

Prospects for the Coming Accelerator and Reactor Neutrino Experiments

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Physics Beyond the Standard Model

Theoretical arguments:

SM does not exist without cutoff

Higgs-doublet = only simplest extension

Gauge hierarchy problem

Why: 3 generations, fermion representations

Many parameters (9+? Masses, 4+? Mixings)

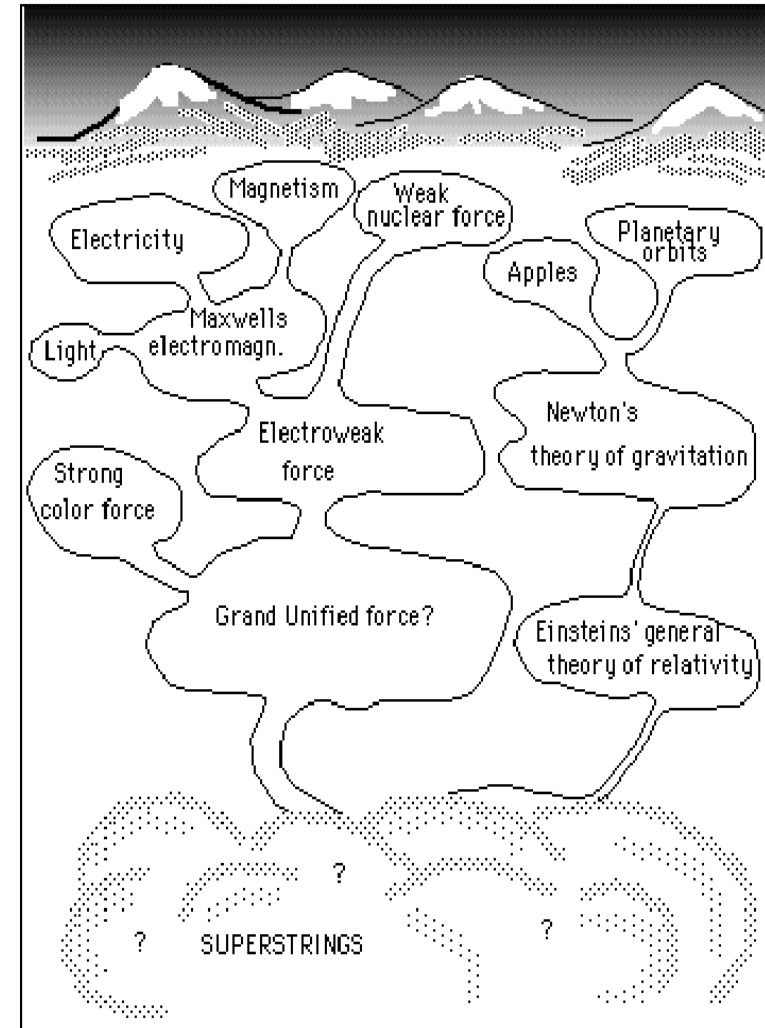
Charge quantisation, unification: GUTs, ...,

Gravitation, ...

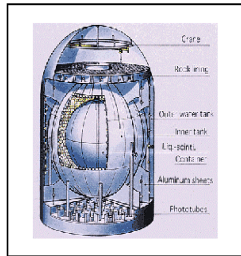
2 directions: Sym. breaking & Flavour

Experimental facts:

- **Dark Matter & Dark Energy exist!**
- **Neutrino masses have been detected!**
- **Baryon asymmetry of the universe $\leftrightarrow m_\nu > 0$**
 - ➔ **physics beyond the standard model**
 - ➔ **results $\leftrightarrow$ implications for theory**



Current Oscillation Results

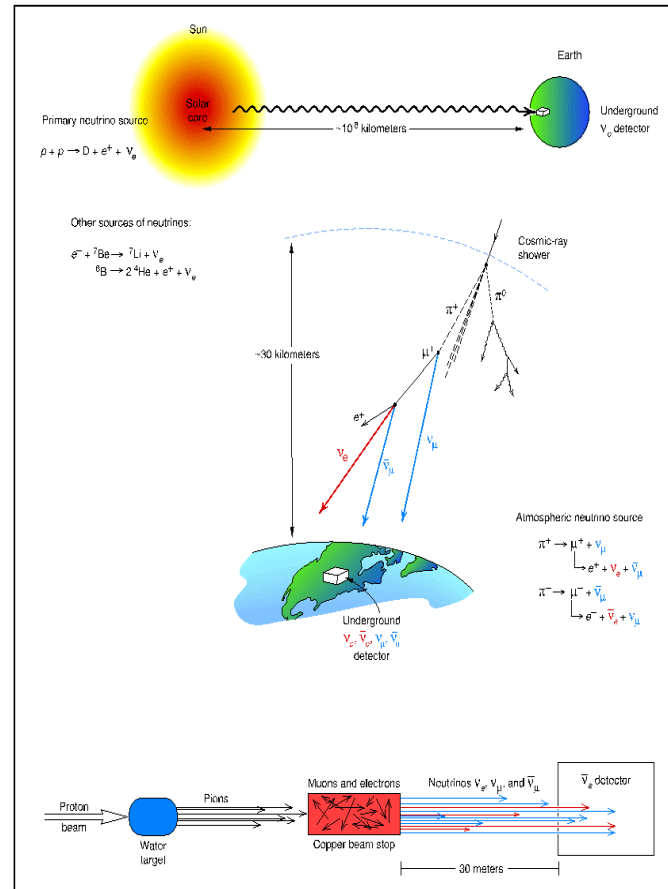


KamLAND **solar**
CHOOZ

atmospheric
+
K2K

neutrino
oscillation
signals

LSND?



$$\Delta m_{12}^2 = 8.2 \pm 0.3 * 10^{-5} \text{eV}^2$$

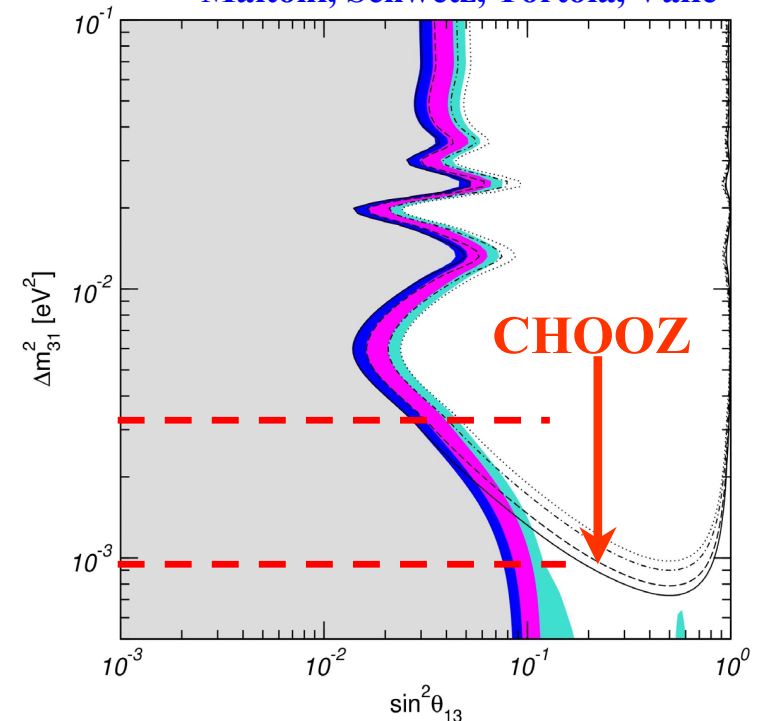
$$\tan^2 \theta_{12} = 0.39 \pm 0.05$$

$$\Delta m_{13}^2 = 2.2 \pm 0.6 * 10^{-3} \text{eV}^2$$

$$\tan^2 \theta_{23} = 1.0 \pm 0.3$$

$$\sin^2 \theta_{13} \leq 0.041 @ 3\sigma$$

Maltoni, Schwetz, Tortola, Valle



→ approximately two 2x2 oscillations

Accelerator & Reactor Experiments Now:

