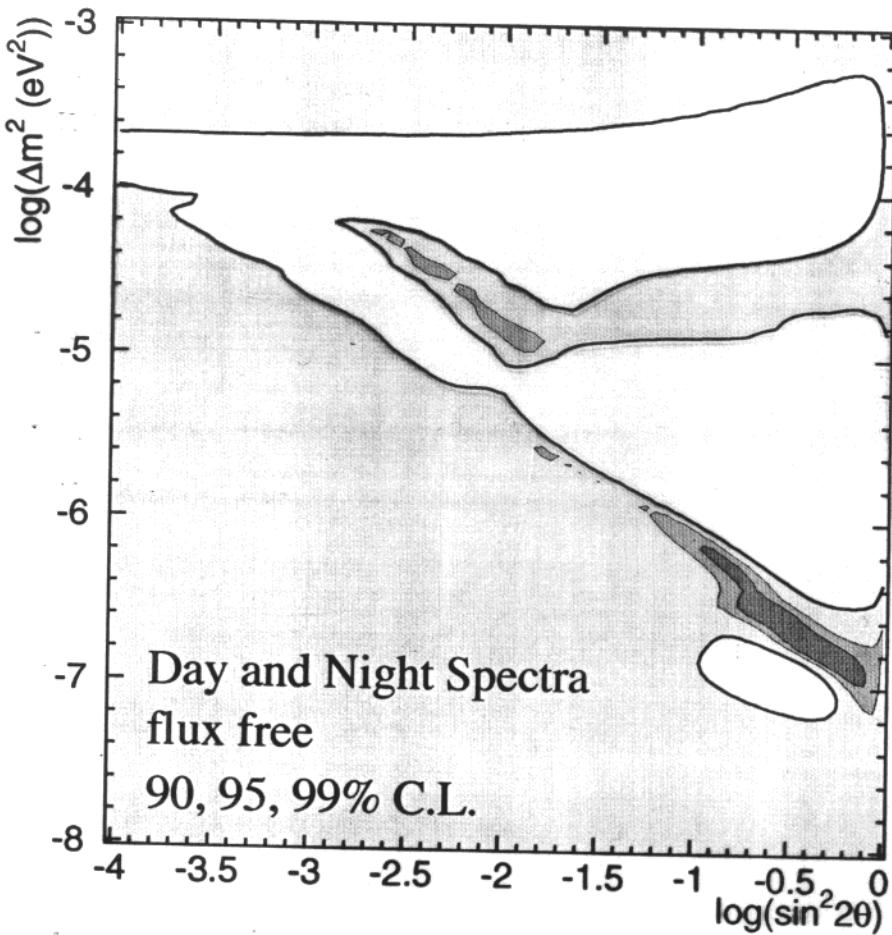
	SLE analysis		LE analysis	
	5.5-6.0	6.0-6.5	6.5-7.0	7.0-
<b>Correlated Error</b> E-Scale, Resolution, $^{8}\text{B}$ spectrum			see figure	
<b>Trigger Efficiency</b>	+2.6	+0.4	+1.1 -0.9	-
<b>Noise event Cut</b>	+2.7 -1.8	+2.4 -1.3	$\pm 0.7$	$\leftarrow$
