

LHC and neutrino physics

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See also: Senjanovic and Mohapatra

The origin of neutrino masses

- * A compelling understanding:

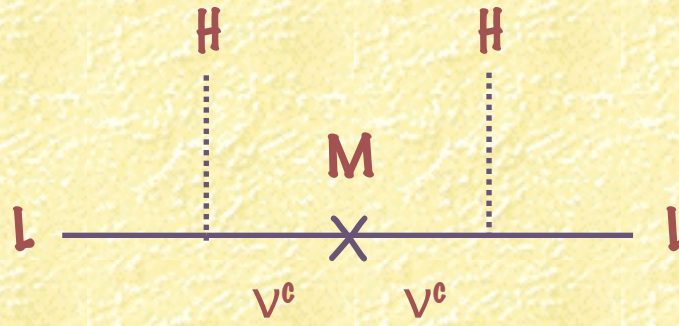
$$\mathcal{L}_{\text{SM}}^{\text{eff}} = \mathcal{L}_{\text{SM}}^{\text{ren}} + h_{ij} \frac{L_i L_j H H}{\Lambda} + \dots$$

Lepton number is an accidental symmetry of the SM

$$m_\nu = h v \times \frac{v}{\Lambda}$$

$$\Lambda \sim 0.5 \times 10^{15} \text{ GeV} \left(\frac{0.05 \text{ eV}}{m_\nu} \right) h$$

- * (close to $M_{\text{GUT}} \approx 2 \times 10^{16} \text{ GeV}$)
- * $\mathcal{L}_{\text{eff}}$ is sensitive to $\Lambda = \Lambda_L \approx 10^{15} \text{ GeV}$, $\Lambda_B > 4 \times 10^{15} \text{ GeV} \gg \text{TeV}$
(L- and B-violating operators = window to the GUT scale)
- * No or very small L, B violation at TeV scale



$$\frac{h}{\Lambda} LLHH$$

$$\frac{h}{\Lambda} \rightarrow -\lambda^T \frac{1}{M} \lambda$$

$$m_\nu = -m_D^T \frac{1}{M} m_D$$

$$\Lambda \lesssim \text{TeV?}$$

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* $\Lambda \lesssim \text{TeV}$ needs $h < 10^{-11}$: why?

- L is “exactly” conserved: $h = 0$
- L is violated by interactions at $\Lambda \lesssim \text{TeV}$ and $h \approx 10^{-11}$ – 10^{-13}

Lepton number is conserved (at first)

- * Neutrino masses need a 'right-handed' neutrino (Dirac neutrinos)
- * In SM: $\lambda \bar{\nu}_R L H \rightarrow m_\nu = \lambda v$ (like the other fermions)
- * Need $\lambda < 10^{-11}$: why?
 - L is conserved + $\lambda \bar{\nu}_R L H$ forbidden by a symmetry, e.g. because it is charged under a U(1) symmetry:

$$\lambda \bar{\nu}_R L H \rightarrow \lambda \left(\frac{\phi}{M} \right)^n \bar{\nu}_R L H, \quad \lambda_{\text{eff}} = \lambda \left(\frac{\langle \phi \rangle}{M} \right)^n$$

[Chacko Hall Okui Oliver ph/0312267
Chacko, Hall Oliver Perelstein ph/0405067
Davoudiasl Kitano Kribs Murayama
ph/0502176]

- interesting (model dependent) consequences for cosmology (and LSND)
- no consequences for LHC: $\frac{\langle H \rangle}{M} \sim \frac{m_\nu}{\langle \phi \rangle} \sim g_{\phi \nu \nu^c} \lesssim 10^{-5}$ (BBN) (n=1)

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 - L is conserved + λ originates in extra-dimensions
 - ν_R lives in the flat bulk of large extra dimensions:
$$\lambda_{\text{eff}} = \frac{\lambda}{(2\pi R M_*)^{\delta/2}} = \lambda \frac{M_*}{M_{\text{Pl}}}$$

[Arkani-Hamed et al. ph/9811448
Dienes Dudas Gherghetta ph/9811428]
 - ν_R and L are localized in distant points of a (warped) extra dimension:
$$\lambda \propto e^{-(\text{superposition of the wave functions})}$$

[Grossman Neubert ph/9912408]

Large extra-dimensions

- * 5D $\nu_R \leftrightarrow$ 4D $(\nu_R)_n$
- * $M_n \approx n/R$ (large n)
- * Brane-bulk mixing: $m \approx \lambda_{\text{eff}} \langle H \rangle$ (for each active neutrinos)

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$$\nu_i = U_{ik} \hat{\nu}_k + \frac{m_i}{M_n} N_n$$

In the presence
of bulk mass terms

[Lukas Ramond R Ross
ph/0008049, ph/0011295]

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Standard
3 (mainly) active
neutrino mixing

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[Lukas Ramond R Ross
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- * Signatures from the interaction $\frac{\lambda}{M_*} \overline{\nu_R} L H$:
 - e^+e^- colliders
 - H decay