- CHOOZ & Palo Verde

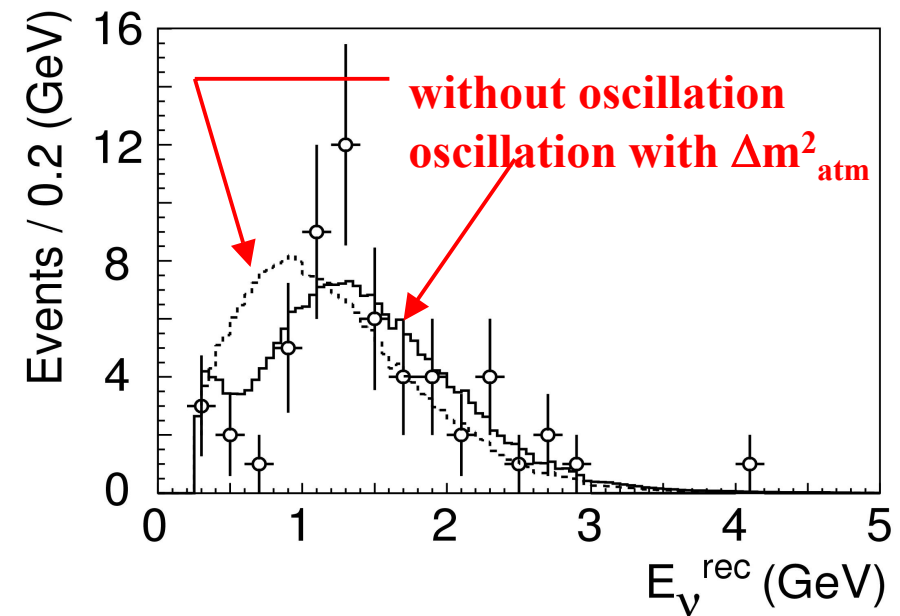
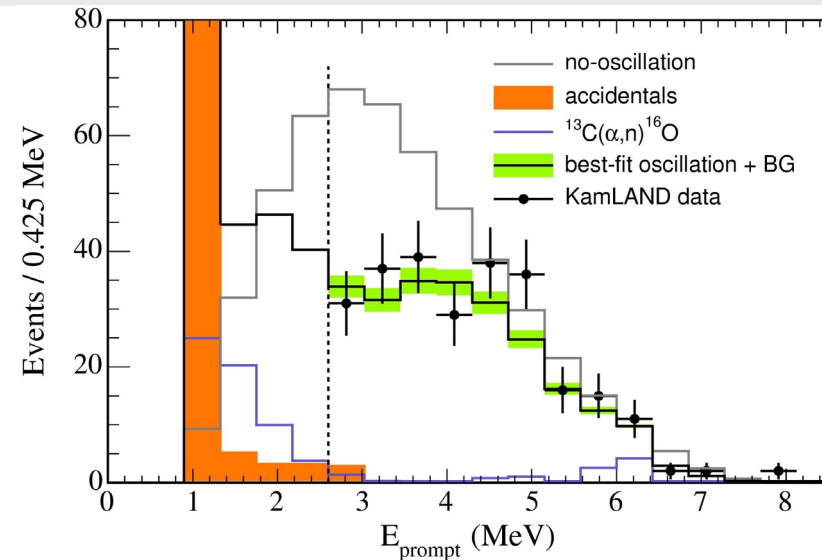
$\longleftrightarrow \theta_{13}$ limit

- KamLAND

$\longleftrightarrow$ solar oscillation

- K2K

$\longleftrightarrow$ atmospheric oscillation



The Future of Oscillation Physics

Δm^2 and θ_{ij} regions fixed:

- precision oscillation experiments
- controlled detectors & neutrino sources

- long baseline accelerator experiments with neutrino beams
- reactor experiments with identical near & far detector

$$\begin{pmatrix} U_{e1} & U_{e2} & U_{e3} \\ U_{\mu 1} & U_{\mu 2} & U_{\mu 3} \\ U_{\tau 1} & U_{\tau 2} & U_{\tau 3} \end{pmatrix} = \underbrace{\begin{pmatrix} 1 & 0 & 0 \\ 0 & c_{23} & s_{23} \\ 0 & -s_{23} & c_{23} \end{pmatrix}}_{\theta_{23}} \underbrace{\begin{pmatrix} c_{13} & 0 & s_{13}e^{-i\delta} \\ 0 & 1 & 0 \\ -s_{13}e^{i\delta} & 0 & c_{13} \end{pmatrix}}_{\substack{S_{13} \leftrightarrow 3f\text{-effects} \\ \rightarrow S_{13} \rightarrow \delta}} \underbrace{\begin{pmatrix} c_{12} & s_{12} & 0 \\ -s_{12} & c_{12} & 0 \\ 0 & 0 & 1 \end{pmatrix}}_{\theta_{12}} \times \text{Majorana-CP-phases}$$

Aims:

- **improved precision** of the leading 2x2 oscillations
- detection of **generic 3-neutrino effects**

Three Flavour Oscillations

$$\begin{Bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{Bmatrix} \xrightarrow{\text{oscillation}} \begin{Bmatrix} \nu_e \\ \nu_\mu \\ \nu_\tau \end{Bmatrix}$$

$$\begin{Bmatrix} \bar{\nu}_e \\ \bar{\nu}_\mu \\ \bar{\nu}_\tau \end{Bmatrix} \xrightarrow{\text{oscillation}} \begin{Bmatrix} \bar{\nu}_e \\ \bar{\nu}_\mu \\ \bar{\nu}_\tau \end{Bmatrix}$$

2 flavour approximations:

$$P_{ab} = \sin^2(2\theta) \sin^2(\Delta m^2 L / 4E)$$

$$P_{aa} = 1 - P_{ab}$$

➔ precision not sufficient

3 flavour-oscillations

$$J_{ij}^{e_l e_m} := U_{li} U_{lj}^* U_{mi}^* U_{mj} \quad \Delta_{ij} := \frac{\Delta m_{ij}^2 L}{4E} = \frac{(m_i^2 - m_j^2) L}{4E}$$

$$P(\nu_{e_l} \rightarrow \nu_{e_m}) = \delta_{lm} - \underbrace{4 \sum_{i>j} \text{Re} J_{ij}^{e_l e_m} \sin^2 \Delta_{ij}}_{P_{CP}} - \underbrace{2 \sum_{i>j} \text{Im} J_{ij}^{e_l e_m} \sin 2\Delta_{ij}}_{P_{CP}}$$

+ matter effects
➔ $\text{sgn}(\Delta m^2)$

- 2→3 flavours ➔ more mixings & CP phase
- MSW matter effects

Analytic Approximations

- $\Delta = \Delta m_{31}^2 L / 4E$
- qualitative understanding $\Rightarrow$ expand in $\alpha = \Delta m_{21}^2 / \Delta m_{31}^2$ and $\sin^2 2\theta_{13}$
- matter effects $\hat{A} = A / \Delta m_{31}^2 = 2VE / \Delta m_{31}^2$; $V = \sqrt{2}G_F n_e$

$$P(\nu_\mu \rightarrow \nu_\mu) \approx 1 - \cos^2 \theta_{13} \sin^2 2\theta_{23} \sin^2 \Delta + 2\alpha \cos^2 \theta_{13} \cos^2 \theta_{12} \sin^2 2\theta_{23} \Delta \cos \Delta$$

$$\begin{aligned}
 P(\nu_e \rightarrow \nu_\mu) \approx & \sin^2 2\theta_{13} \sin^2 \theta_{23} \frac{\sin^2((1-\hat{A})\Delta)}{(1-\hat{A})^2} \\
 & \pm \sin \delta_{\text{CP}} \alpha \sin 2\theta_{12} \cos \theta_{13} \sin 2\theta_{23} \sin(\Delta) \frac{\sin(\hat{A}\Delta) \sin((1-\hat{A})\Delta)}{\hat{A}(1-\hat{A})} \\
 & + \cos \delta_{\text{CP}} \alpha \sin 2\theta_{12} \cos \theta_{13} \sin 2\theta_{23} \cos(\Delta) \frac{\sin(\hat{A}\Delta) \sin((1-\hat{A})\Delta)}{\hat{A}(1-\hat{A})} \\
 & + \alpha^2 \sin^2 2\theta_{12} \cos^2 \theta_{23} \frac{\sin^2(\hat{A}\Delta)}{\hat{A}^2}
 \end{aligned}$$

$\rightarrow$ analytic discussion / full simulations
 $\rightarrow$ degeneracies, correlations $\rightarrow (\sin^2 2\theta_{13})_{\text{eff}}$

Cervera et al.
 Freund, Huber, ML
 Akhmedov, Johansson, ML, Ohlsson, Schwetz

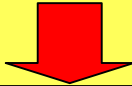
Degeneracies, Correlations, ...