<b>Reduction</b>	$\pm 0.2$	$\leftarrow$	$\leftarrow$	$\leftarrow$
<b>Vertex Shift</b>	$\pm 0.5$	$\leftarrow$	$\leftarrow$	$\leftarrow$
<b>non-flat B.G.</b>	$\pm 0.1$	$\leftarrow$	$\leftarrow$	$\leftarrow$
<b>Direction</b>	$\pm 1.0$	$\leftarrow$	$\leftarrow$	$\leftarrow$
<b>Cross Section</b>	$\pm 0.5$	$\leftarrow$	$\leftarrow$	$\leftarrow$
<b>Spallation Dead Time</b>	-	$\leftarrow$	$\leftarrow$	$\leftarrow$
<b>Live Time</b>	-	$\leftarrow$	$\leftarrow$	$\leftarrow$
<b>Relative between SLE &amp; LE</b>	$\pm 2.0$	$\pm 2.0$	-	$\leftarrow$
<b>Total</b>			see figure	



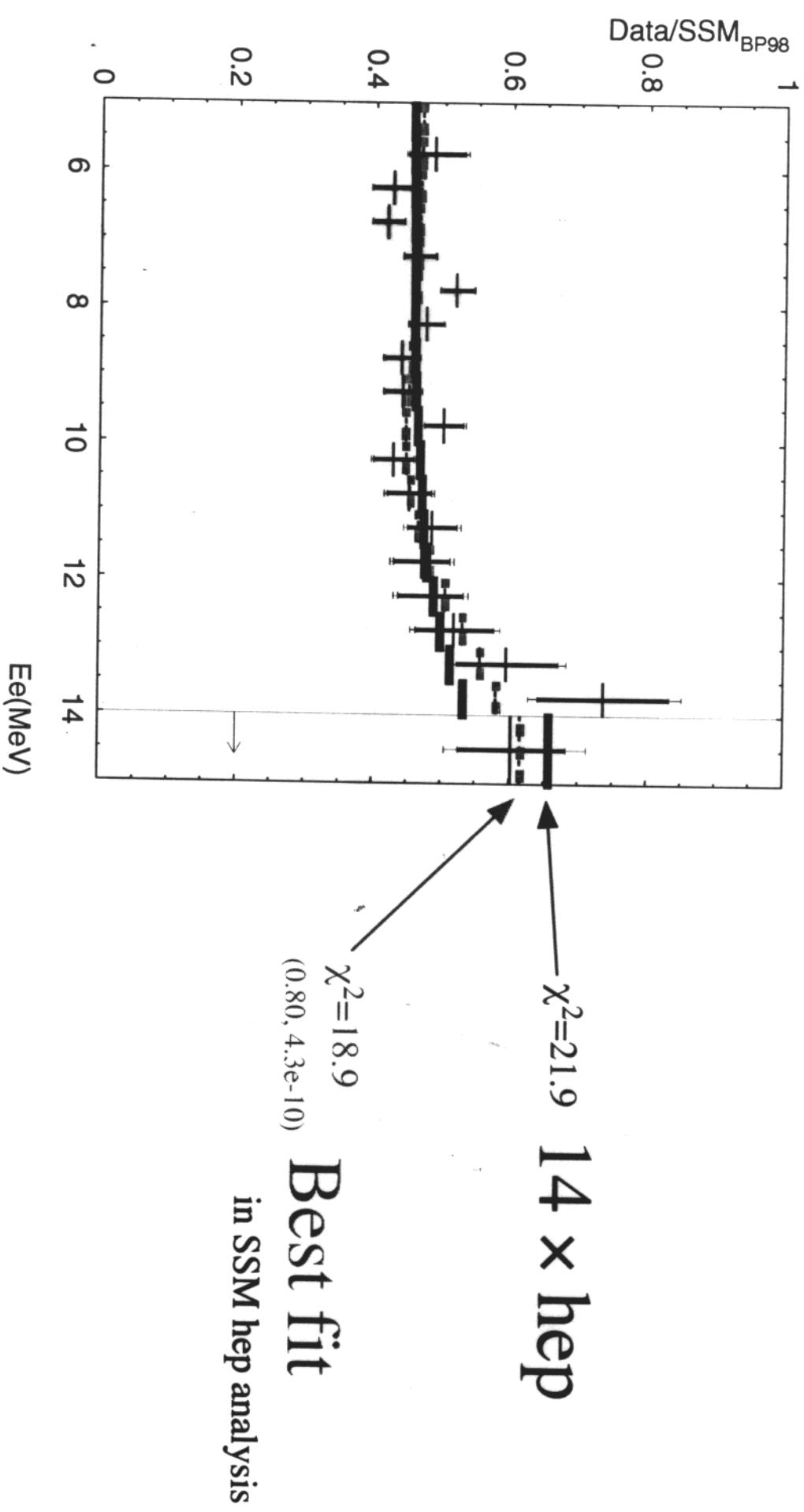
# Spectrum (18 bins, 8B flux free)



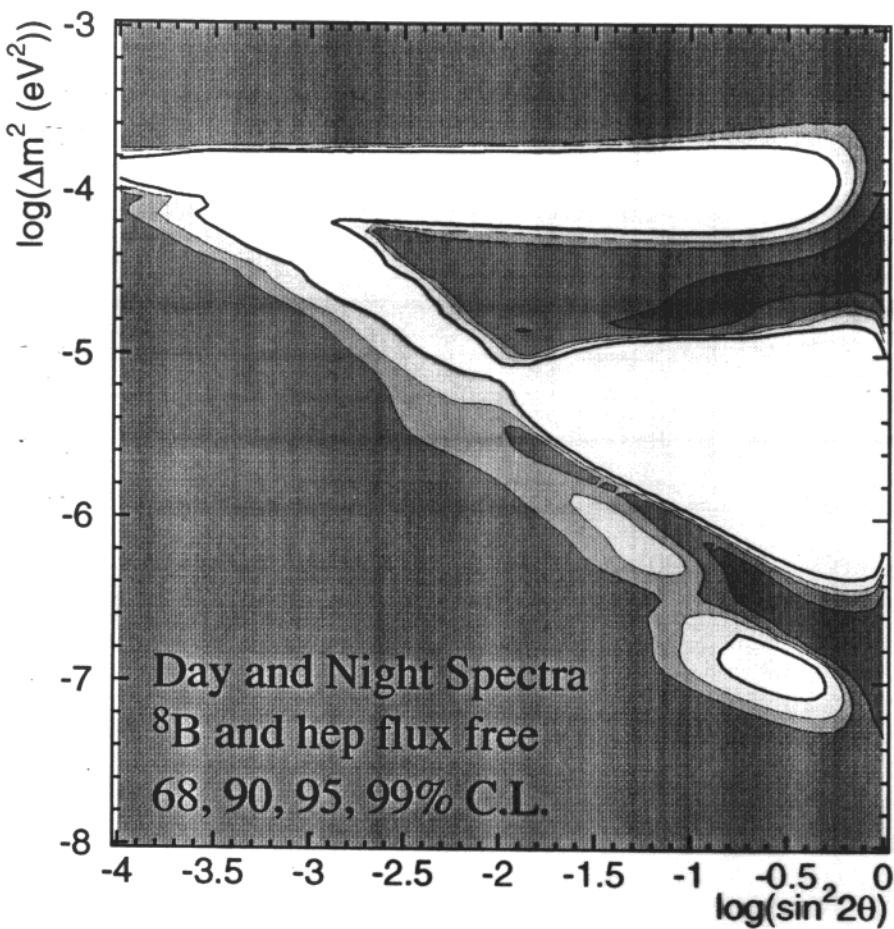