Signatures

- * $h \rightarrow \bar{\nu}_L(\nu_R)_n$ proceeds through $\lambda_{\text{eff}}^2 \sim \left(\frac{M_*}{M_{\text{Pl}}}\right)^2$
but is enhanced by the $(M_* R)$ final states ($D=5$)

$$\frac{\Gamma(H^0 \rightarrow \bar{\nu}_L \nu_R)}{\Gamma(H^0 \rightarrow b\bar{b})} \sim 10^{6-\delta} \frac{m^2}{10^{-5} \text{ eV}^2} \left(\frac{m_H}{100 \text{ GeV}}\right)^\delta \left(\frac{\text{TeV}}{M_*}\right)^{2+\delta} \quad (D=4+\delta)$$

The invisible decay can easily dominate

[Arkani-Hamed et al. ph/9811448
Martin Wells ph/9903259]

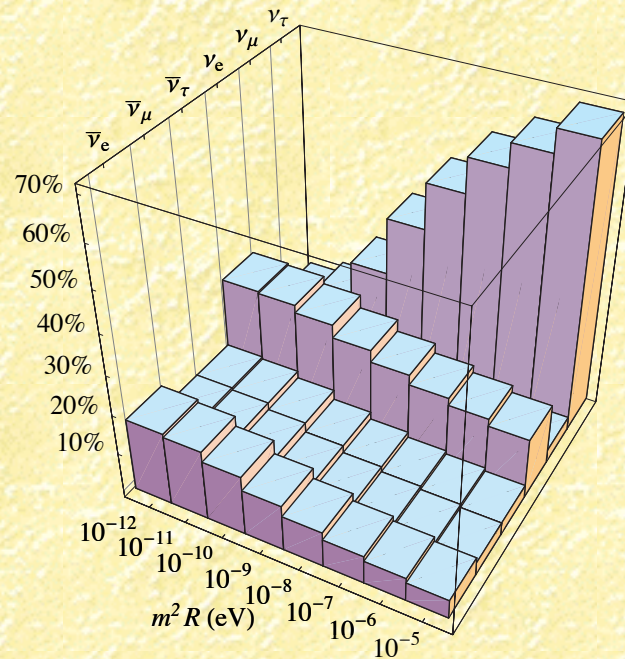
- * For a 2HDM (type II):

$$\frac{\Gamma(H^- \rightarrow \bar{\tau}_L \nu_R)}{\Gamma(H^- \rightarrow \bar{\tau}_R \nu_L)} \sim \cot^4 \beta \left(\frac{\lambda}{\lambda_\tau}\right)^2 \left(\frac{m_H}{M_*}\right)^\delta$$

[Agashe Deshpande Wu ph/0006122]

Large extra-dimensions and SN

- * Neutrino KK = dangerous invisible energy loss channel (resonantly enhanced)
- * A double feedback mechanism weakens the naive bounds on $m^2 R$ by 5 orders of magnitude to $5 \cdot 10^{-5} \text{eV}$
- * The neutrino spectrum is modified



KK neutrinos mix
predominantly with ν_τ

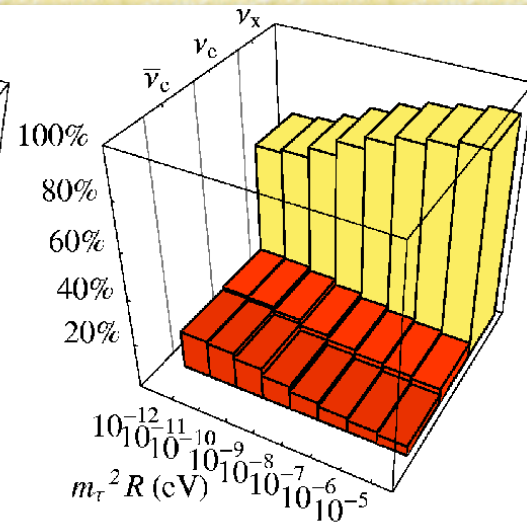
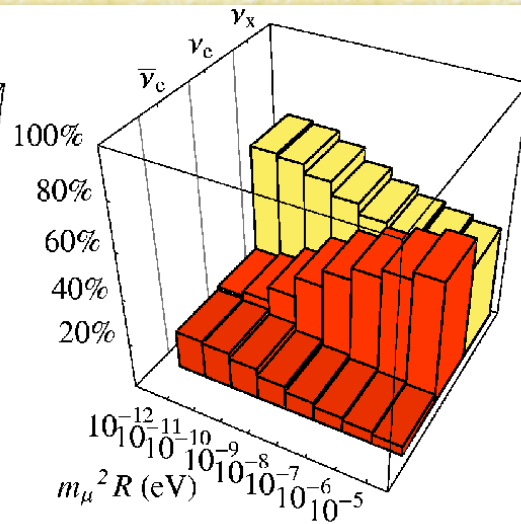
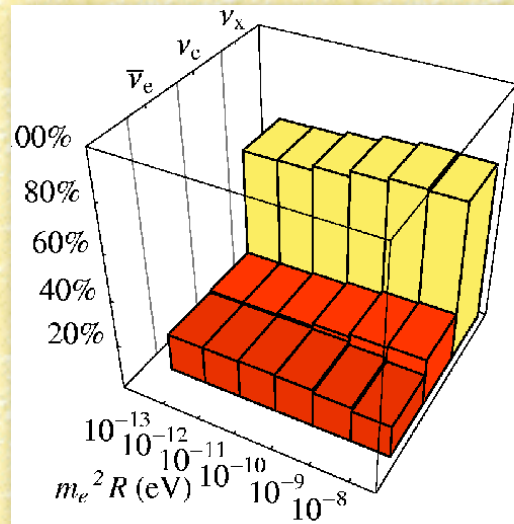
[Cacciapaglia Cirelli Yin R ph/0209063
Cacciapaglia Cirelli R ph/0302246]