Fixed L/E $\rightarrow$ probabilities invariant under transformations:

- $\theta_{23} \rightarrow \pi/2 - \theta_{23}$ Fogli, Lisi
 $P(\nu_e \rightarrow \nu_\mu)$ not really invariant $\rightarrow$ compensation by small parameter off-sets
- $\Delta m^2 \rightarrow -\Delta m^2$ compensated by offset in δ Minakata, Nunokawa
- $P(\nu_e \rightarrow \nu_\mu) = \text{const.} \rightarrow \delta - \theta_{13}$ manifolds Koike, Ota, Sato & Burguet-Castell et al.
- $\rightarrow$ 8-fold degeneracy Barger, Marfatia, Whisnant
- $P(\nu_e \rightarrow \nu_\mu)$ depends on $\sin(\Delta\phi) = \sin(VL/2)$
 $VL=2\pi \rightarrow$ magic baseline: $L_{\text{magic}} = 2\pi/V = 8100 \text{ km}$ Huber, Winter

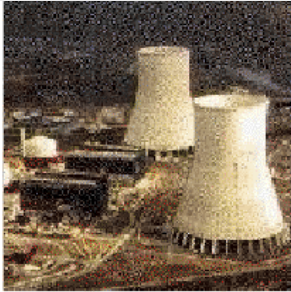
- parameter extraction suffers from correlations & degeneracies
- how to break degeneracies & correlations?

Future Long Baseline Experiments

K2K	analysis	establish atmospheric oscillations with beam
MINOS OPERA , ICARUS	running construction	<u>expected precision:</u> 8% for Δm^2_{13} , 25% for $\sin^2\theta_{23}$, θ_{13} ?
T2K	approved	4% for Δm^2_{13} , 15% for $\sin^2\theta_{23}$, $\rightarrow \theta_{13}$
NOvA	LOI	3% for Δm^2_{13} , 15% for $\sin^2\theta_{23}$ (combined with T2K) , $\rightarrow \theta_{13}$, $\rightarrow \delta$? , $\rightarrow \text{sgn}(\Delta m^2_{13})$
T2H	R&D	
β -beams	R&D	
neutrino factory	R&D	
...muon collider	...	
		precision neutrino physics

- every stage is a **necessary prerequisite** for the next
- continuous line of **improvements for beams, detectors, physics**

New Reactor Experiments



$\overline{\nu}_e$

near detector (170m)

$\overline{\nu}_e$

far detector (1700m)

identical detectors $\rightarrow$ many errors cancel

- **The survival probability:**

- expand in small quantities

$$P_{ee} \approx 1 - \sin^2 2\theta_{13} \sin^2 \frac{\Delta m_{31}^2 L}{4E_\nu} + \left(\frac{\Delta m_{21}^2 L}{4E_\nu} \right)^2 \cos^4 \theta_{13} \sin^2 2\theta_{12}$$

- **last term negligible** for $\frac{\Delta m_{31}^2 L}{4E_\nu} \sim \pi/2$ and $\sin^2 2\theta_{13} \gtrsim 10^{-3}$
- **atmospheric frequency** is dominant
- most important:

- **No degeneracies!**
- **Practically no correlations!**
- **No matter effects!**

Simulation of Future Experiments

- select a setup (beam, detector, baseline, ...)
- take „most realistic“ parameters $\leftrightarrow$ best guess!
- simulate all relevant aspects as good as possible

Source	⊗	Oscillation	⊗	Detector
- neutrino energy E		- oscillation channels		- effective mass, material
- flux and spectrum		- realistic baselines		- threshold, resolution
- flavour composition		- MSW matter profile		- particle ID (flavour, charge, event reconstruction, ...)
- contamination		- degeneracies		- backgrounds
- symmetric $\nu/\bar{\nu}$ operation		- correlations		- x-sections (at low E)

- determine the potential: „true“ $\leftrightarrow$ fitted parameters
- compare only realistic simulations (all relevant effects, errors & uncertainties)

Apologies

Simulations of various options by various:

Barger, Geer, Raja, Whisnant, Marfatia

Cervera, Donini, Gavela, Gomez-Cadenaz, Hernandez, Mena, Rigolin, ...

Bueno, Campanelli, Rubbia

Minakata, Yasuda

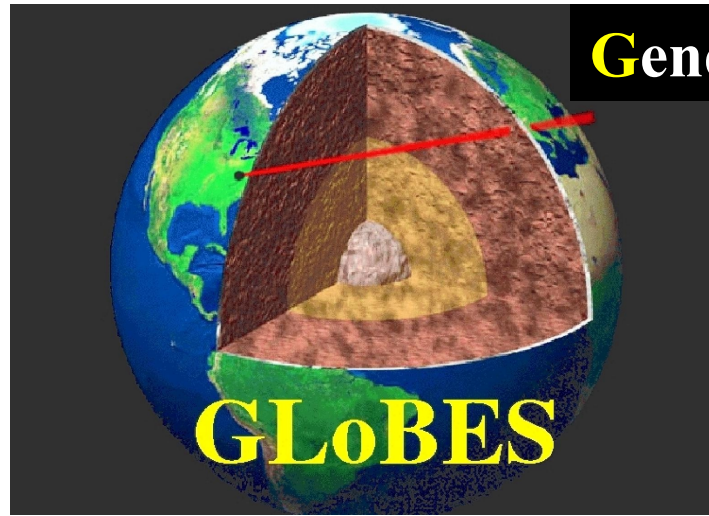
...

Freund, Huber, ML, Winter

Huber, ML, Rolinec, Schwetz, Winter

...