# SK Day and Night Spectra (36 bins, ${}^8\text{B}$ flux free)



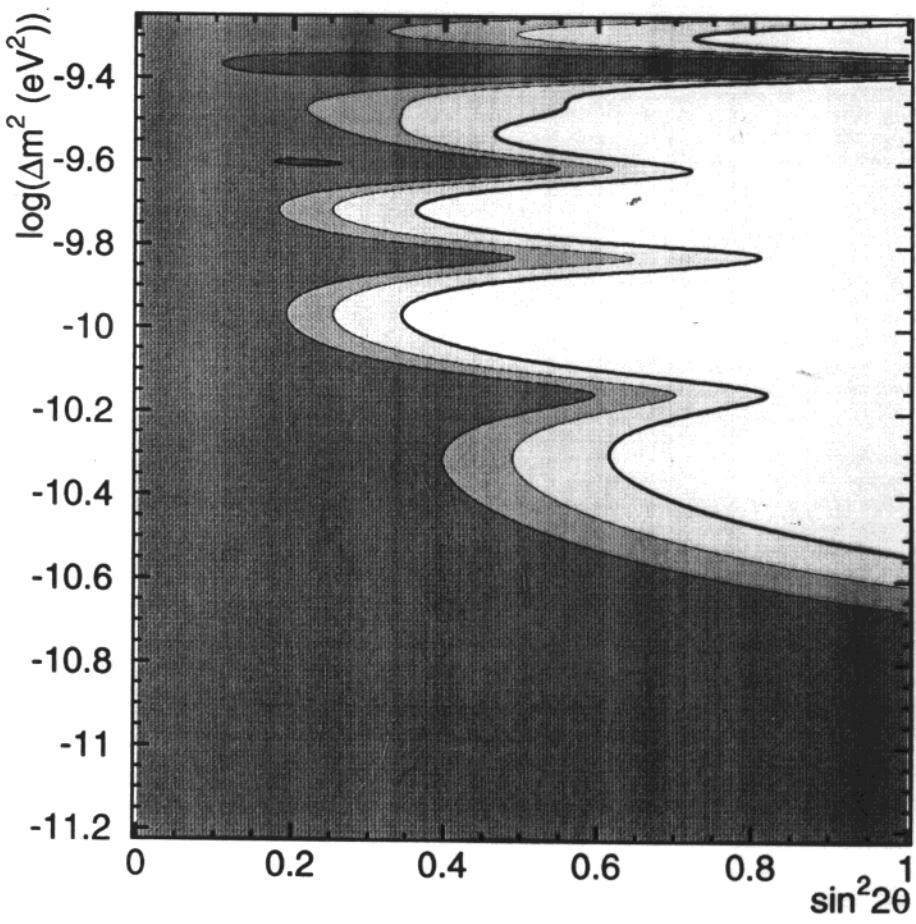
# Spectrum (18 bins, ${}^{18}\text{B}$ flux free, hep flux free)



# SK DN spectra (36 bins, ${}^8\text{B}$ and hep flux free)

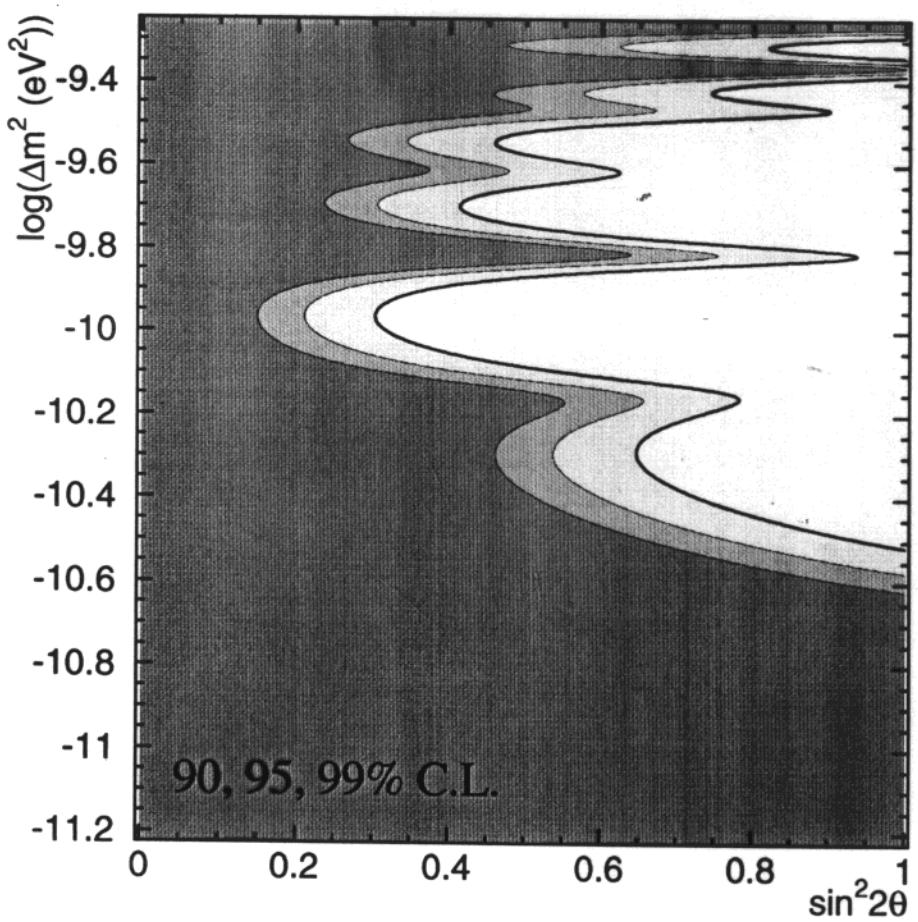
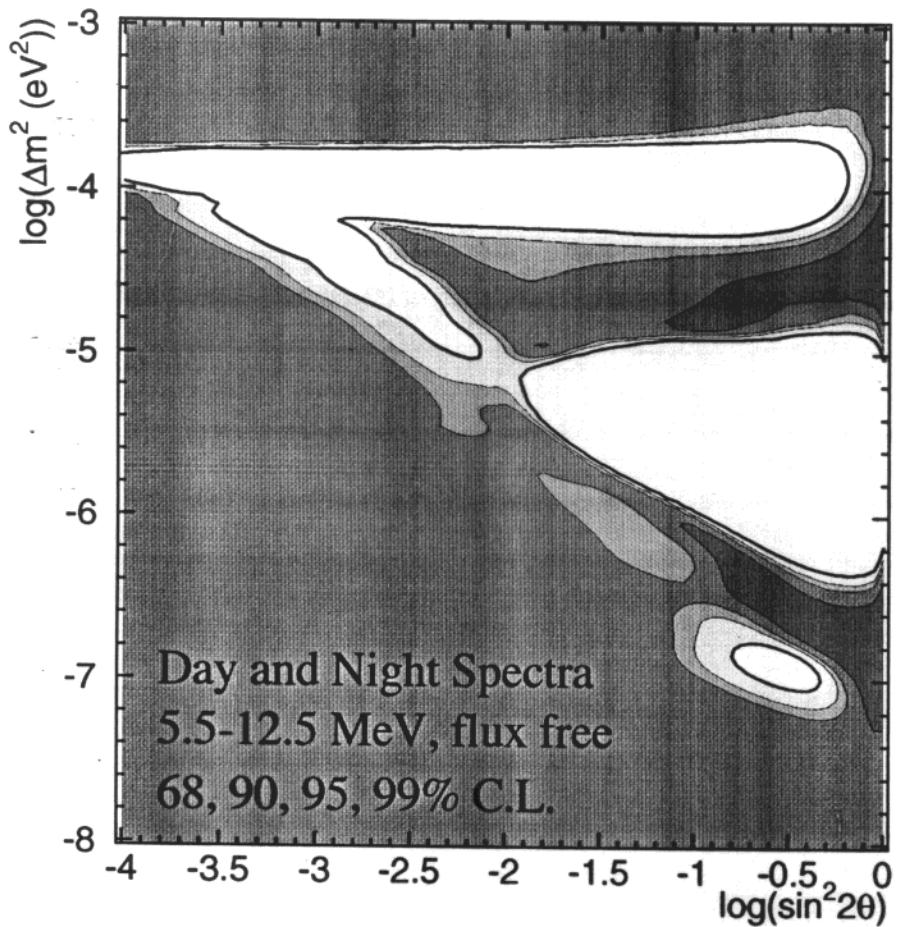