ν_R mixes predominantly with

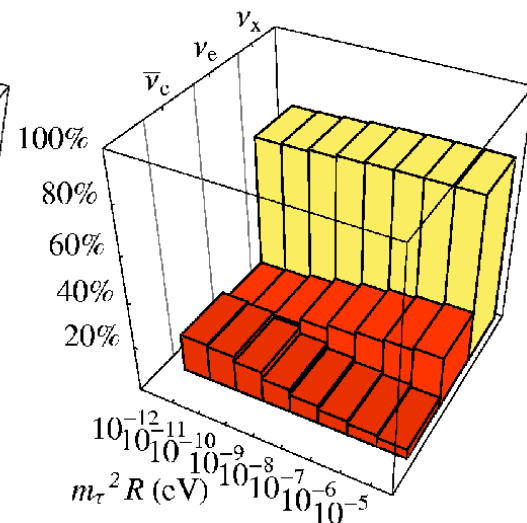
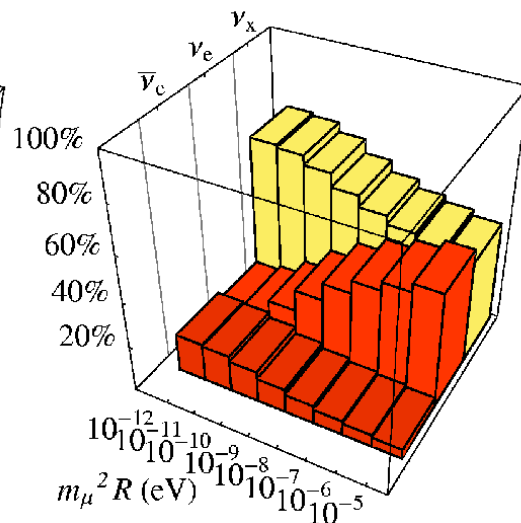
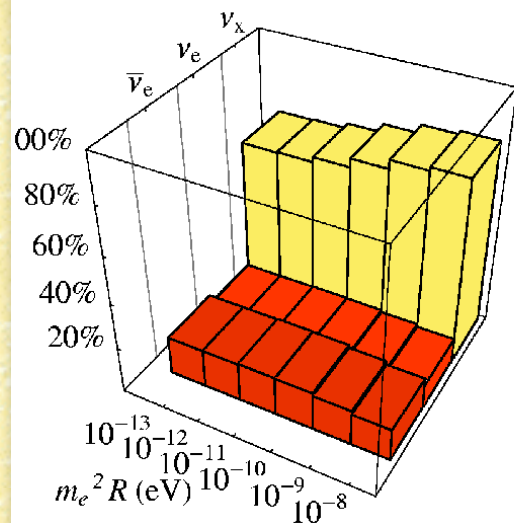
ν_e

ν_μ

ν_τ



$S^2_{13} < 10^{-6}$



$S^2_{13} > 0.5 \cdot 10^{-3}$

(normal hierarchy)

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$$m_\nu = h v \times \frac{v}{\Lambda}$$

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- * Need $h \approx 10^{-11}-10^{-13}$: why?

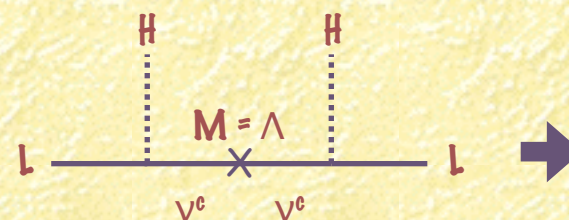
LLHH forbidden by a symmetry

$$h \frac{LLHH}{\Lambda} \rightarrow h \left(\frac{\phi}{M} \right)^n \frac{LLHH}{\Lambda}, \quad h_{\text{eff}} = h \left(\frac{\langle \phi \rangle}{M} \right)^n$$

as before: consequences for cosmology but not for LHC

- * Assume $h \approx 10^{-11}-10^{-13}$, $\Lambda \approx \text{TeV}$: what is the origin of LLHH?