A Powerful Simulation Tool



General Long Baseline Experiment Simulator

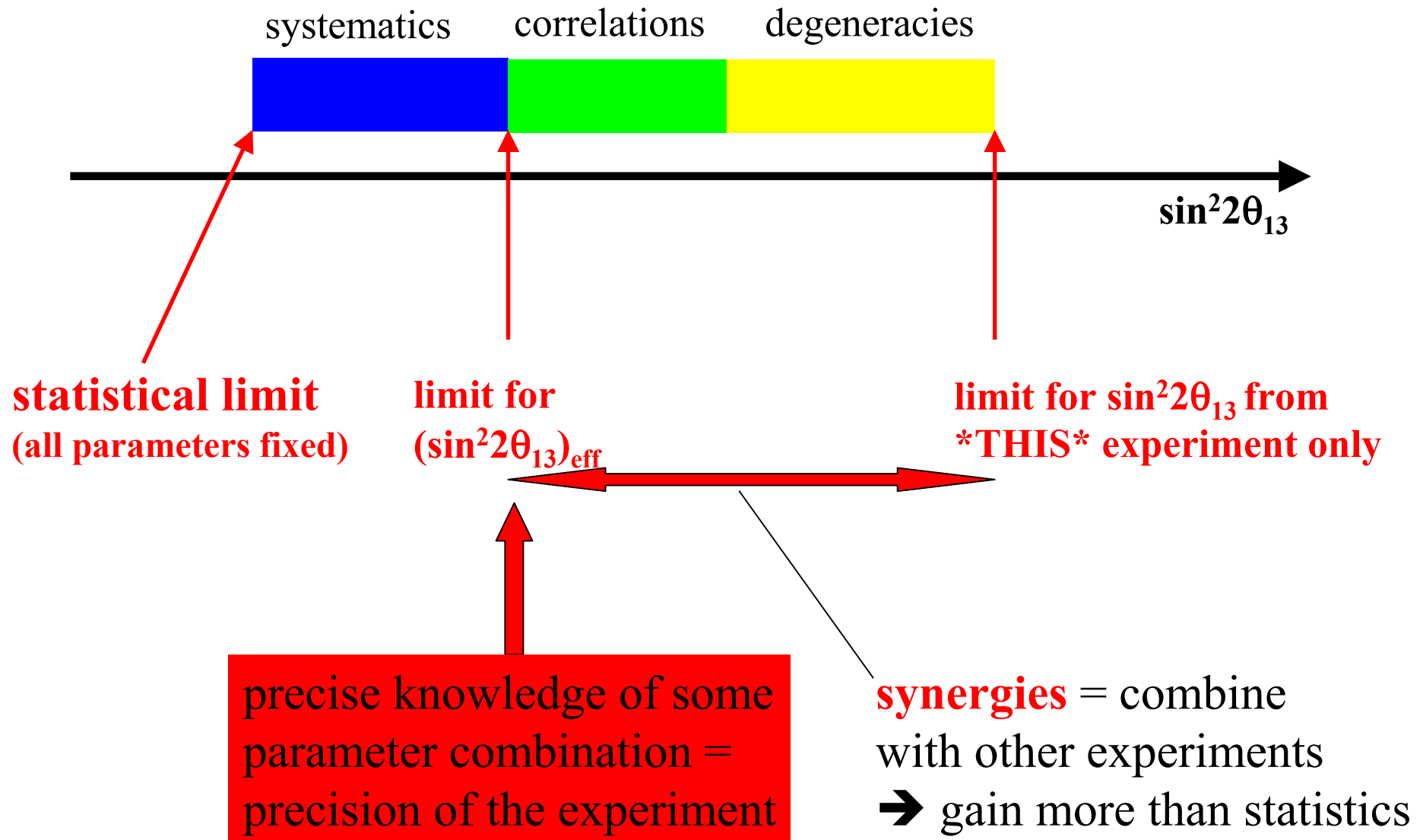
<http://www.ph.tum.de/~globes>

hep-ph/0407333

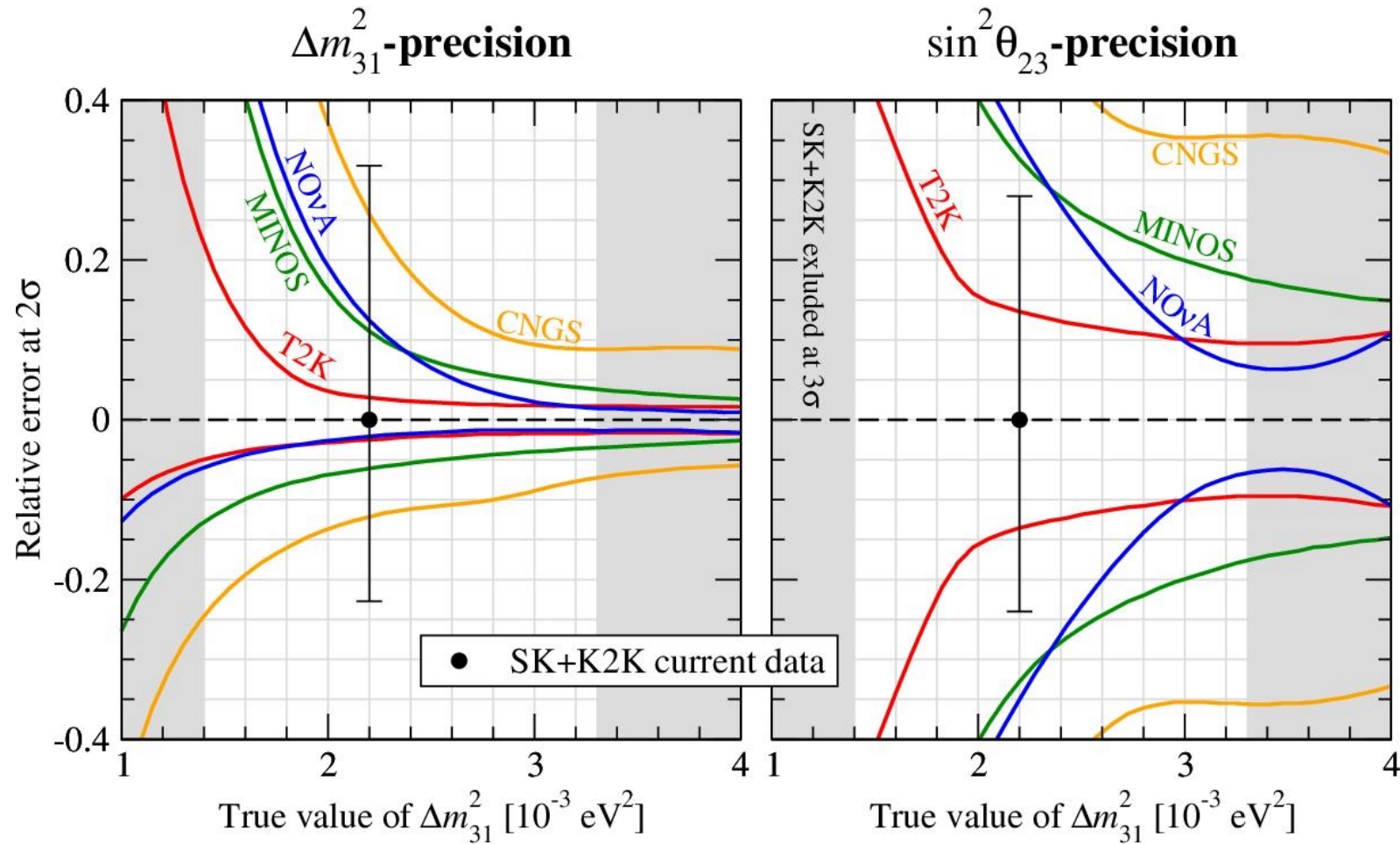
P. Huber, ML, W. Winter

- C-based simulation software (GPL – free, for Linux systems)
- extensive documentation & examples
- 3 phase approach:
 - **AEDL** (Abstract Experiment Definition Language)
 - simulation of an experiment → 3-ν oscillations; scan „true values“
 - analysis → event distributions, ..., sensitivities, ...

Sensitivity Plots

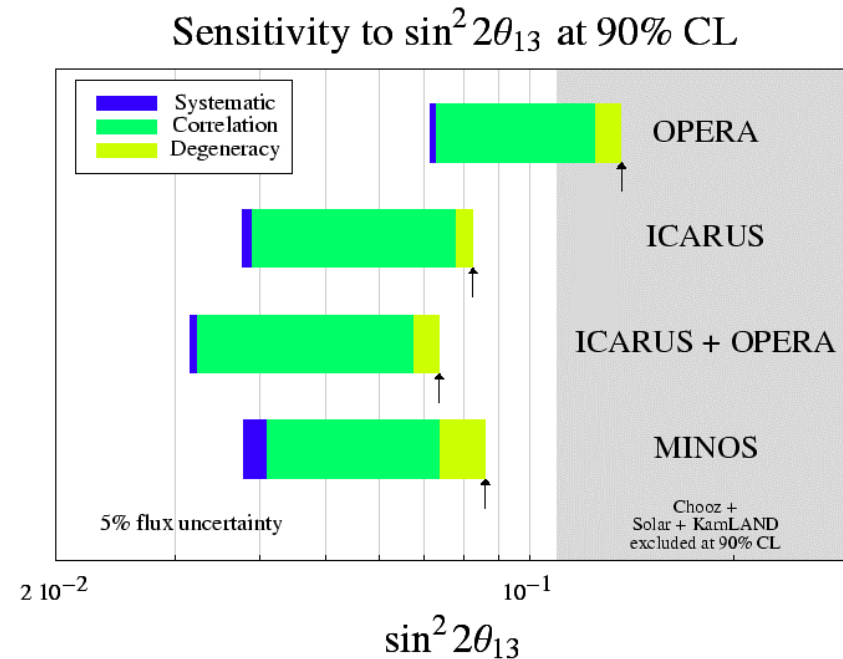
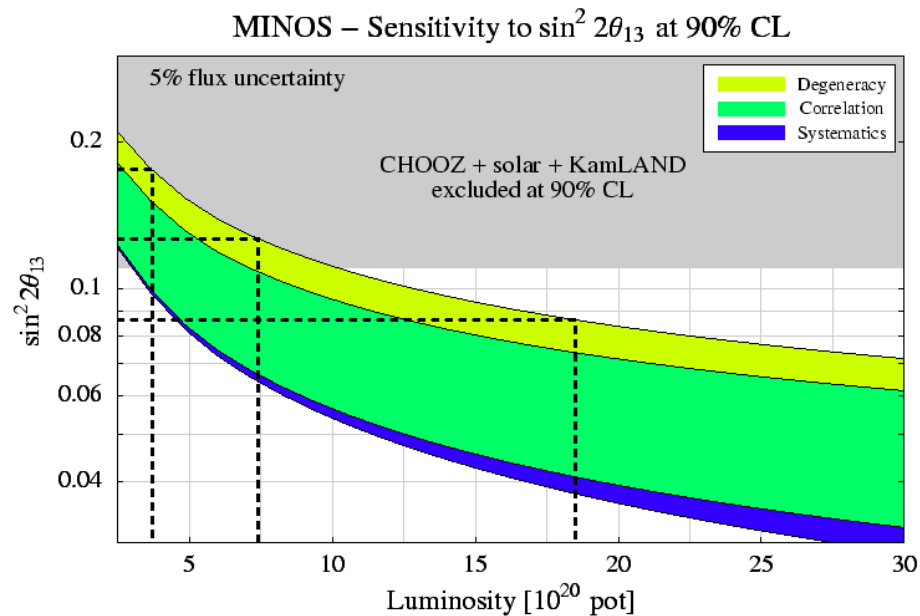


Improvement of Δm_{31}^2 and $\sin^2 \theta_{23}$



Huber, ML, Rolinec, Schwetz, Winter

θ_{13} in the Coming LBL Generation



MINOS sensitivity as a function of time:

- MINOS: $3.7 \cdot 10^{20}$ pot/y
- 1,2,5 years

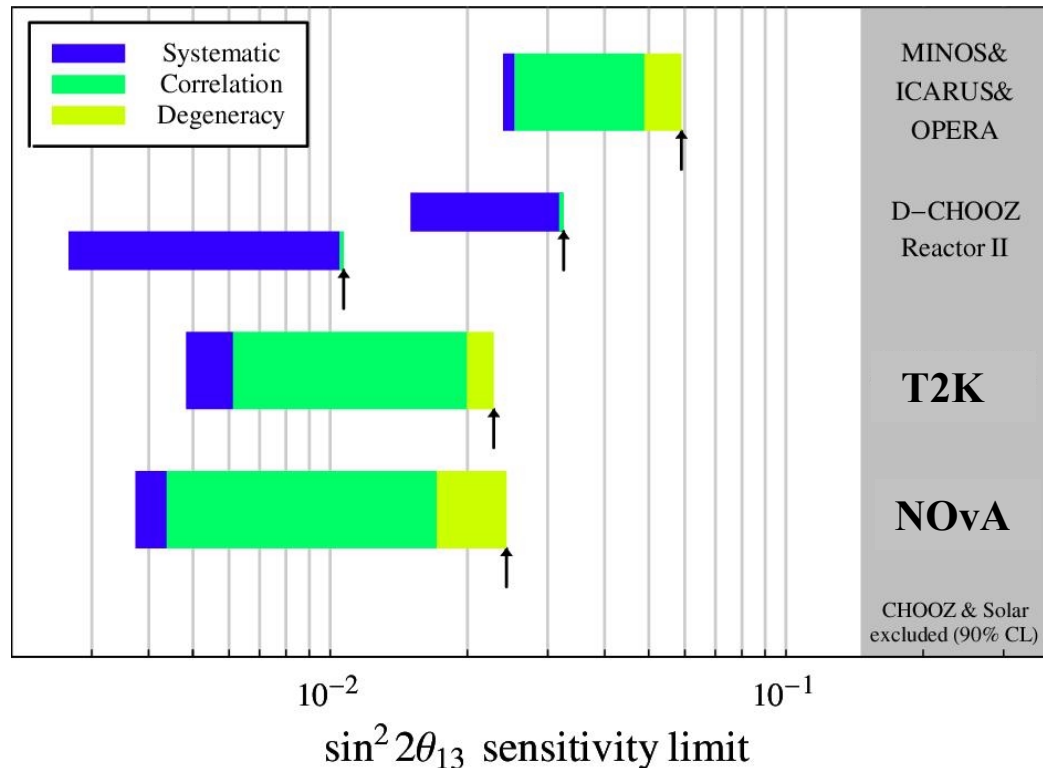
Compare: 5 years, 5% flux uncertainty

- CNGS: $4.5 \cdot 10^{19}$ pot/y

- only modest improvements for θ_{13}
- other objectives...

θ_{13} Sensitivity in the Next Generation

Sensitivity to $\sin^2 2\theta_{13}$ at 90% CL



coming long baseline experiments

1 reactor + 2 detectors

next generation long baseline experiments

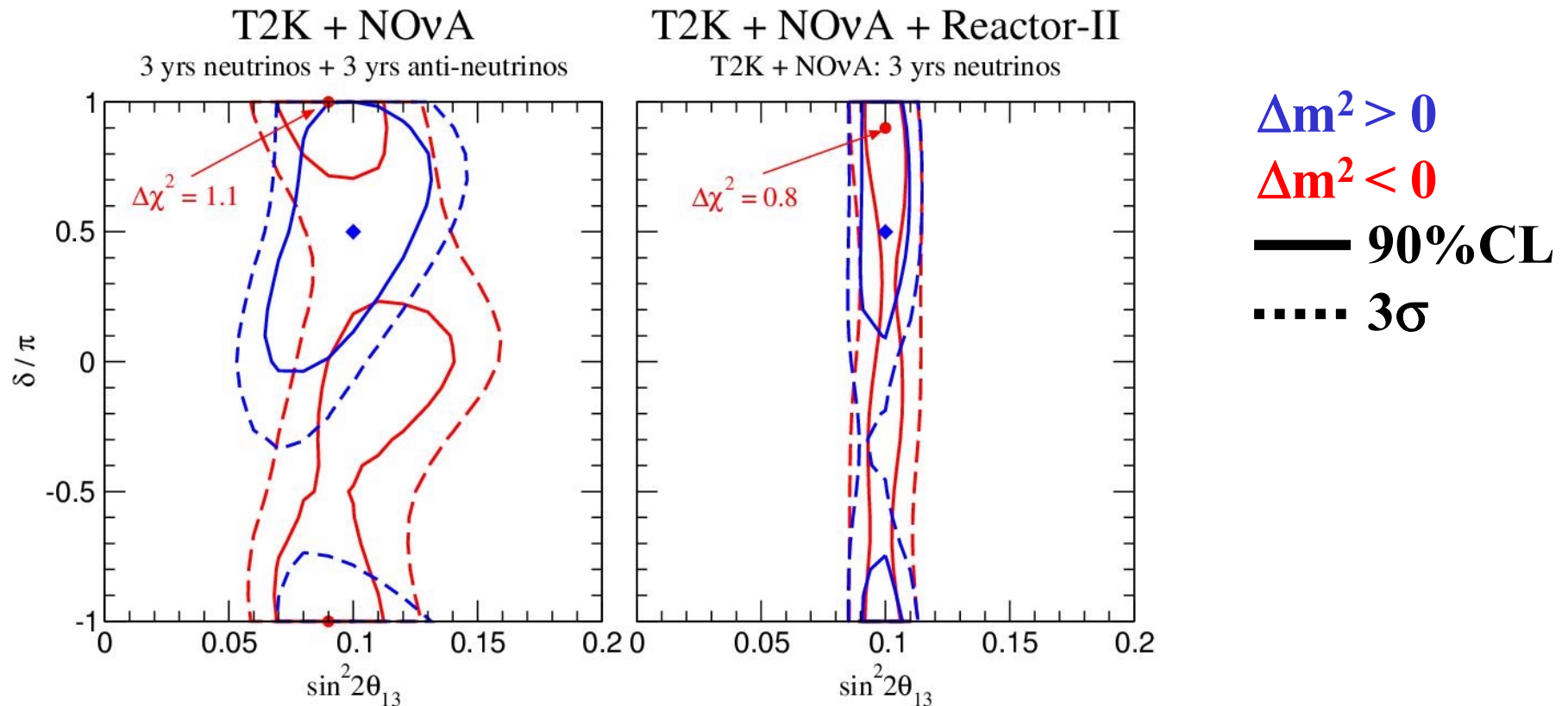
Compare:

- 5 years each
- 5% flux uncertainty

- one order of magnitude improvement for θ_{13}
- synergies between reactor and accelerator experiments
 - reactor anti-neutrinos $\Rightarrow$ only neutrino beams (x-section)
 - reactor: uncorrelated $\theta_{13} \Rightarrow$ combine with beams & resolve correlations
- synergy between beams $\Rightarrow$ NOvA at large baseline $\Rightarrow$ matter effects

Leptonic CP Violation – Best Case

assume: $\sin^2 2\theta_{13} = 0.1$, $\delta = \pi/2 \rightarrow$ combine: T2K+NOvA+Reactor