no oscillation  
 $\chi^2 = 49.9$  hep=14

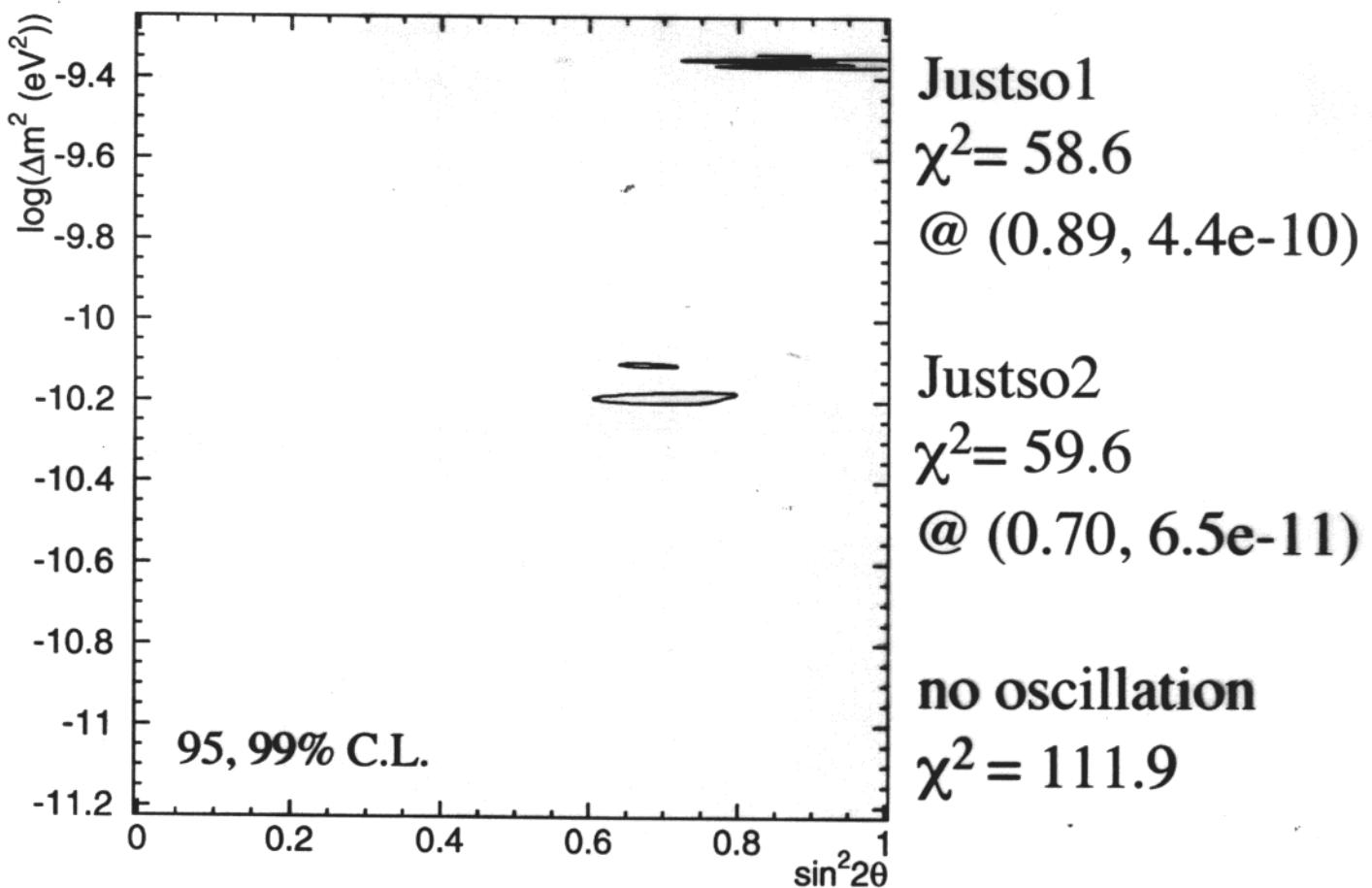
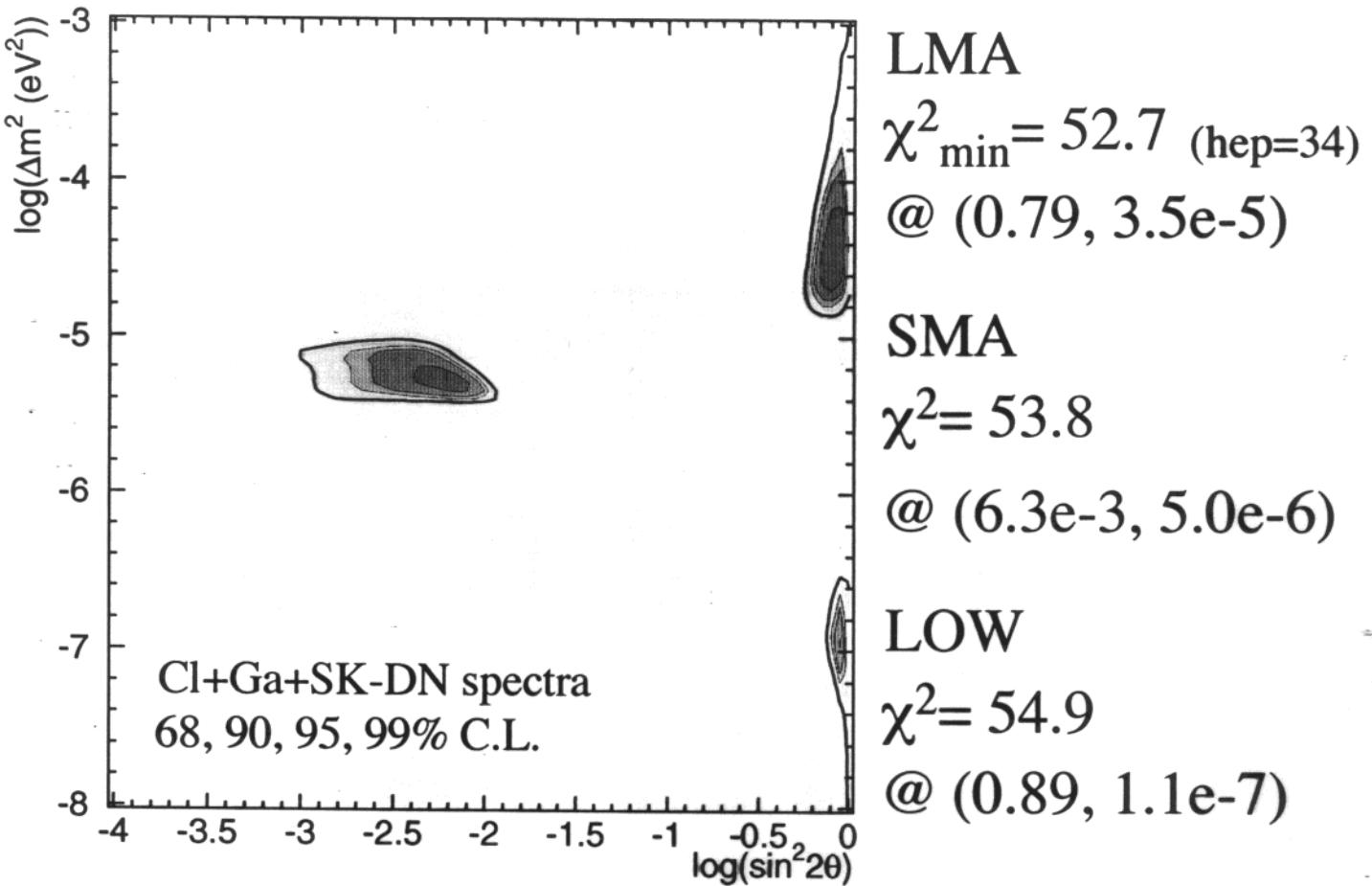


$\chi^2_{\min} = 47.0$   
@ (0.80, 4.3e-10)

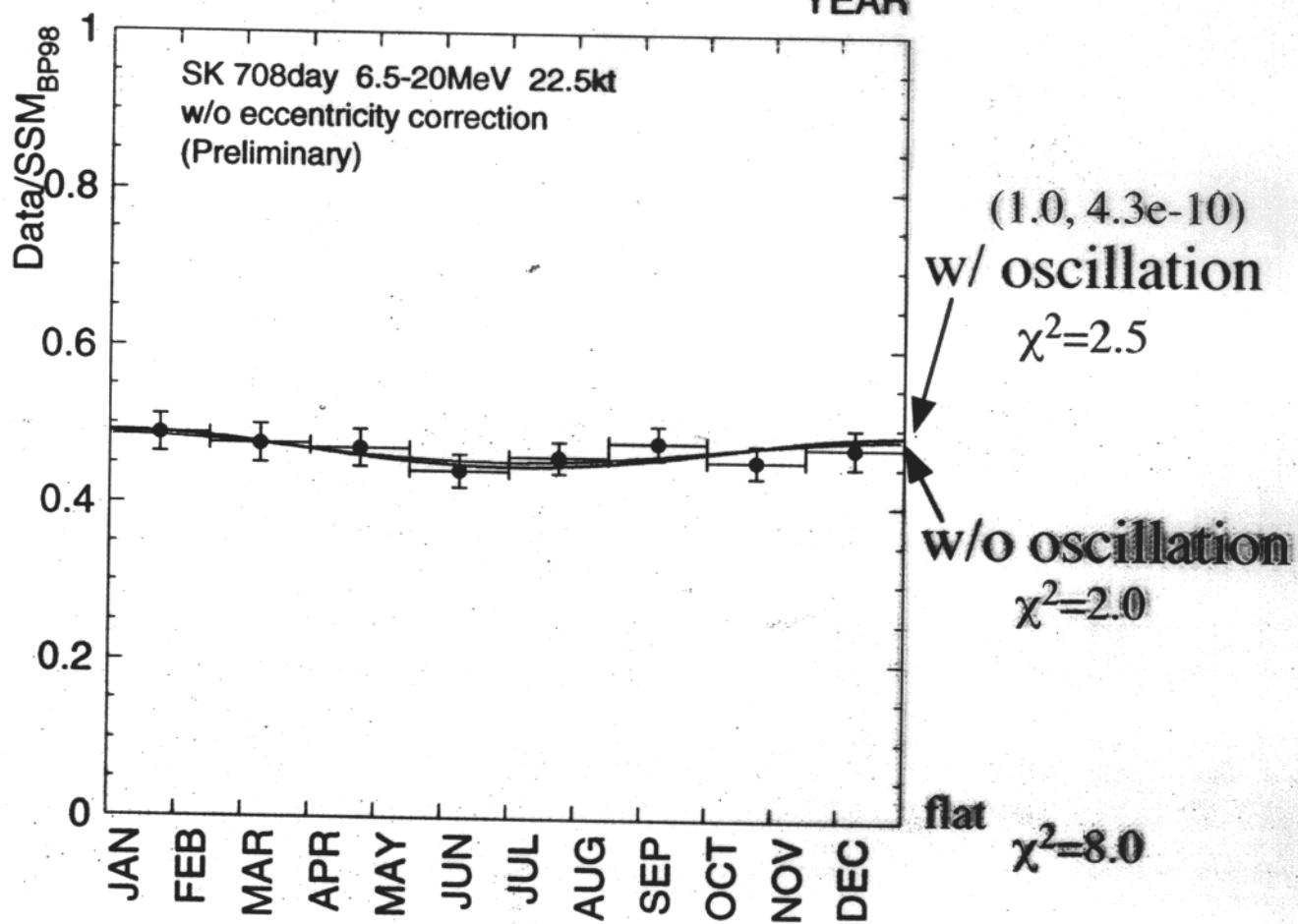
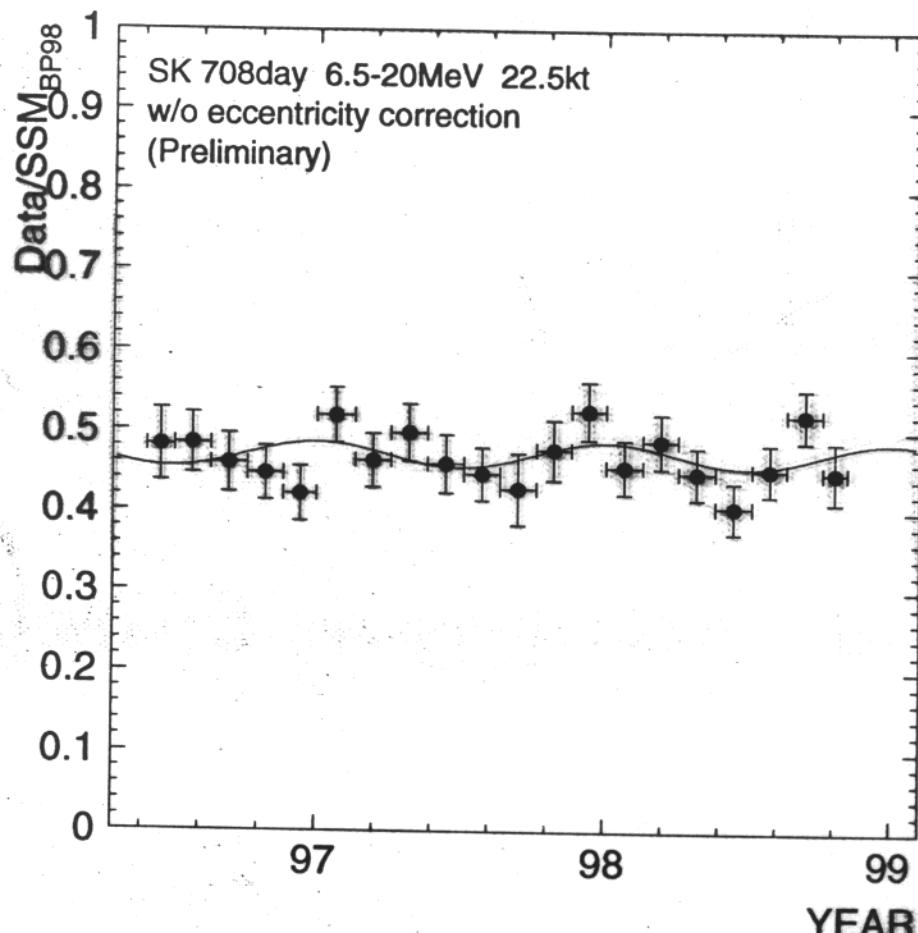
# DN spectra, 5.5–12.5 MeV (28 bins, ${}^8\text{B}$ flux free)



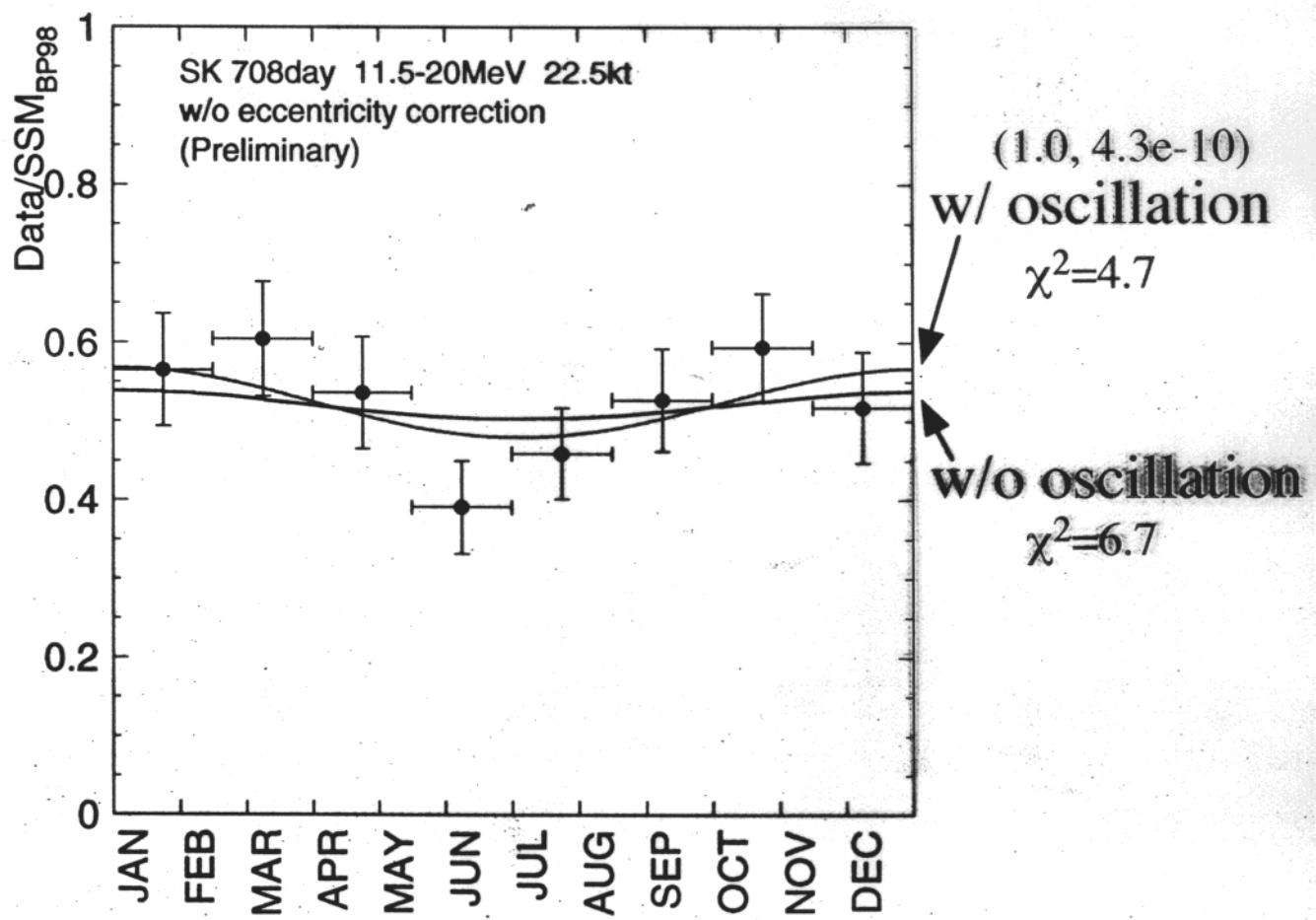