- TeV-scale see-saw



$$\frac{h}{\Lambda} \rightarrow -\lambda^T \frac{1}{M} \lambda$$

$$m_\nu = -m_D^T \frac{1}{M} m^D$$

- Supersymmetry

TeV-scale see-saw

* ν_R with $M_R \sim \text{TeV}$

* Probe ν_R through $\lambda \bar{\nu}_R L H$: $m_\nu = -m_D^T \frac{1}{M} m_D$, $m_D = \lambda \langle H \rangle$

* $M_R \sim \text{TeV} \Rightarrow \lambda = \frac{m_D}{M_R} \approx 2 \times 10^{-7} \left(\frac{m_\nu}{0.05 \text{ eV}} \right)^{1/2} \left(\frac{\text{TeV}}{M_R} \right)^{1/2}$ too small for LHC

* Unless $\lambda \gg 10^{-7}$ + cancellations in $m_\nu = -m_D^T \frac{1}{M} m_D$ (2 or more ν_R 's)

• “magical”, e.g.:

$$m_{nj}^D = \alpha_n \beta_j m_0, \quad M_R = \text{Diag}(M_1 \dots M_n), \quad \sum_n \alpha_n^2 M_n = 0 \quad [\text{Buchmüller Greub NPB363}]$$

$m_\nu = 0 + \text{corrections}$

• natural, e.g.:

$L_e, L_\mu, L_\tau, (\nu_R)_1 \equiv N$ have $L = 1$, $(\nu_R)_2 \equiv N'$ has $L = -1$

then $-\mathcal{L}_{\text{mass}}^\nu = (L_e, L_\mu, L_\tau, N, N') \mathcal{M} (L_e, L_\mu, L_\tau, N, N')^T$

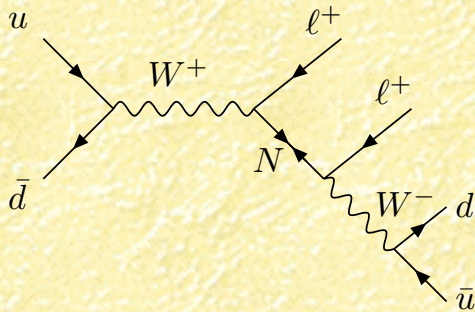
$$\mathcal{M} = \begin{pmatrix} & & & 0 & m_e \\ & \mathbf{0} & & 0 & m_\mu \\ & & & 0 & m_\tau \\ 0 & 0 & 0 & 0 & M \\ m_e & m_\mu & m_\tau & M & 0 \end{pmatrix}$$

- rank = 2: 3 massless neutrinos independently of the size of m_i

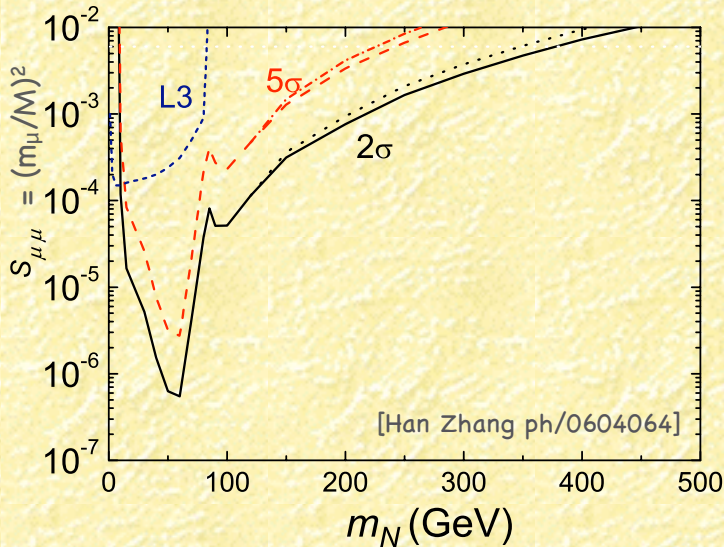
- $\nu_i = U_{ik} \hat{\nu}_k + \frac{m_i}{M} N$

* Constraints: $(m_e/M)^2 < 0.006$, $(m_\mu/M)^2 < 0.003$, $(m_\tau/M)^2 < 0.003$

* LNV at LHC (!):
 $qq' \rightarrow \mu^\pm \mu^\pm W^\mp$
 cleanly probes m_μ/M



* No connection with m_ν



[Nardi Roulet Tommasini NPB386 (1992),
 ph/9402224, ph/9409310
 Bergmann Kagan, ph/9803305]

[Dicus Karatas Roy PRD44 (1991)
 Datta Guchait Pilaftsis ph/9311257
 Almeida et al ph/0002024
 Ali Borisov Zamorin ph/0104123
 Panella et al ph/0107308
 Han Zhang ph/0604064
 Aguila Aguilar-Saavedra Pittau ph/0606198
 Bar-Shalom et al ph/0608309
 Atwood Bar-Shalom Soni ph/0701005
 Bray Lee Pilaftsis ph/0702294]

- * Need $h \approx 10^{-11}$ – 10^{-13} : why? Extra structure:

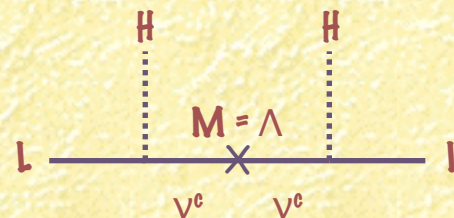
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- Supersymmetry

(R_P violating) supersymmetry

- * Supersymmetry does not guarantee (accidental) L (or B) conservation, unlike the SM: $H_d \approx \tilde{L}_i$

$$W = \lambda_{ij}^U u_i^c Q_j H_u + \lambda_{ij}^D d_i^c Q_j H_d + \lambda_{ij}^E e_i^c L_j H_d + \mu H_u H_d \\ + \lambda_{ijk}'' u_i^c d_j^c d_k^c + \lambda_{ijk}' L_i Q_j d_k^c + \lambda_{ijk} L_i L_j e_k^c + \mu_i H_u L_i$$

$$\mathcal{L}_{\text{soft}} = A_{ij}^U \tilde{u}_i^c \tilde{Q}_j H_u + A_{ij}^D \tilde{d}_i^c \tilde{Q}_j H_d + A_{ij}^E \tilde{e}_i^c \tilde{L}_j H_d + B\mu H_u H_d \\ + A_{ijk}'' \tilde{u}_i^c \tilde{d}_j^c \tilde{d}_k^c + A_{ijk}' \tilde{L}_i \tilde{Q}_j \tilde{d}_k^c + A_{ijk} \tilde{L}_i \tilde{L}_j \tilde{e}_k^c + (B\mu)_i H_u \tilde{L}_i$$

$$+ \tilde{m}_Q^2 \tilde{Q}^\dagger \tilde{Q} + (\tilde{m}_i^2 H_d^\dagger \tilde{L}_i + \text{h.c.}) + \text{gaugino masses}$$

- * L and B violating terms controlled by $R_P = (-1)^{3(B-L)+2s}$
- * A small R_P breaking:
 - induces $(h_{ij}/\Lambda) L_i L_j H H$, with $\Lambda = \tilde{m}$, $h \leftrightarrow$ small R_P breaking parameters
 - makes the LSP unstable (could be any susy partner)

Bilinear R_P violation

[Hall Suzuki NPB231 (1984), Lee PLB138 (1984), NPB246 (1984), Dawson NPB261 (1985), Hempfling ph/9511288, Nilles Polonsky ph/9606388, Kaplan Nelson ph/9901254, Chun Kang ph/9909429, Hirsch Diaz Porod Romao Valle ph/0004115, Joshipura Vempati ph/9808232, Takayama Yamaguchi ph/9910320, Joshipura Vaidya Vempati ph/0203182, Chun Jung Park ph/0211310]

- * In an appropriate basis for $L_\alpha = (H_d, L_e, L_\mu, L_\tau)$:
 - No R_P -violating **trilinear** terms
 - $W \supset \mu H_u H_d + \mu_i H_u L_i$, $\mathcal{L}_{\text{soft}} \supset B\mu H_u H_d + (B\mu)_i H_u L_i +$
- * Predictive + might follow from spontaneous R_P breaking

[Aulakh Mohapatra PLB119, Ellis et al PLB150, Ross Valle PLB151, Chikashige Mohapatra Peccei PLB98]

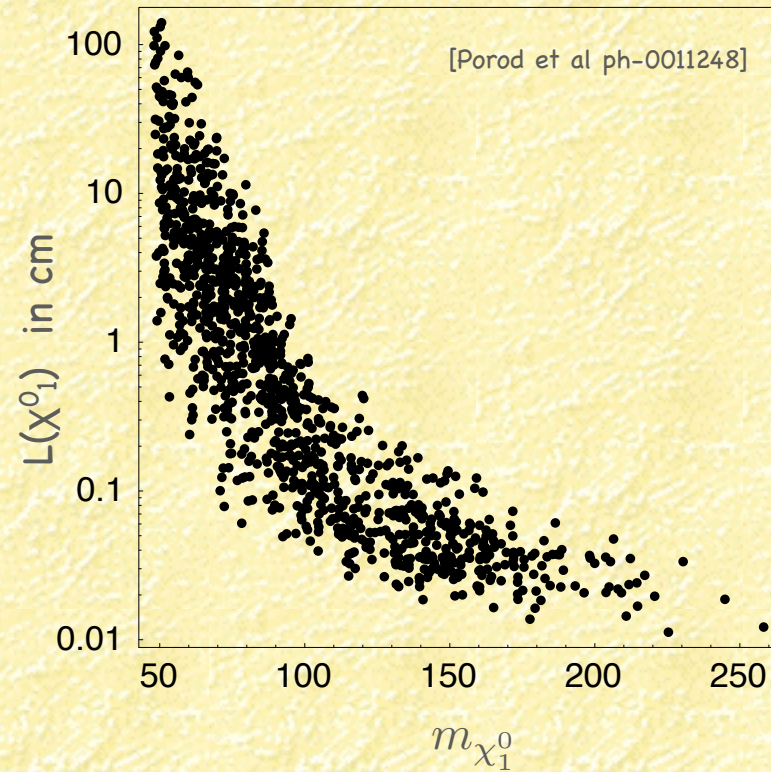
- * $\langle H_u \rangle_2 = v_u$, $\langle H_d \rangle_1 = v_d$, $\langle \tilde{L}_i \rangle_2 = v_i \rightarrow$ neutrino-neutralino mixing $\rightarrow m_\nu$
- * Tree level: $\frac{h_{ij}}{\Lambda} \approx \frac{g^2}{2M_2} \xi_i \xi_j \rightarrow (m_\nu)_{ij} \approx \frac{M_Z^2}{M_2} \xi_i \xi_j$ controlled by $\xi_i = \frac{v_i \mu - \mu_i v_d}{\mu v_d}$
 (misalignment of μ_α and $(B\mu)_\alpha$): $\xi = |\vec{\xi}| \approx 2.5 \times 10^{-6} / \cos \beta \left(\frac{M_2}{\text{TeV}} \right)$
 - $m_1 = m_2 = 0$ (normal hierarchy), $\tan \theta_{23} = \xi_2 / \xi_3$, $\tan \theta_{13} = \xi_1 / (\xi_2^2 + \xi_3^2)^{1/2}$
- * $(\Delta m^2)_{12}$ and θ_{12} at 1-loop, controlled by μ_i / μ , $\tan \theta_{12} = \mu_1 / (\mu_2^2 + \mu_3^2)^{1/2}$
- * **LHC**: production and decay to LSP almost unaffected
- * Small R_P breaking effect ξ_i , μ_i visible through **LSP decay**