Huber, ML, Rolinec, Schwetz, Winter

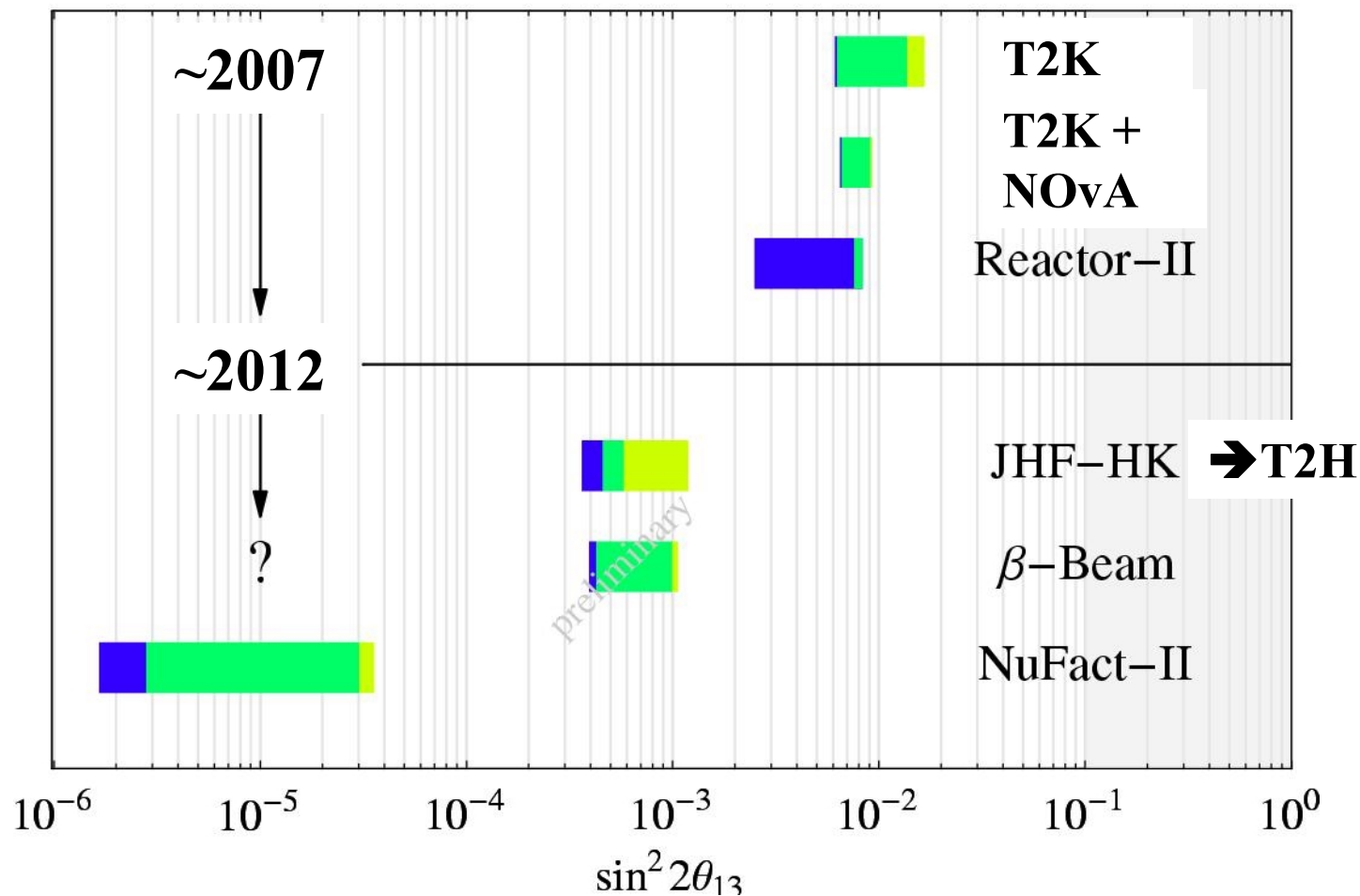
Long Term Perspectives

- superbeams: $E_\nu \approx \text{GeV}$ → large low Z sampling calorimeters $\approx 50 \text{ kt}$
- superbeams, β -beams: $E_\nu \approx \text{GeV}$
 - huge Cerenkov detectors $\approx 1000 \text{ t}$
 - huge liquid Ar detectors $\approx 100 \text{ kt}$
 - huge scintillator detectors $\approx 30 \text{ kt}$
- neutrino factory: $E_\nu \approx 20\text{-}50 \text{ GeV}$
 - large magnetized iron Calorimeters $\approx 40 \text{ kt}$
 - large magnetized liquid Ar detectors $\approx 20 \text{ kt}$
 - large OPERA-like emulsion detectors $\approx 5 \text{ kt}$

$L=3000 \text{ km}$, magnetized iron → wrong sign muons

	P(MW)	$\mu^{\prime}\text{s/year}$	$T_\nu + T_{\bar{\nu}}$ (y)	M(kt)
Neutrino factory I:	0.75	10^{20}	5	10
Neutrino factory II:	4.00	$5.3 \cdot 10^{20}$	8	50

Sensitivity to $\sin^2 2\theta_{13}$ at 90% cl

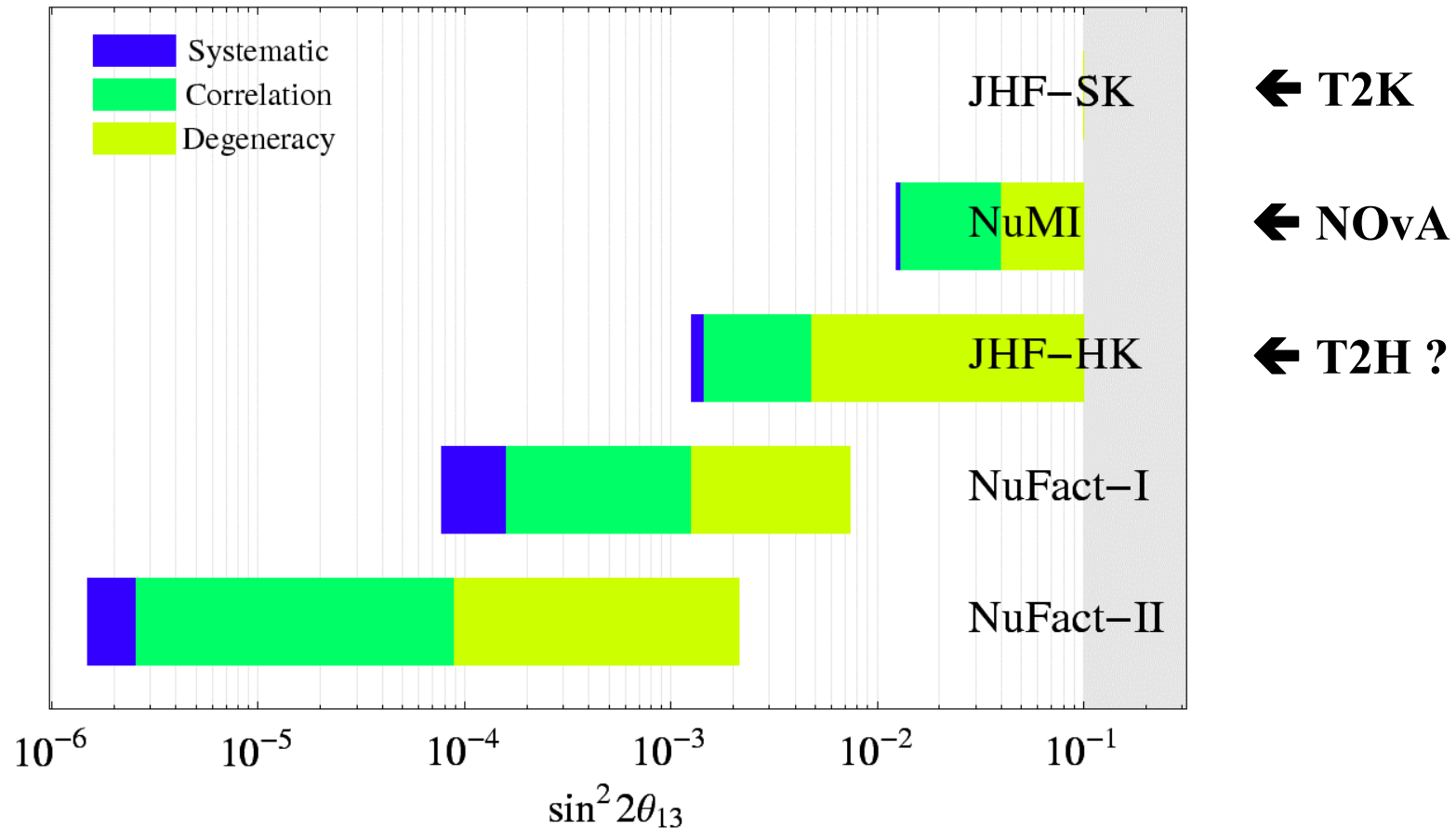


- very powerful
- many good ideas
- ...challenges

➔ further R&D

- different sensitivity reductions by systematics
- correlations & degeneracies lead to severe sensitivity reductions
- break C&D by combining different experiments of comparable potential

Sensitivity to the sign of Δm_{31}^2



- $sign(\Delta m_{31}^2)$ very hard to determine with superbeams
 - degeneracies with δ_{CP} are the main problem
- ⇒ combine experiments!