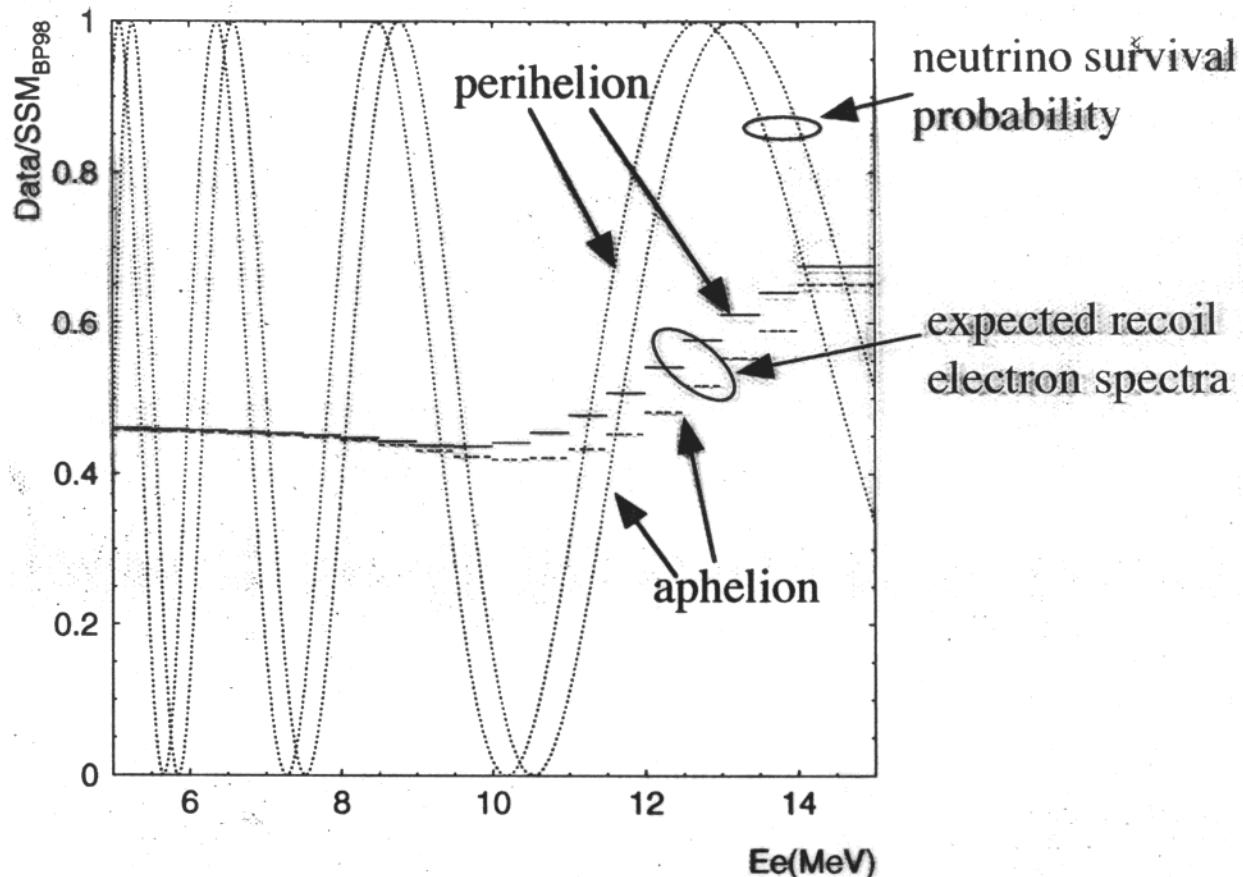
# Global Fit with SK DN spectra, hep-free



# Seasonal variation



# Seasonal variation in high energy



# Summary

- ◊  ${}^8\text{B}$  solar neutrino flux was obtained with 708 days of data.  
 $\phi({}^8\text{B}) = 2.419 {}^{+0.042}_{-0.039} \text{ (stat)} {}^{+0.068}_{-0.065} \text{ (sys)} \times 10^6 / \text{cm}^2/\text{sec}$
- ◊ No significant but small DN asymmetry was seen.
- ◊  $N/D - 1 = 0.060 \pm 0.036 \text{ (stat)} \pm 0.008 \text{ (sys)}$   
 $N5/\langle D, N1-N4 \rangle - 1 = -0.013 {}^{+0.060}_{-0.058} \text{ (stat)} \pm 0.013 \text{ (sys)}$
- ◊ The analysis threshold has been lowered to 5.5 MeV with SLE data.  
Flat in low energy side and slope in high energy side  
Just so or large hep-flux seem to be necessary.
- ◊ No significant seasonal variation is seen, yet.  
If the slope is due to the justso oscillation,  
seasonal variation will be seen in high energy bins.