LSP decay

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- * In particular:

$$\frac{\text{BR}(\chi^0_1 \rightarrow \mu^\mp \bar{q} q')}{\text{BR}(\chi^0_1 \rightarrow \tau^\mp \bar{q} q')} \approx \frac{\text{BR}(\chi^0_1 \rightarrow W^\pm \mu^\mp)}{\text{BR}(\chi^0_1 \rightarrow W^\pm \tau^\mp)} \approx \left(\frac{\xi_2}{\xi_3}\right)^2 \approx \tan^2 \theta_{23}$$

[Mukhopadyaya Roy Vissani ph/9808265
Datta Mukhopadyaya Vissani ph/9910296
Porod et al ph/0011248]

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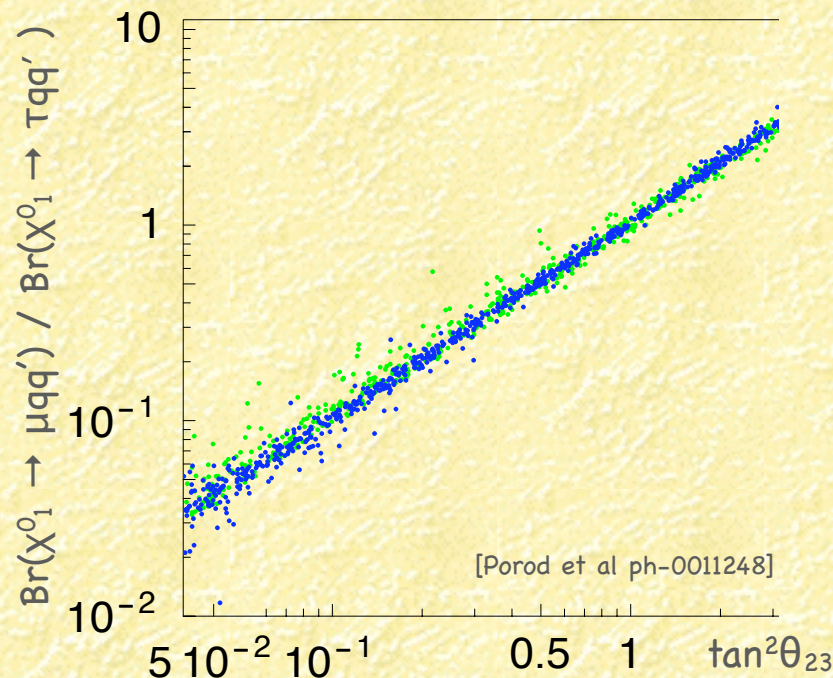
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[Mukhopadhyaya Roy Vissani ph/9808265
Datta Mukhopadhyaya Vissani ph/9910296
Porod et al ph/0011248]