How to Break Degeneracies & Correlations

Rates only → degeneracies ...broken by

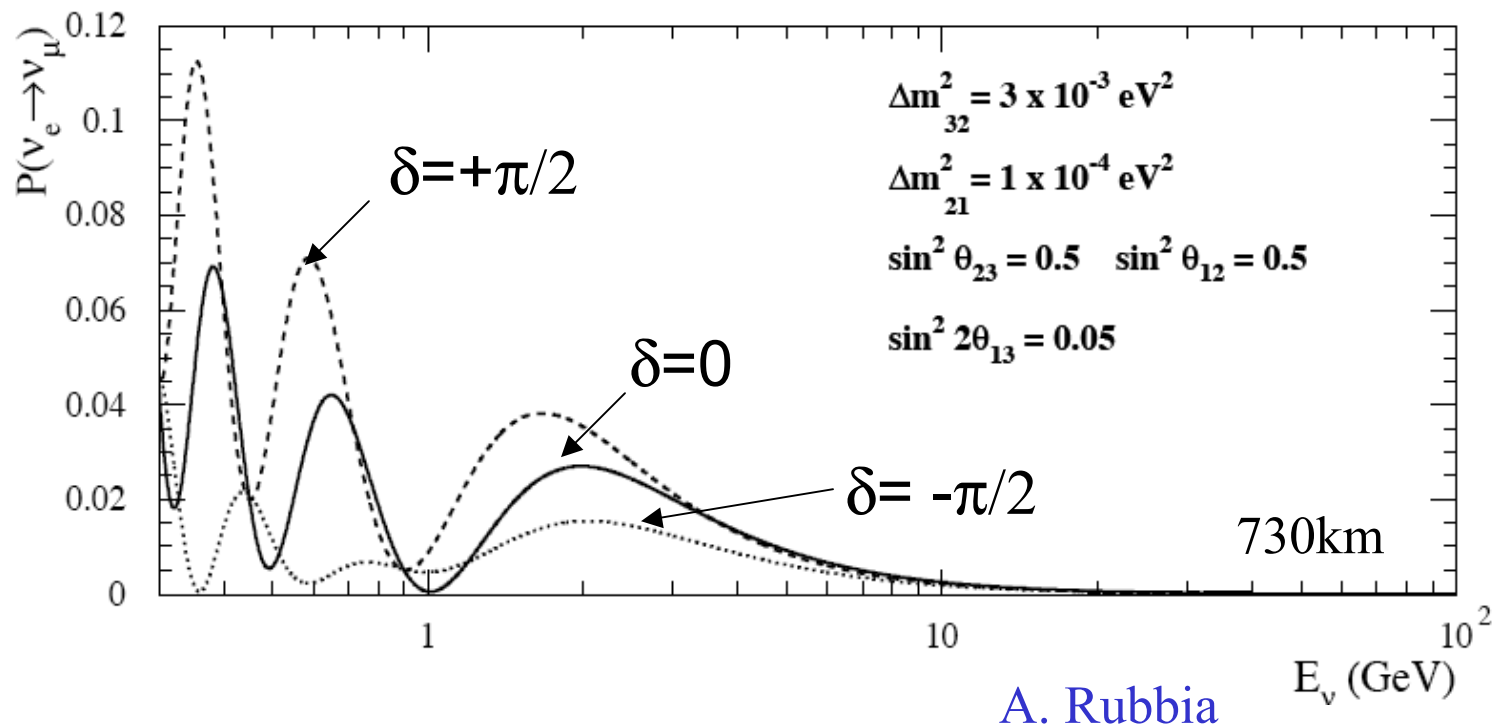
- combination of different oscillation channels
- use different baselines
- combine different energies
- use energy spectrum
- go to „magic baseline“

All degeneracies can in principle be broken

Optimal strategy (physics output / time, money, feasibility)
depends on further R&D

Energy Resolution

Rate based degeneracies have **different energy spectra**



→ use energy resolution to break degeneracies

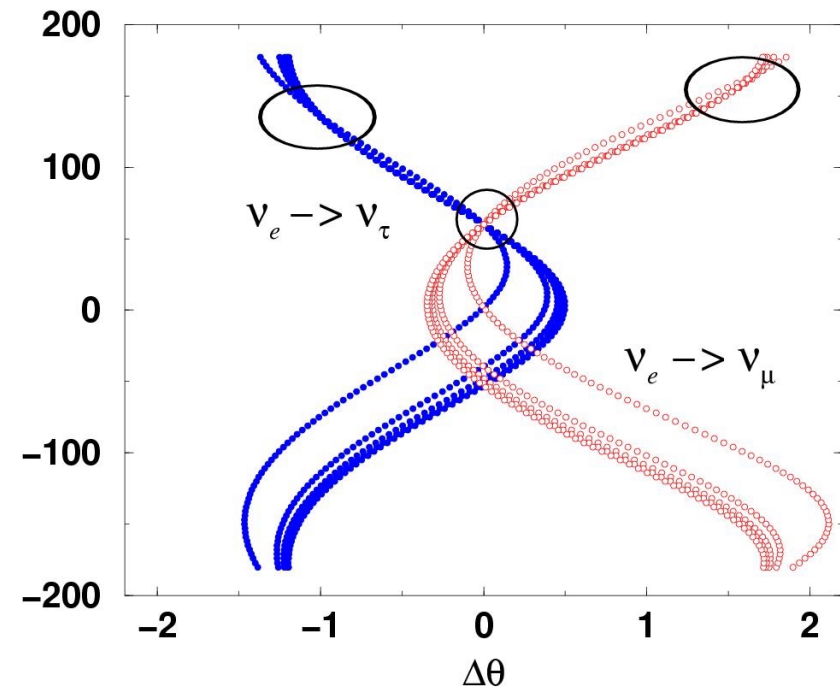
Silver Channels

Neutrino factory:

- golden channel: wrong sign μ 's
- silver channel : τ 's

→ different oscillation probabilities_o

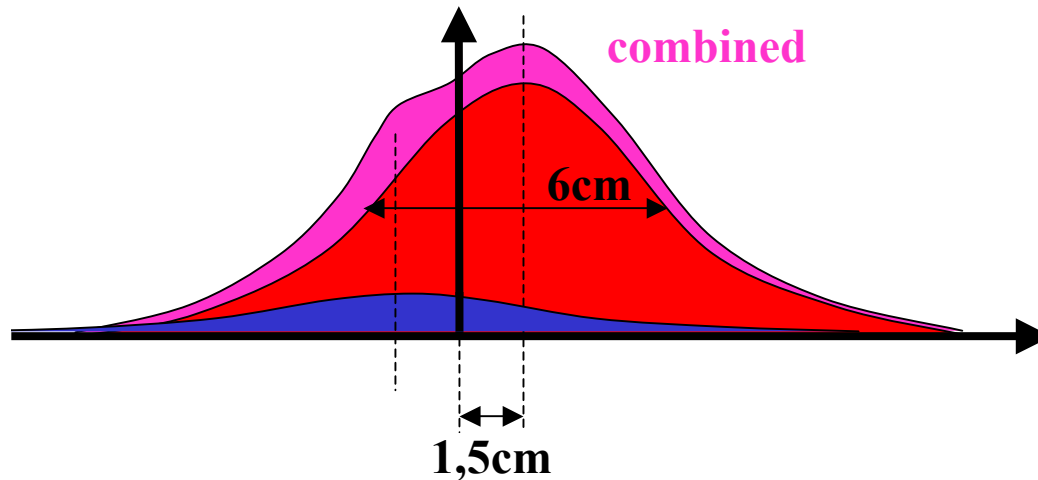
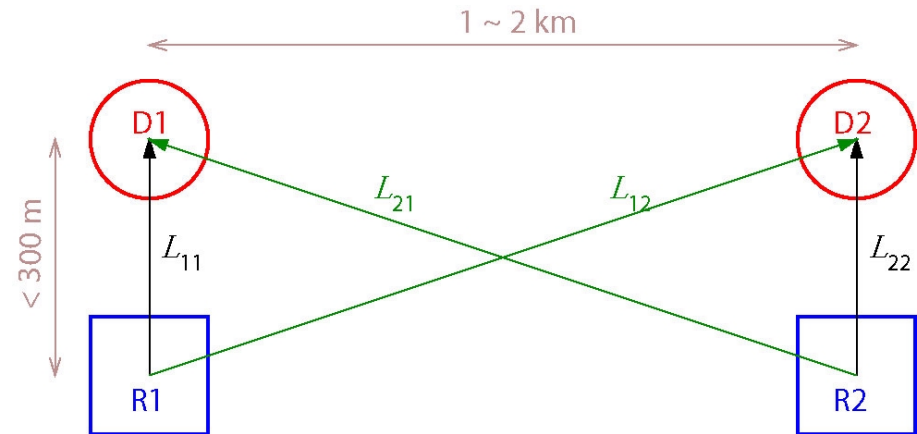
→ break degeneracies!



Donini, Meloni, Migliozi
Autiero, et al.