$m_0 = (0.2-1) \text{ TeV}$

$M_2 \mu = (0-1) \text{ TeV}$

$A_0, B = (-3, 3) m_0$

$\tan \beta = (2.5-10)$

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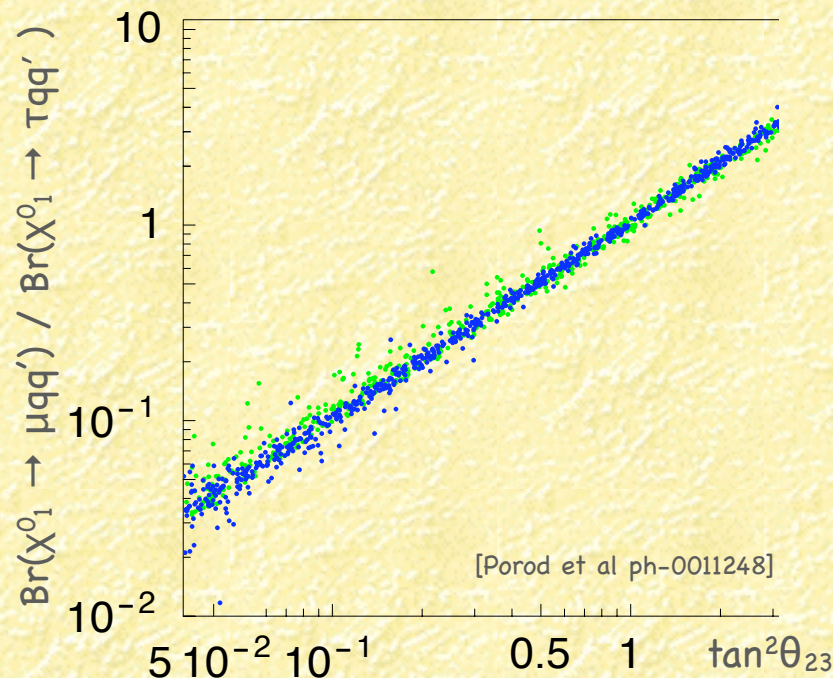
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[Mukhopadhyaya Roy Vissani ph/9808265
Datta Mukhopadhyaya Vissani ph/9910296
Porod et al ph/0011248]

LHC: BR ratio
at 3% with 100 fb⁻¹

[Porod Skands ph/0401077]



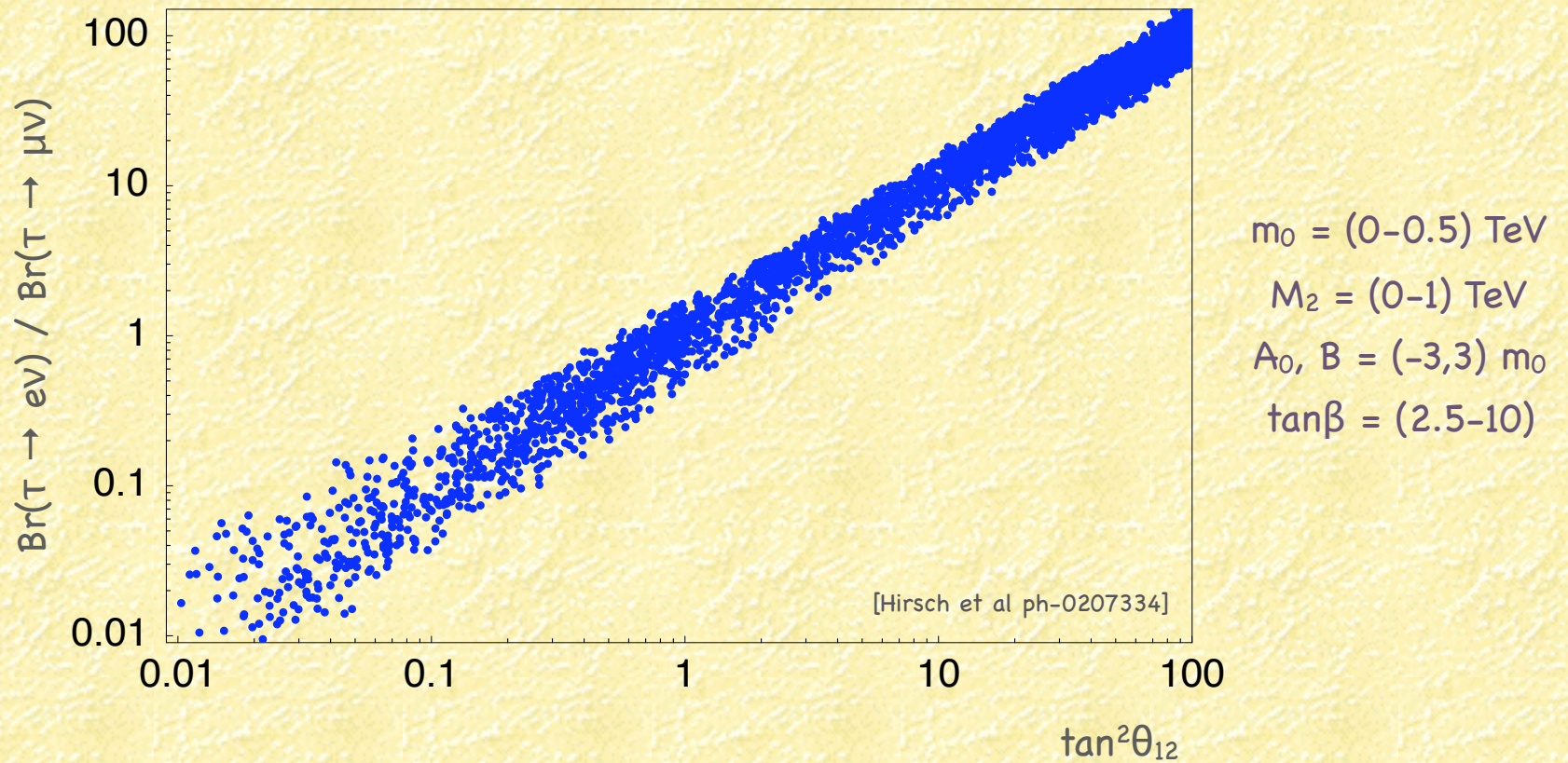
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$\tan \beta = (2.5-10)$