New Ideas: R2D2 - Reactor Experiments

- **Symmetric reactors, detectors:**
 - R1, R2, D1, D2 - may be different
 - $L_{11}=L_{22}$ and $L_{12}=L_{21}$
- **Separate events from R_1 and R_2**
 - R1 and R2 on/off times
 - Neutron displacement
- **Simplest case: 1d line-up**



High statistics:

- ➔ precise statistical separation
- ➔ $N_{11}, N_{21}, N_{12}, N_{22}$
- ➔ self-calibration: $N_{11}/N_{21} = N_{22}/N_{12}$
- ➔ $\frac{N_{11} * N_{22}}{N_{21} * N_{12}} = \frac{r^4}{R^4} \quad \longleftrightarrow \text{oscillation}$
- ➔ stable against size, backgrounds, ...
- ➔ **factor 2 – 10 compared to R1D2**

Theoretical Implications of Precision

Precision allows to identify / exclude:

- special angles: $\theta_{13}=0^\circ$, $\theta_{23}=45^\circ$, ... $\leftrightarrow$ discrete f. symmetries?
- special relations: $\theta_{12} + \theta_C = 45^\circ$? $\leftrightarrow$ lepton–quark relation?
- quantum corrections $\leftrightarrow$ renormalization group evolution

Provides also tests of:

- **MSW effect** (coherent forward scattering and matter profiles)
- 3 neutrino unitarity $\leftrightarrow$ sterile neutrinos with small mixings
- neutrino decay
- decoherence
- NSI
- MVN, ...

The Value of Precision for θ_{13}

- models for masses & mixings
- input: Known masses & mixings
→ distribution of θ_{13} „predictions“
- θ_{13} often close to experimental bounds
→ motivates new experiments
→ θ_{13} controls 3-flavour effects
like CP-violation

for example: $\sin^2 2\theta_{13} < 0.01$ →

physics question: small θ_{13}

→ numerical coincidence

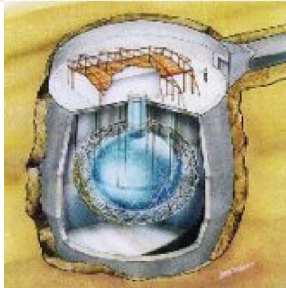
→ systematic (symmetry,...)

• how small?

• precision!

Reference	$\sin \theta_{13}$	$\sin^2 2\theta_{13}$
$SO(10)$		
Goh, Mohapatra, Ng [40]	0.18	0.13
$O(10)$		
Asaka, Buchmüller, Covi [41]	0.1	0.04
$SO(10) + \text{flavor symmetry}$		
Babu, Pati, Wilczek [42]	$5.5 \cdot 10^{-4}$	$1.2 \cdot 10^{-6}$
Blazek, Raby, Iobe [43]	0.05	0.01
Kitano, Mimura [44]	0.22	0.18
Albright, Barr [45]	0.014	$7.8 \cdot 10^{-4}$
Machawa [46]	0.22	0.18
Ross, Velasco Sevilla [47]	0.07	0.02
Chen, Mahanthappa [48]	0.15	0.09
Raby [49]	0.1	0.04
$SO(10) + \text{texture}$		
Buchmüller, Wyler [50]	0.1	0.04
Bando, Obara [51]	0.01 .. 0.06	$4 \cdot 10^{-4}$.. 0.01
Flavor symmetries		
Grimus, Lavoura [52, 53]	0	0
Grimus, Lavoura [52]	0.3	0.3
Babu, Ma, Valle [54]	0.14	0.08
Kuchimanchi, Mohapatra [55]	0.08 .. 0.4	0.03 .. 0.3
Ohlsson, Seidl [56]	0.07 .. 0.14	0.02 .. 0.08
King, Ross [57]	0.2	0.15
Textures		
Honda, Kaneko, Tanimoto [58]	0.08 .. 0.20	0.03 .. 0.15
Lebed, Martin [59]	0.1	0.04
Bando, Kaneko, Obara, Tanimoto [60]	0.01 .. 0.05	$4 \cdot 10^{-4}$.. 0.01
Ibarra, Ross [61]	0.2	0.15
$3 \times 2 \text{ see-saw}$		
Appelquist, Piai, Shrock [62, 63]	0.05	0.01
Freempton, Glashow, Yanagida [64]	0.1	0.04
Mei, Xing [65] (normal hierarchy)	0.07	0.02
(inverted hierarchy)	> 0.006	$> 1.6 \cdot 10^{-4}$
Anarchy		
de Gouvêa, Murayama [66]	> 0.1	> 0.04
$\text{Renormalization group enhancement}$		
Mohapatra, Parida, Rajasekaran [67]	0.08 .. 0.1	0.03 .. 0.04

Mass Models & Renormalization



low energies:

- small masses
- large mixings

renormalization group running



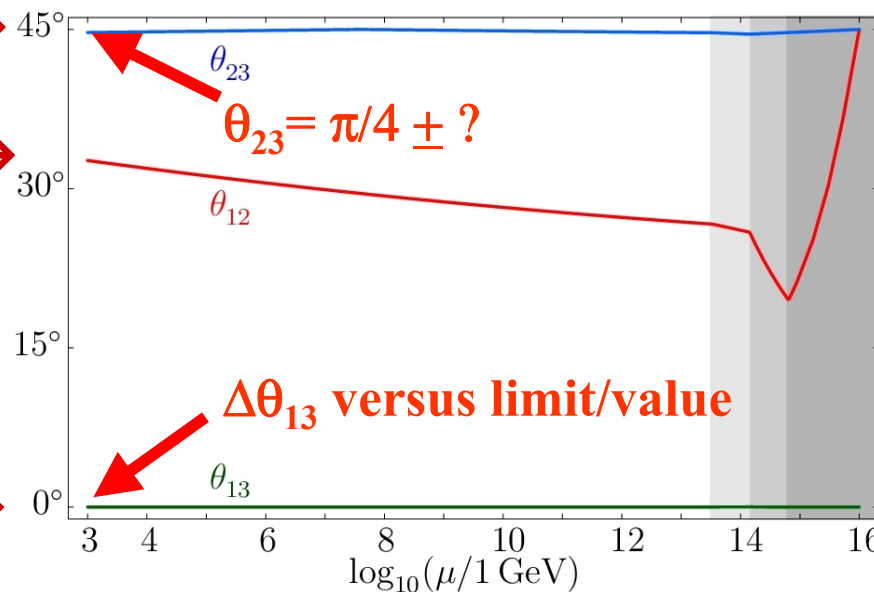
High energies:

- mass models
- flavour-symmetries
- GUT-models, ...

atmospheric $\rightarrow$ 45° $\leftarrow$ bi-maximal

solar $\rightarrow$

reactor $\rightarrow$



MSSM example:
Antusch, Kersten, ML, Ratz

$\leftarrow$ small

The Interplay of Various Topics

Standard Model extensions

flavour symmetries

m_{LL}	Scheme	$0\nu\beta\beta$ -decay
$\frac{m}{2} \begin{pmatrix} 0 & 0 & 0 \\ 0 & 1 & 1 \\ 0 & 1 & 1 \end{pmatrix} + \text{corr.}$	hierarchical, normal	small
$\frac{m}{\sqrt{2}} \begin{pmatrix} 0 & 1 & 1 \\ 1 & 0 & 0 \\ 1 & 0 & 0 \end{pmatrix} + \text{corr.}$	hierarchical, inverted	small
$m \begin{pmatrix} 1 & 0 & 0 \\ 0 & \frac{1}{2} & \frac{1}{2} \\ 0 & \frac{1}{2} & \frac{1}{2} \end{pmatrix} + \text{corr.}$	hierarchical, inverted	large
$m \begin{pmatrix} 0 & \frac{1}{\sqrt{2}} & \frac{1}{\sqrt{2}} \\ \frac{1}{\sqrt{2}} & \frac{1}{2} & \frac{1}{2} \\ \frac{1}{\sqrt{2}} & \frac{1}{2} & \frac{1}{2} \end{pmatrix} + \text{corr.}$	degenerate	small
$m \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix} + \text{corr.}$	degenerate	large

leptogenesis mechanisms
supernovae
nucleosynthesis
structure formation
...

renormalization group

mass spectrum, mixings, CP-phases, LVF, $0\nu 2\beta\beta$ -decay, ...

→ ν -parameters extremely valuable

→ long term: most precise flavour info

Conclusions

- very precise neutrino oscillation parameters can be obtained
- long term: most precise flavour information
- $\sin^2 2\theta_{13}$, $\text{sign}(\Delta m^2)$ and CP phase should be measured
- unique impact on our understanding of flavour
 - model building (symmetries, GUTs, mixing angle relations,...)
 - quantum corrections
 - limits on 3v unitarity, decay, decoherence
- tests also
 - coherent forward scattering, matter profiles
 - NSI
 - MVN, extra dimensions, ...

→ very promising program!