* Stau LSP:



* Other candidates → [Hirsch Porod ph/0307364]

Babu-Zee model

* Fields: **SM** + h^+ + k^{--} (SU(3)xSU(2) singlets, Y = 1, -2 respectively)

* $\mathcal{L} = \mathcal{L}_{\text{SM}} + f_{ij} l_i l_j h^+ + h_{ij} e_i^c e_j^c k^{--} + \mu h^+ h^+ k^{--}$ breaks L

* Neutrino masses generated at 2 loops

$$\text{BR}(h^+ \rightarrow e\nu) = \frac{\epsilon^2 + \epsilon'^2}{2(1 + \epsilon^2 + \epsilon'^2)}$$

$$\text{BR}(h^+ \rightarrow \mu\nu) = \frac{1 + \epsilon'^2}{2(1 + \epsilon^2 + \epsilon'^2)}$$

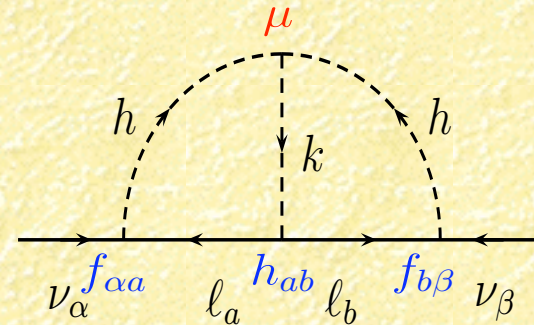
$$\text{BR}(h^+ \rightarrow \tau\nu) = \frac{\epsilon^2 + \epsilon'^2}{2(1 + \epsilon^2 + \epsilon'^2)}$$

where $\epsilon = \tan \theta_{12} \frac{\cos \theta_{23}}{\cos \theta_{13}}$ or $-\frac{\sin \theta_{23}}{\tan \theta_{13}}$

$\epsilon' = \tan \theta_{12} \frac{\sin \theta_{23}}{\cos \theta_{13}}$ or $\frac{\cos \theta_{23}}{\tan \theta_{13}}$

NH

IH



Babu-Zee model

* Fields: $SM + h^+ + k^{--}$ ($SU(3) \times SU(2)$ singlets, $Y = 1, -2$ respectively)

* $\mathcal{L} = \mathcal{L}_{SM} + f_{ij} l_i l_j h^+ + h_{ij} e_i^c e_j^c k^{--} + \mu h^+ h^+ k^{--}$ breaks L

* Neutrino masses generated at 2 loops

$$* \text{BR}(h^+ \rightarrow e\nu) = \frac{\epsilon^2 + \epsilon'^2}{2(1 + \epsilon^2 + \epsilon'^2)}$$

$$\text{BR}(h^+ \rightarrow \mu\nu) = \frac{1 + \epsilon'^2}{2(1 + \epsilon^2 + \epsilon'^2)}$$

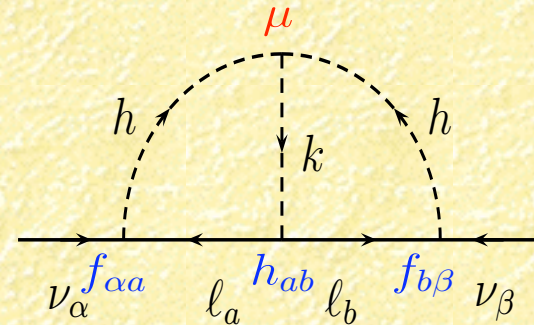
$$\text{BR}(h^+ \rightarrow \tau\nu) = \frac{\epsilon^2 + \epsilon'^2}{2(1 + \epsilon^2 + \epsilon'^2)}$$

where $\epsilon = \tan \theta_{12} \frac{\cos \theta_{23}}{\cos \theta_{13}}$ or $-\frac{\sin \theta_{23}}{\tan \theta_{13}}$

$$\epsilon' = \tan \theta_{12} \frac{\sin \theta_{23}}{\cos \theta_{13}}$$

NH

IH



[Zee 1985, Babu 1988]

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Summary

- * The standard understanding of neutrino masses in terms of a high scale breaking of an accidentally conserved lepton number is solid, economical, and general
- * Alternative, TeV-scale origins of neutrino masses are also allowed and offer the opportunity to probe such an origin at the LHC